



SEARCH AND DISCOVER

THE HUNT FOR RARE DECAYS OF ¹²⁴Xe & TRACE DETECTION OF KRYPTON IN XENON

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THE HUNT FOR RARE DECAYS OF ¹²⁴Xe

TRACE DETECTION OF KYPTON IN XENON

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DECAYS OF ¹²⁴Xe

- 124 124 124 54 Xe 124 72 2 Te
- Abundance of 0.095 % in natural xenon β -decay into ¹²⁴I energetically forbidden Double β -decay into ¹²⁴Te predicted Q-Value of 2856.73(12) keV allows several decay modes

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124 Xe + 2 e^{-1}	$T \rightarrow {}^{124}\text{Te}^* + 2\nu_e$	$(2\nu DEC)$
124 Xe + e^{-}	\rightarrow ¹²⁴ Te [*] + e^+ + $2\nu_e$	$(2\nu\beta^+\text{EC})$
¹²⁴ Xe	\rightarrow ¹²⁴ Te + 2 e^+ + 2 ν_e	$(2\nu 2\beta^+)$

Processes predicted by Standard Model

DECAYS OF ¹²⁴Xe





THE XENON DARK MATTER PROJECT

Dual-phase xenon detectors with extremly low background designed for dark matter search located at Laboratori Nazionali del Gran Sasso (LNGS)

- International collaboration (140 physicists)
- Successfully explores parameter space for potential dark matter particles (WIMPs)
- Different stages with the upcoming XENON1T to become the most sensitive dark matter detector in the world
- Suitable for other rare event searches

XENON100XENON10Image: Second second

Analysis of 225 live days will be shown in this talk

Will provide data this year. Sensitivity study shown in this talk

www.xenon1t.org

GXe

LXe

An energy deposition in xenon results in Scintillation & Ionization

Anti-correlated simultaneous production of electrons and photons

Number of photons and electrons is proportional to the deposited energy → Energy calibration

Different light and charge yields depending on interaction \rightarrow Particle discrimination

If electrons completly recombine they also produce only direct scintillation light

→ How to make use of both channels?

178 nm photons produced by dimers and therefore not absorbed by single xenon atoms → Transparency



The photons are detected with photomultipliers optimized for VUV-light (quantum efficiency 25% - 40%)

Photomultiplier tubes (PMTs)



Fast light signal from the emitted photons with a width of a few 100 ns \rightarrow S1

0.1 0.2 0.3 0.4 0.5 0.6 0.7 0.8

0.1 0.2 0.3 0.4 0.5 0.6 0.7 0.8

Time (us)

+2.089e2

+2.089e

70

60

50

40

30

20

10

Amplitude (pe/bin)

By applying an electric field (~kV/cm) the electrons generated by ionization are removed from the interaction site

They are drifted with a constant velocity (~mm/µs) towards the gas phase

There they are pushed and accelerated by a secondary strong field (10 kV/cm) and produce a spreadout light signal by proportional scintillation ("electroluminescence")

 \rightarrow S2

- kV/cm





x,y – position reconstruction via pattern (O(mm) resolution) S2 top 10 16 -10

> 15 -15 - 10-5 0 5 10

resolution of O(0.1 mm)

3D-reconstruction possible

- Discrimination of double scatters
- Selection of a fiducial volume with reduced external background

 $2\nu DEC$

<u>WHY SEARCH FOR DOUBLE</u> <u>ELECTRON CAPTURE WITH THIS</u> <u>DETECTOR?</u>

Expected signal for two K-shell electron capture

64.33 keV

- Source = Detector
- High self-shielding capacity
- 3D vertex reconstruction allows for selection of a fiducial volume
- Careful screening of materials and active removal of radioactive krypton

Due to the small range (<0.5 mm) and time difference (~10⁻¹⁵ s) the individual X-rays (Auger electrons) at cannot be resolved

Extremely low background experiment with keV-scale optimized sensitivity

Data is for "free" as it is the same as for the dark matter search

diffe



Select 34 kg fiducial volume, corresponding to 29 g of ¹²⁴Xe (0.095%)

Apply data quality and selection cuts and estimate their acceptance using ²³²Th & ⁶⁰Co calibration sources



 Γ : decay rate η : ¹²⁴Xe abundance mt: exposure N_A : Avogadro's constant M_{XE} : molar mass of xenon σ_{sig} : peak width μ_{sig} : peak position f_{bkg} : linear background Bayesian fit from 10 keV to 135 keV with two models:

Linear background only Linear background and a Gaussian signal







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Linear background only Linear background and a Gaussian signal





Use knowledge about parameters Implement systematical uncertainties as Gaussian priors Evaluate signal significance with Bayes factor (BF)



A. Caldwell, D. Kollar, and K. Kroninger, Comput. 438 Phys. Commun. 180, 2197 (2009).

Expected signal: single energy peak at Bayesian fit from 10 keV to 135 keV with two models:

Background + Signal model (p = 0.92)
Background only (p = 0.89)

Result on the 90% lower credibility limit of the double K-shell electron capture from XENON100 data

$$T_{\frac{1}{2}} > 6.5 \ x \ 10^{20} \ yr$$

arXiv:1609.03354 Submitted to Phys. Rev. C

Other results on the half-life

XMASS (Abe et al.): > $4.7 \cdot 10^{21} yr$ Gavrilyuk et al. : > $2.0 \cdot 10^{21} yr$ This result supersedes the obtained limit by Mei et al. (> $1.6 \cdot 10^{21} yr$)





Favors background only model Calculate lower limit on the half-life

A. Caldwell, D. Kollar, and K. Kroninger, Comput. 438 Phys. Commun. 180, 2197 (2009).

XENON1T – THE NEXT STEP

30 x less background



Assuming the same light yield as XENON100 (pessimistic) and a fiducial volume of 1 ton (1kg of ¹²⁴Xe):

5 live days to achieve the highest sensitivity 2 live years to be sensitive up to $6.1 \cdot 10^{22}$ yr half-life @ 90 % C.L.

Additionally: Due to improved detector technology the search for high energy signals (e.g. $2vEC\beta^+$, 0vDEC) becomes possible

- Located in Hall B at LNGS underground laboratory
- Immersed in a water tank providing an active muon veto
- Built to improve the existing dark matter search sensitivity by 2 orders of magnitude
- sensitivity by 2 orders of magnitude
 > 3.2 t of xenon in total with 2 t of LXe in the active volume



FIRST DATA EXPECTED THIS YEAR





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Background ⁸⁵Kr

- Beta decay of ⁸⁵Kr into ⁸⁵Rb is a significant intrinsic background for XENON1T
- Determination of the ⁸⁵Kr concentration in xenon is of crucial importance for knowledge of the signal background
- Ratio of ⁸⁵Kr to natural krypton is at 10⁻¹¹, while natural krypton is aimed to be below <0.2 ppt in the used xenon

Background ⁸⁵Kr

• Beta decay of ⁸⁵Kr into ⁸⁵Rb is a significant

Measuring the Krypton concentration at the sub-ppb level is not trivial!

while natural krypton is aimed to be below <0.2 ppt in the used xenon



Reduce pressure to meet hardware specs of the mass spectrometer



Reduce pressure to meet hardware specs of the mass spectrometer

> Differential Pumping Sections

Butterfly

You want to now the kryptor content

This way we reached a sensitivity of ~ 60 (200) ppt (masterthesis) Still not 0.2 ppt!

You measure the krypton amount in the remaining xenon

Turbo

You reduce the **xenon** by freezing down to its vapor pressure

Cold Trap ৰ

Krypton content stays constant since vapor pressure is not reached This keeps the vacuum ufficient to not blow u the mass spectrometer





Analysis ongoing Promising preliminary results Project finished soon

