

# Measuring the neutrino mass with KATRIN

Annual Retreat of RTG 2149 “*Strong and Weak Interactions – From Hadrons to Dark Matter*”  
WWU Münster / Telgte, 24-26 November 2015

Kathrin Valerius (KIT Center Elementary Particle and Astroparticle Physics, KCETA)



 **Graduiertenkolleg 2149**  
**Research Training Group**





# Outline

brief  
motivation

direct  $\nu$ -mass  
measurement

status  
& outlook

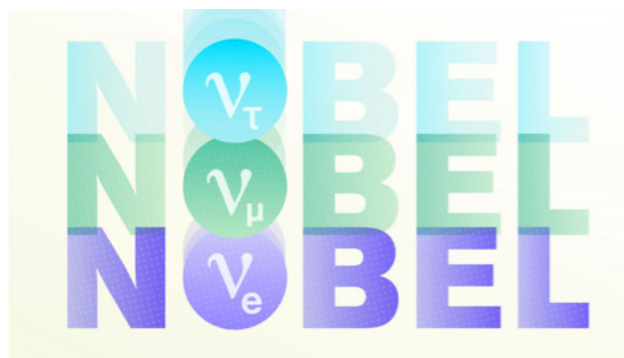
- I. Why study massive neutrinos?
- II. How does neutrino mass measurement with KATRIN work?
- III. What are current steps to prepare the start of measurements with KATRIN?

puzzling  
neutrinos

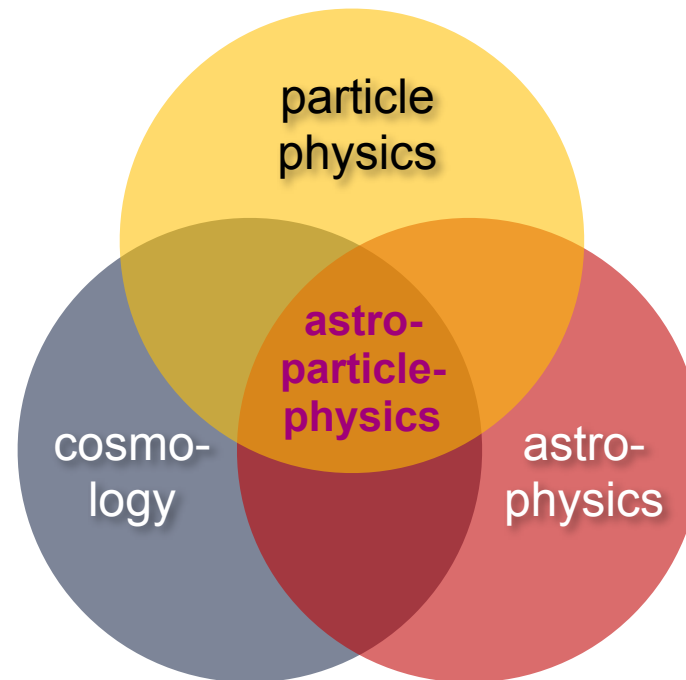
elusive  
neutrinos

challenging  
neutrinos

# I. Motivation: Massive neutrinos

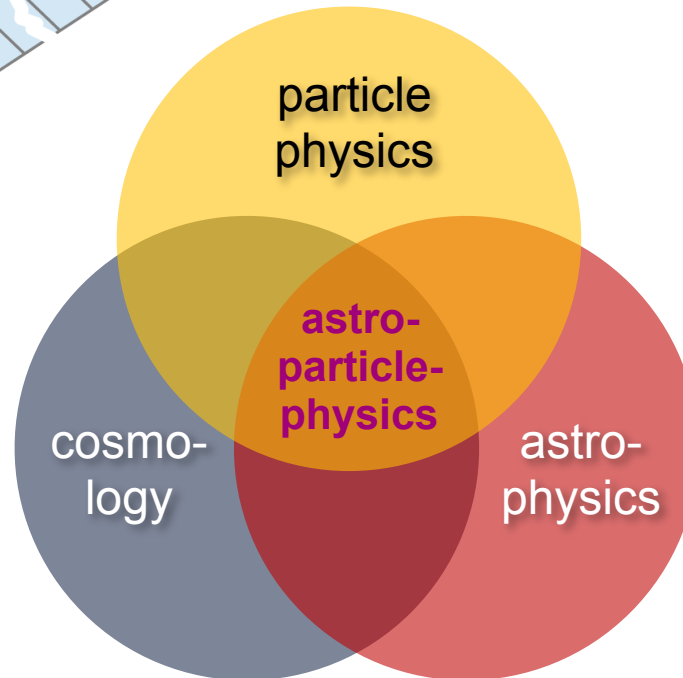
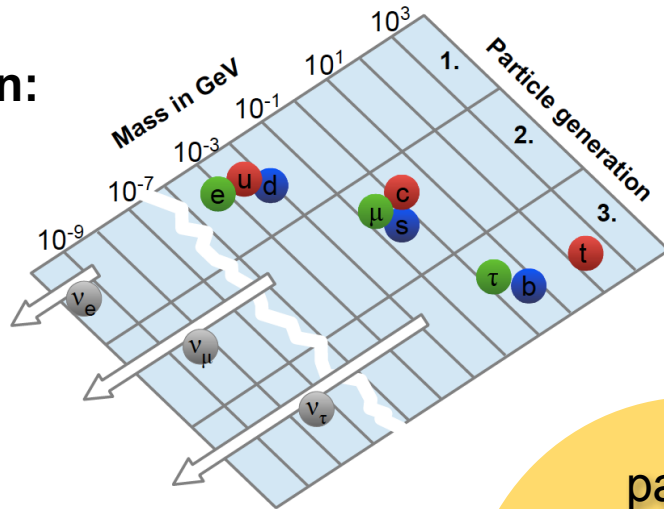


# Massive neutrinos in astroparticle physics



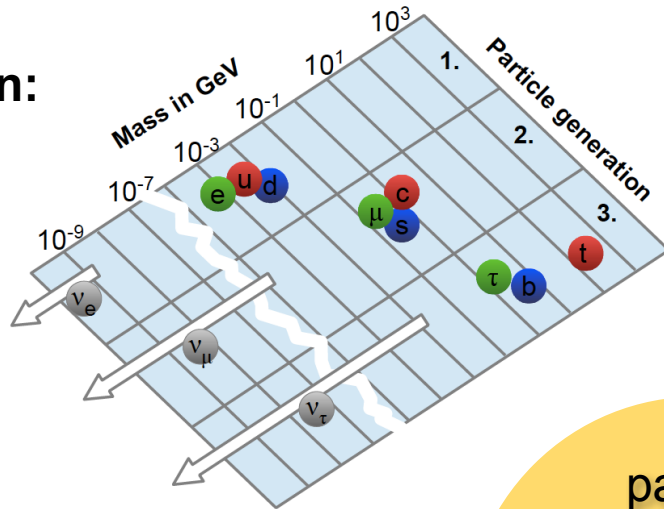
# Massive neutrinos in astroparticle physics

mass generation:  
new concepts

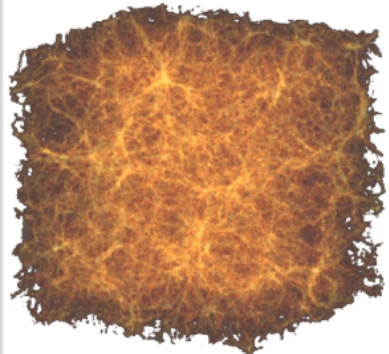


# Massive neutrinos in astroparticle physics

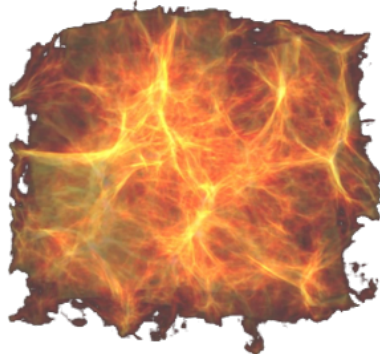
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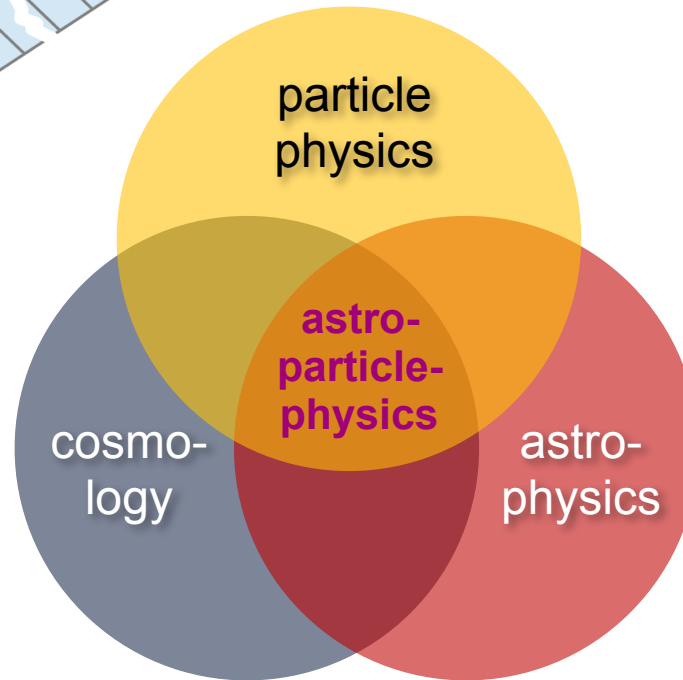
massive neutrinos as  
“cosmic architects”



$\Sigma m_\nu = 0 \text{ eV}$



6.9 eV

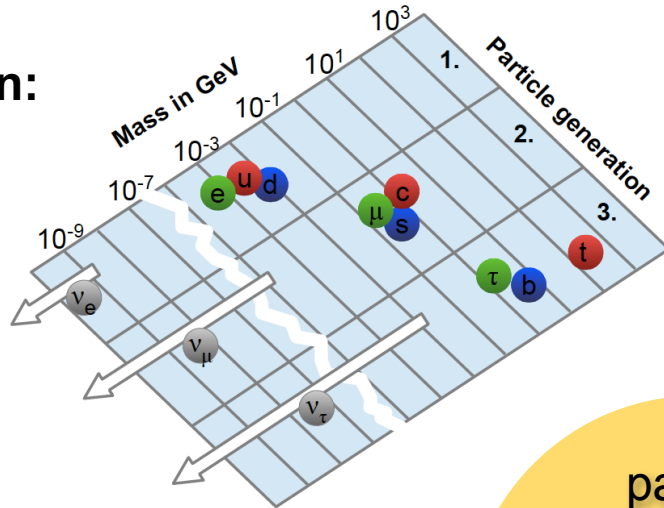


© T. Haugbølle



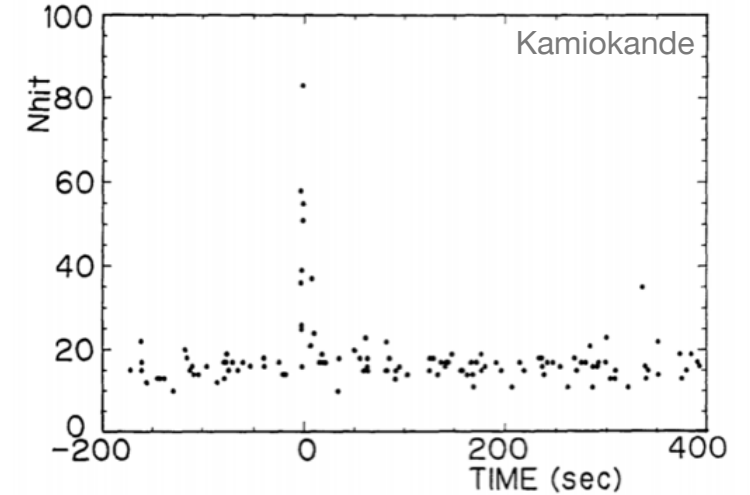
# Massive neutrinos in astroparticle physics

mass generation:  
new concepts

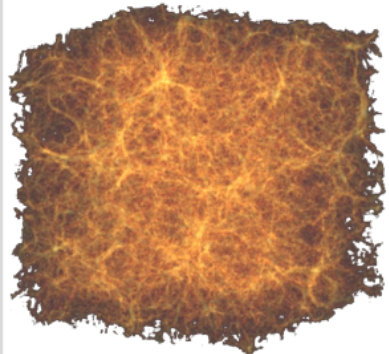


understanding  
astrophysical processes

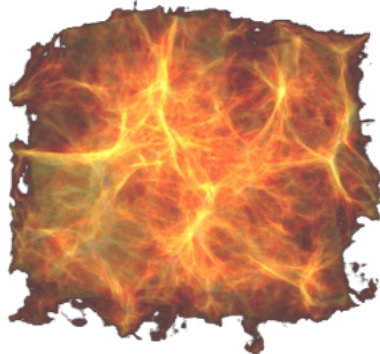
Neutrino burst from SN 1987a



massive neutrinos as  
“cosmic architects”

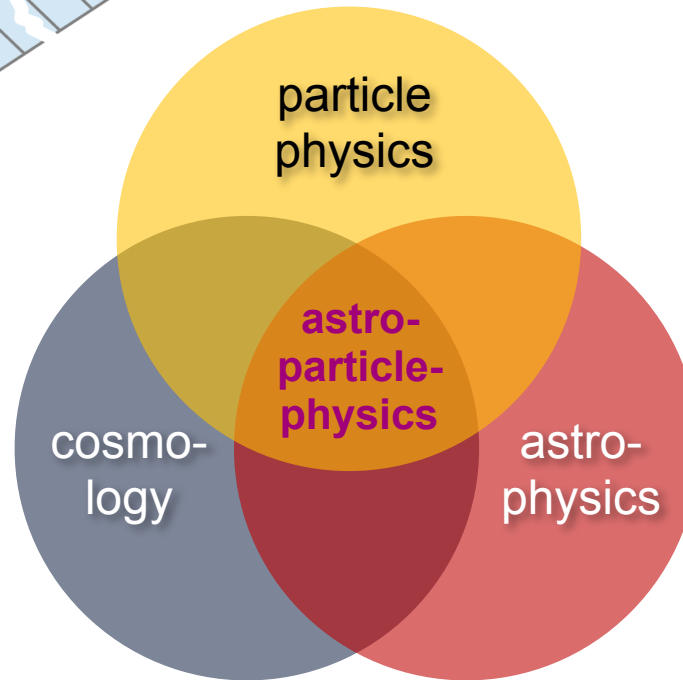


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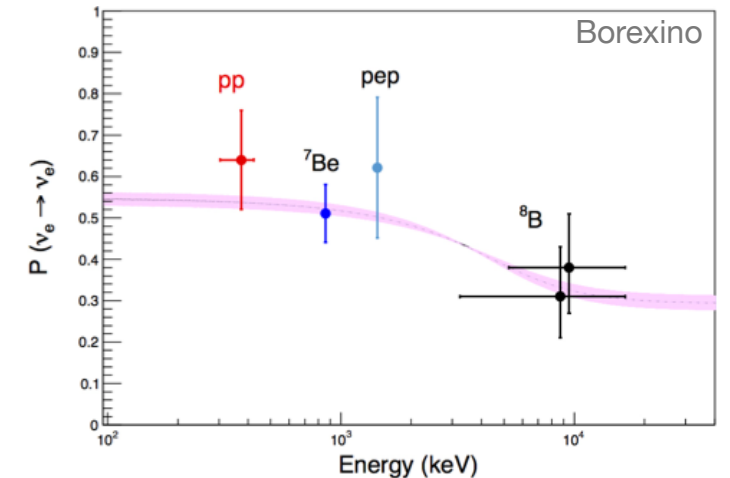


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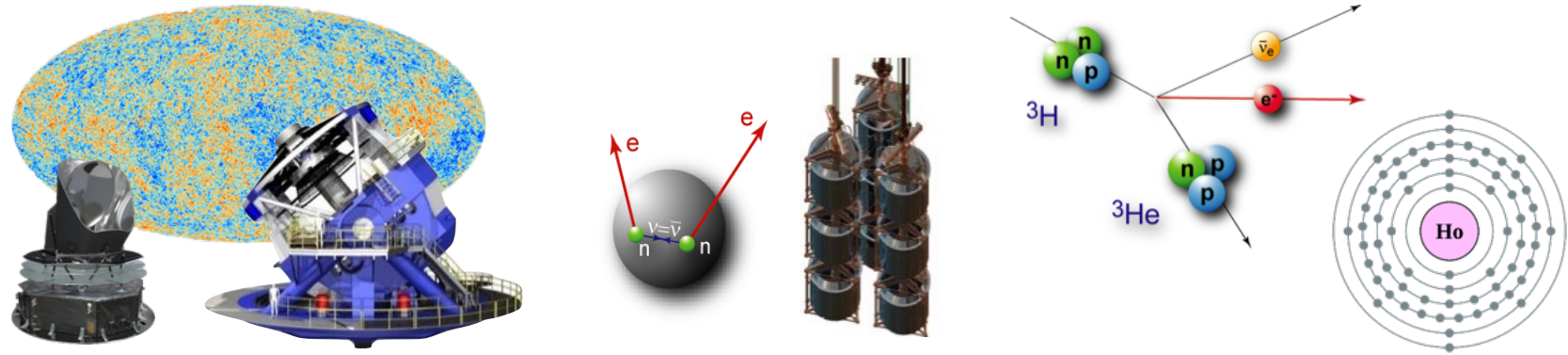
© T. Haugbølle



Matter effects in the sun

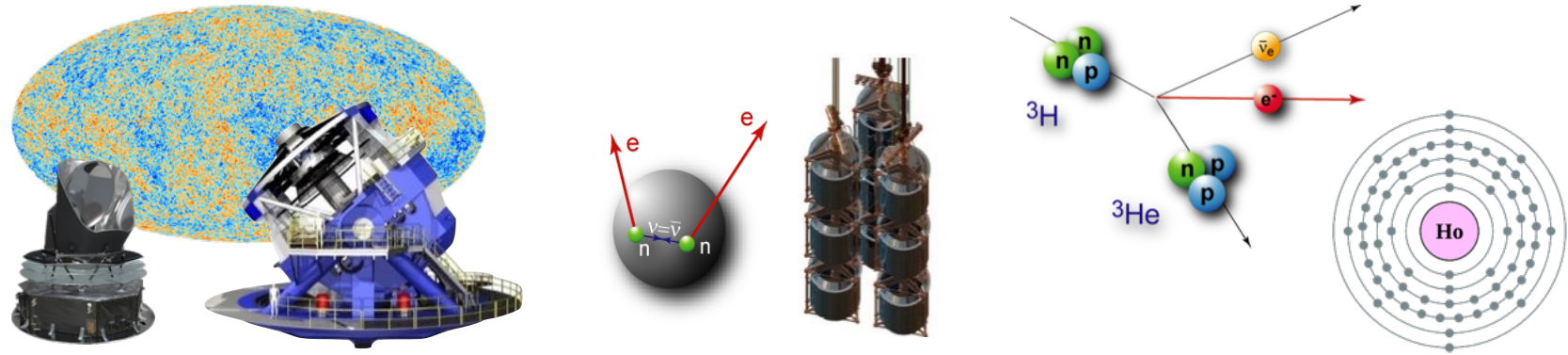


# Complementary paths towards $\nu$ masses



Tool	Cosmology CMB + LSS + ...	Neutrinoless double $\beta$ -decay	$\beta$ -decay endpoint and EC
Observable	$\sum m_\nu = \sum_{i=1}^3 m_i$	$\langle m_{\beta\beta} \rangle = \left  \sum_{j=1}^3  U_{ej} ^2 m_j e^{i\alpha_j} \right $	$m_\beta^2 = \sum_{i=1}^3  U_{ei} ^2 m_i^2$
Present upper limit	0.2 – 1 eV	0.2 – 0.4 eV	2 eV
Potential	20 – 50 meV	20 – 50 meV	200 meV
Model dependence	Multi-parameter cosmological model	<ul style="list-style-type: none"> <li>- Majorana vs. Dirac</li> <li>- nucl. matrix elements</li> </ul>	<ul style="list-style-type: none"> <li>- <b>Direct</b>, only kinematics;</li> <li>- agnostic to Dirac/ Majorana nature</li> </ul>

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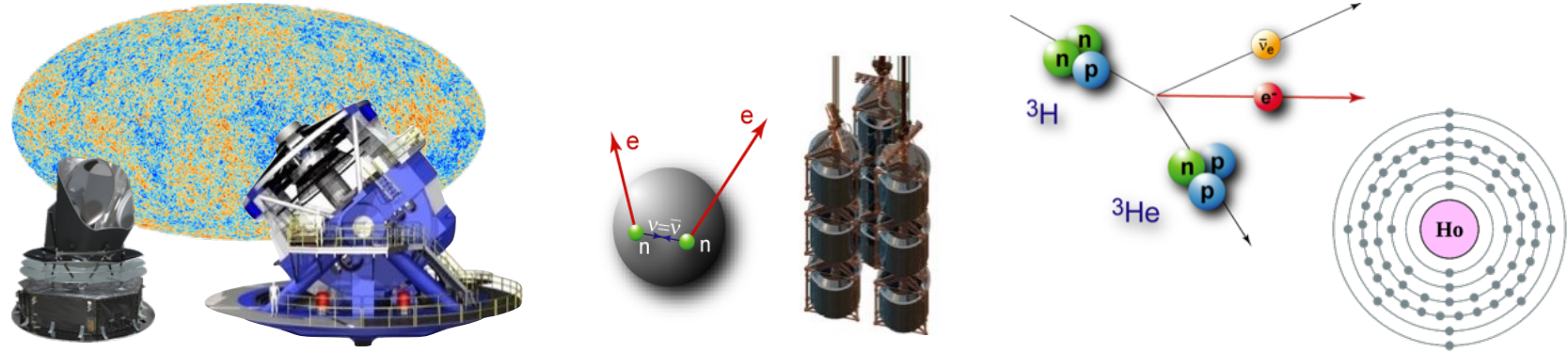


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S. Hannestad

M. Lindner

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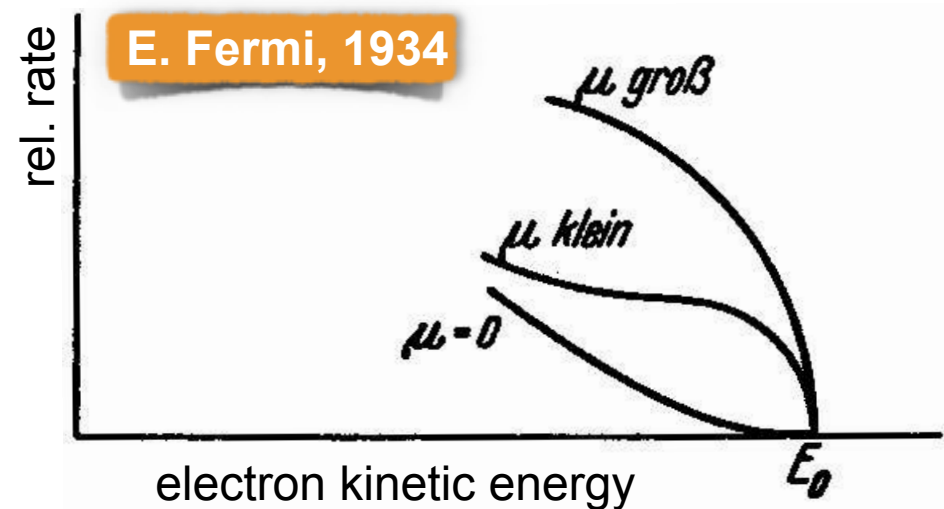
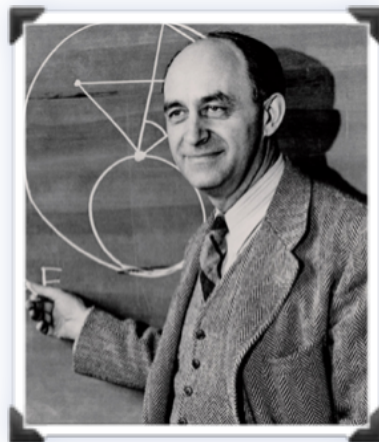


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S. Hannestad

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## II. Method: Direct neutrino mass measurement in the laboratory



# Direct neutrino mass measurement

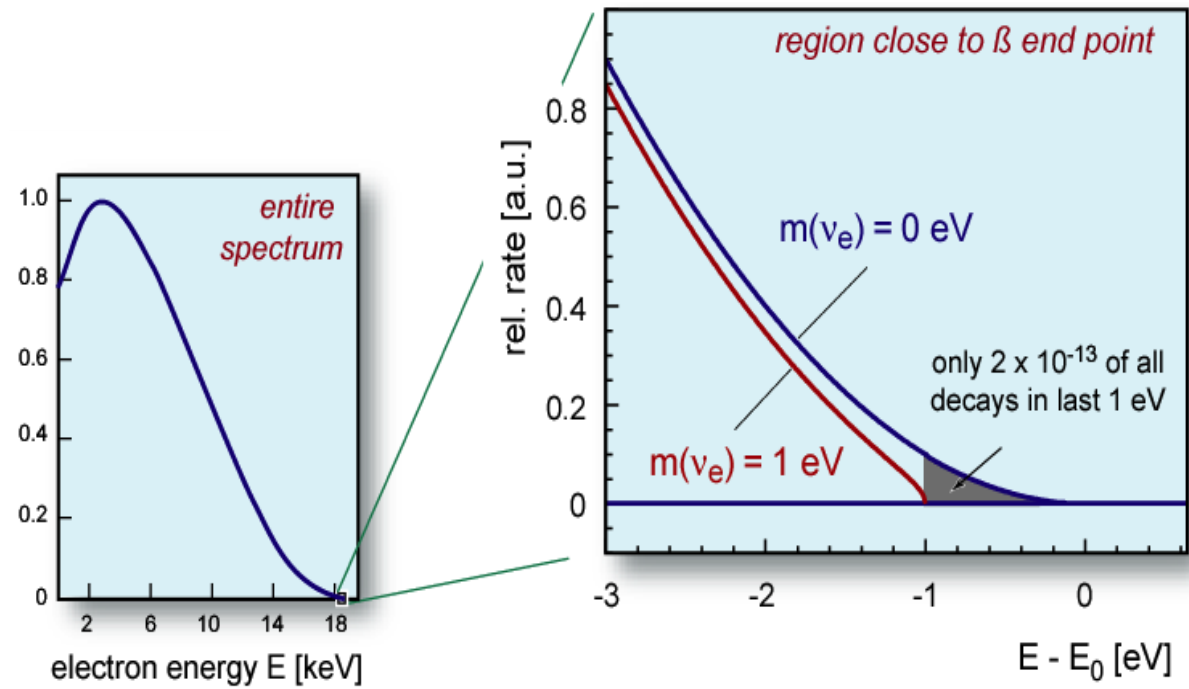
Imprint of  $m_\nu$  on **endpoint region** of  $\beta$  spectrum (similar for EC):

$$\frac{dN}{dE} = C \cdot F(Z, E) \cdot p \cdot (E + m_e) \cdot (E_0 - E) \cdot \sqrt{(E_0 - E)^2 - m^2(\nu_e)}$$

$$m^2(\nu_e) = \sum |U_{ei}|^2 m_i^2$$

observable: effective squared mass

- Key requirements**
- Source isotope:
    - Low spectral endpoint Q
    - Large decay rate (short  $T_{1/2}$ )
  - Instrument:
    - Excellent energy resolution
    - Very low background

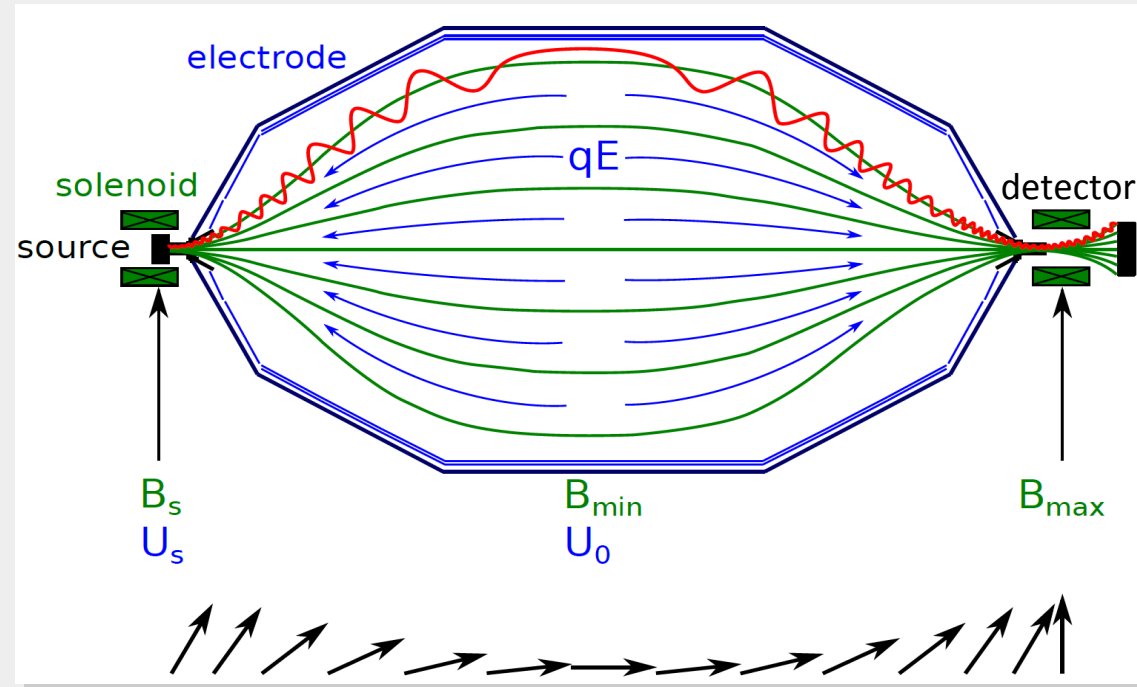


# Spectroscopic technique for $\beta$ decay

## MAC-E filter technique

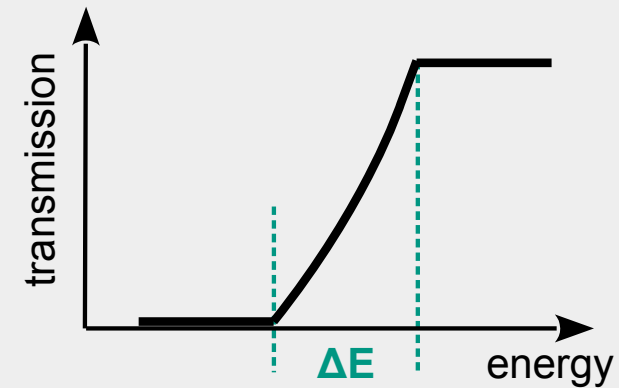
Magnetic Adiabatic Collimation with Electrostatic filter

Picard et al., NIM B63 (1992) 345



$$\mu = \frac{E_{\perp}}{B} = \text{const.}$$

Sharp high-pass filter:



Steps of filter potential  
→ integrated  $\beta$  spectrum

Combination of high luminosity  
and high energy resolution:

$$\frac{\Delta E}{E} = \frac{B_{\min}}{B_{\max}} = \frac{1}{20000}$$

(at KATRIN)



# KATRIN: A next generation tritium beta decay experiment with sub-eV sensitivity for the electron neutrino mass







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PDG 2014

<u>VALUE (eV)</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	
<b>&lt; 2 OUR EVALUATION</b>			
< 2.05	95	<sup>1</sup> ASEEV	11
< 2.3	95	<sup>2</sup> KRAUS	05

Troitsk exp.

Mainz exp.



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Troitsk exp.  
Mainz exp.

## Sensitivity on $m(\nu_e)$ :

2 eV → **0.2 eV** (90% CL, 3 net years)

→ Requires **x100** improvement on  $m^2(\nu_e)$

→ Use expertise and infrastructure at KIT (Tritium Laboratory Karlsruhe, TLK)





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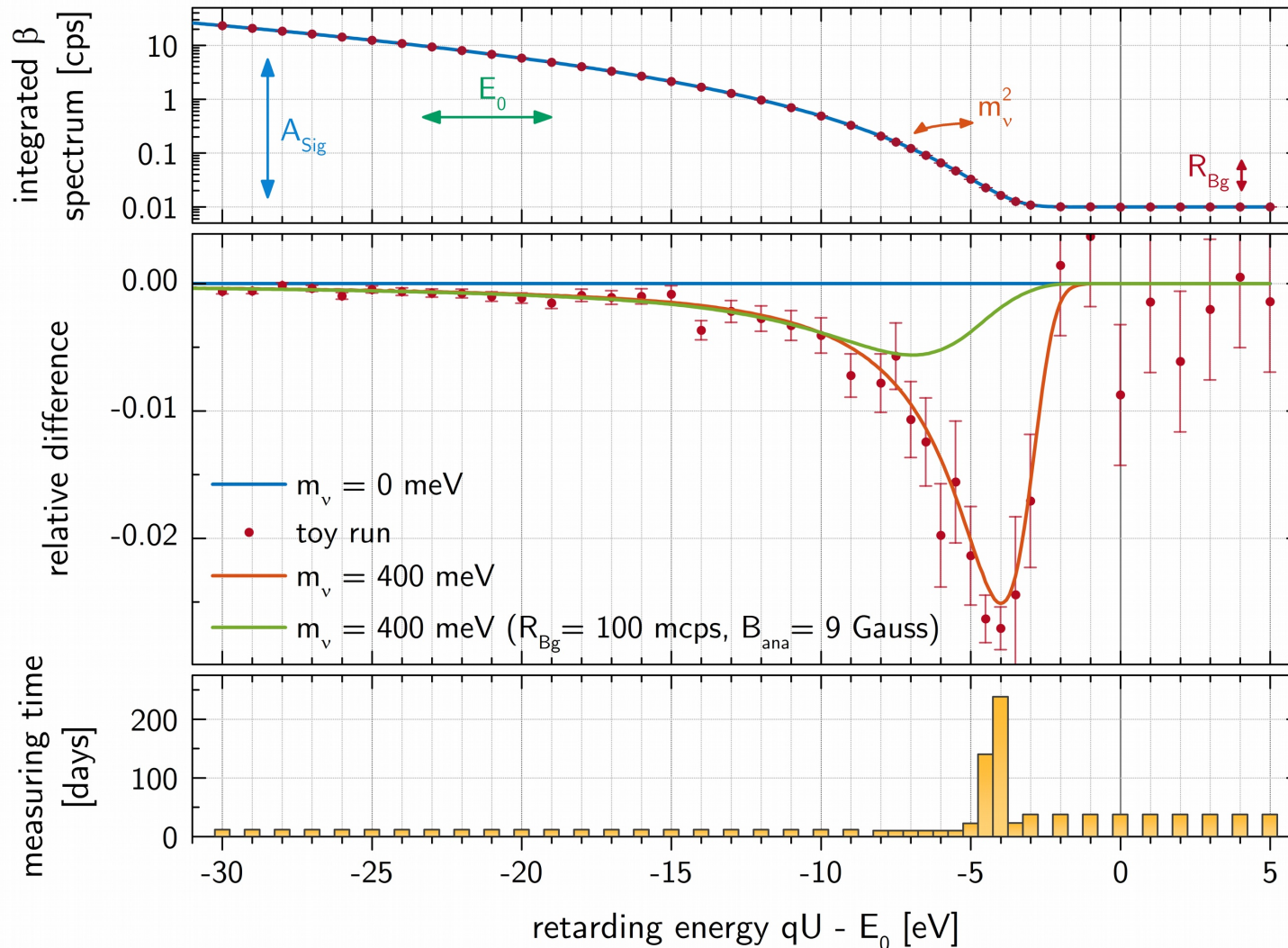


**KATRIN collaboration:**  
~120 members  
from 15 institutions in 5 countries

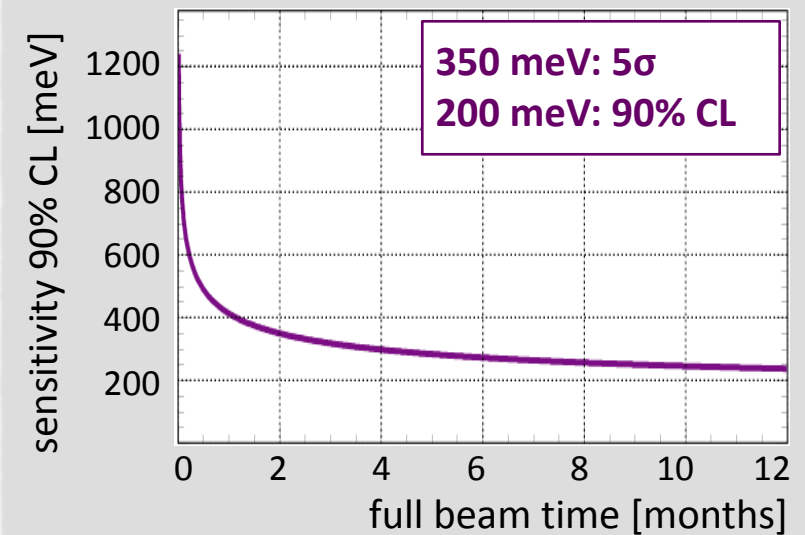


# KATRIN: spectral fit & $\nu$ -mass sensitivity

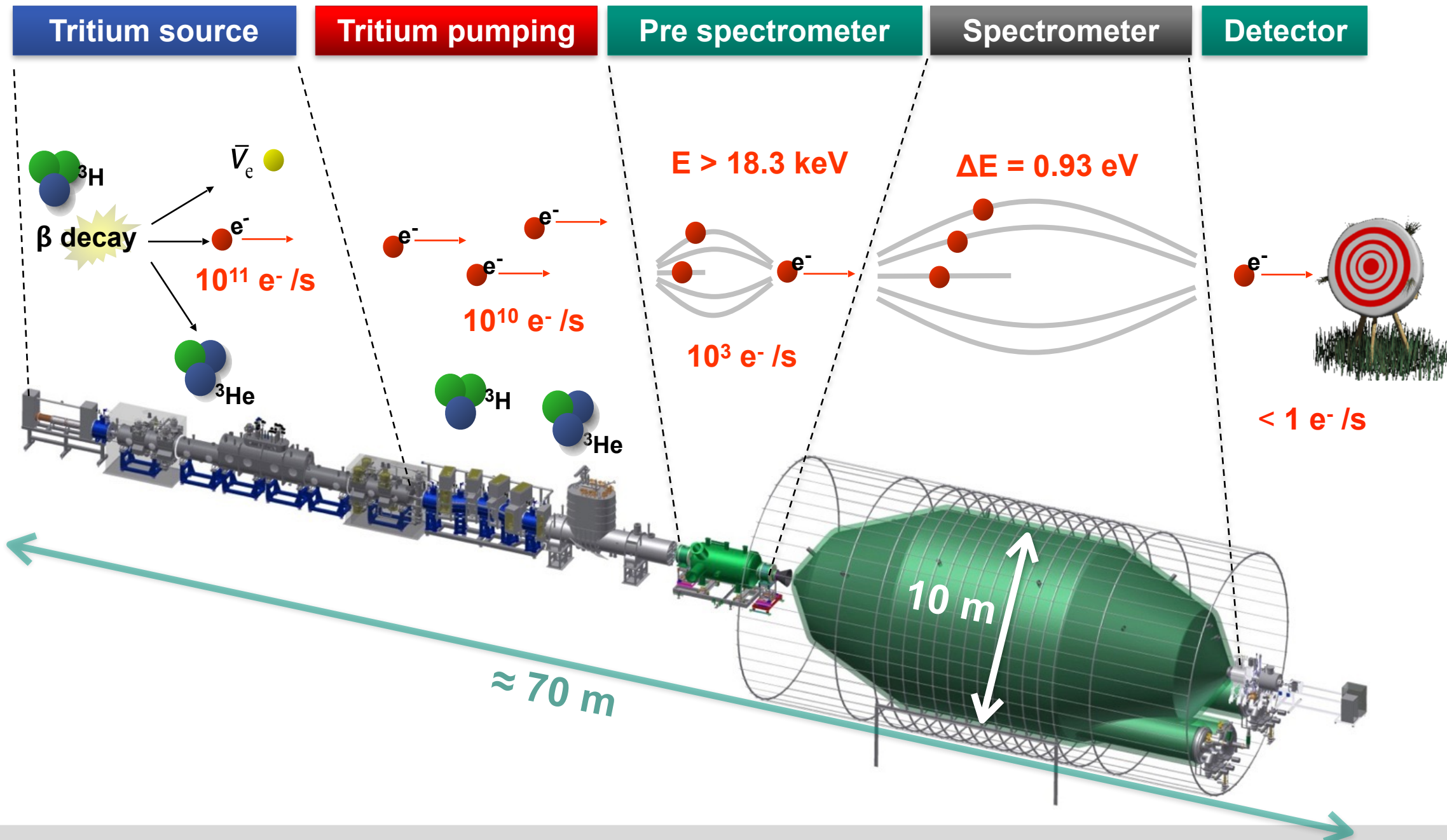
Relative shape measurement of integrated  $\beta$  spectrum:



- 4 fit parameters:  
 $m_\nu^2$ ,  $E_0$ ,  $A_S$ ,  $R_{Bg}$
- After 3 yrs of data (~5 cal. yrs):  
balance of statistics and systematics



# KATRIN overview: 70 m beamline

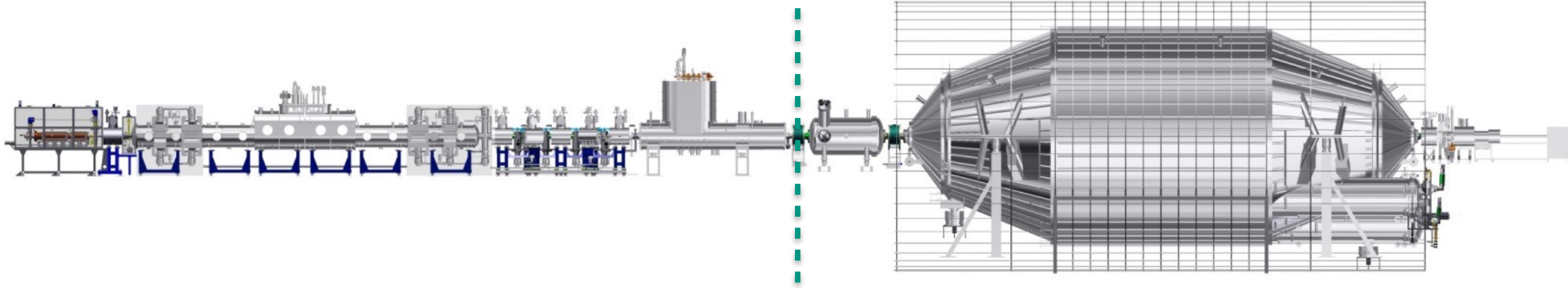


Factor of **10** in sensitivity  
seems easy on paper, but ...

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seems easy on paper, but ...

Eggenstein near Karlsruhe,  
Nov. 25, 2006





## Source and Transport Section

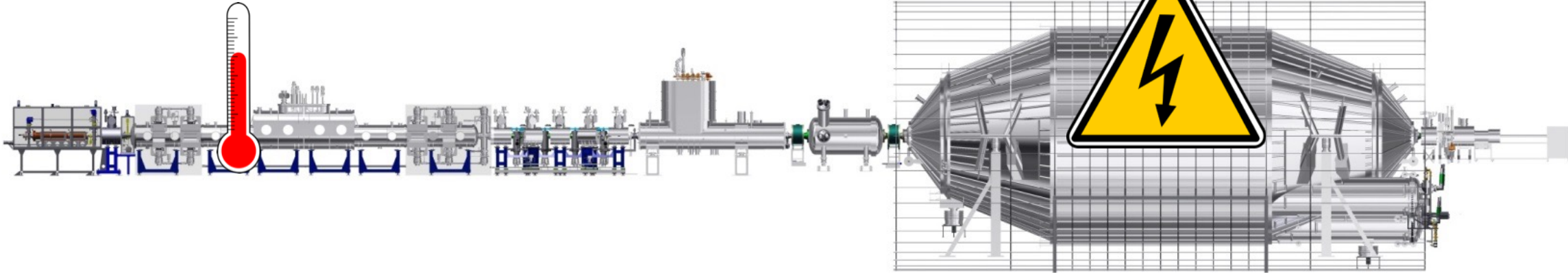
- Windowless gaseous tritium source
  - Intensity ( $10^{11}$  decays/s)
  - Stability ( $10^{-3}$  h $^{-1}$ )
  - Isotopic purity ( $> 95\%$ )
- Tritium retention (factor  $> 10^{14}$ )
- Adiabatic transport of electrons

## Spectrometer and Detector Section

- Spectrometer UHV ( $p < 10^{-11}$  mbar)
- Energy resolution ( $< 1$  eV at 18.6 keV)
- High voltage stability (sub-ppm/month)
- High detection efficiency ( $10^{-3}$ - $10^3$  cps)
- Low background rate ( $10^{-2}$  cps)

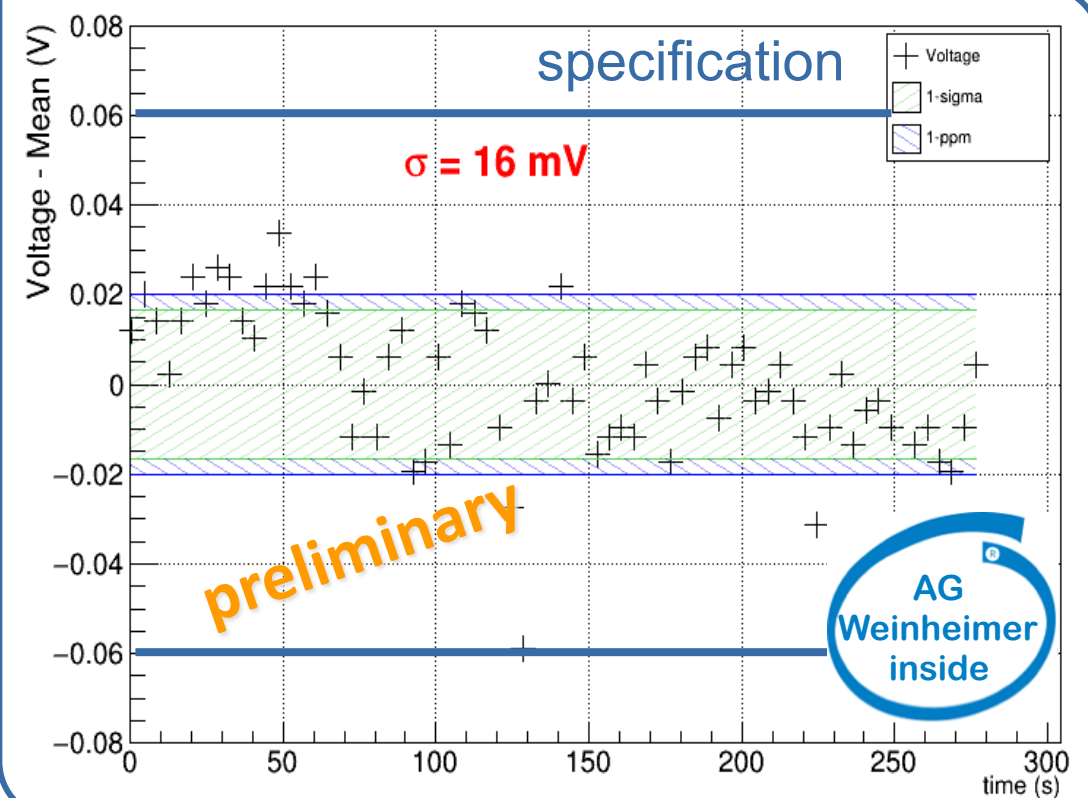
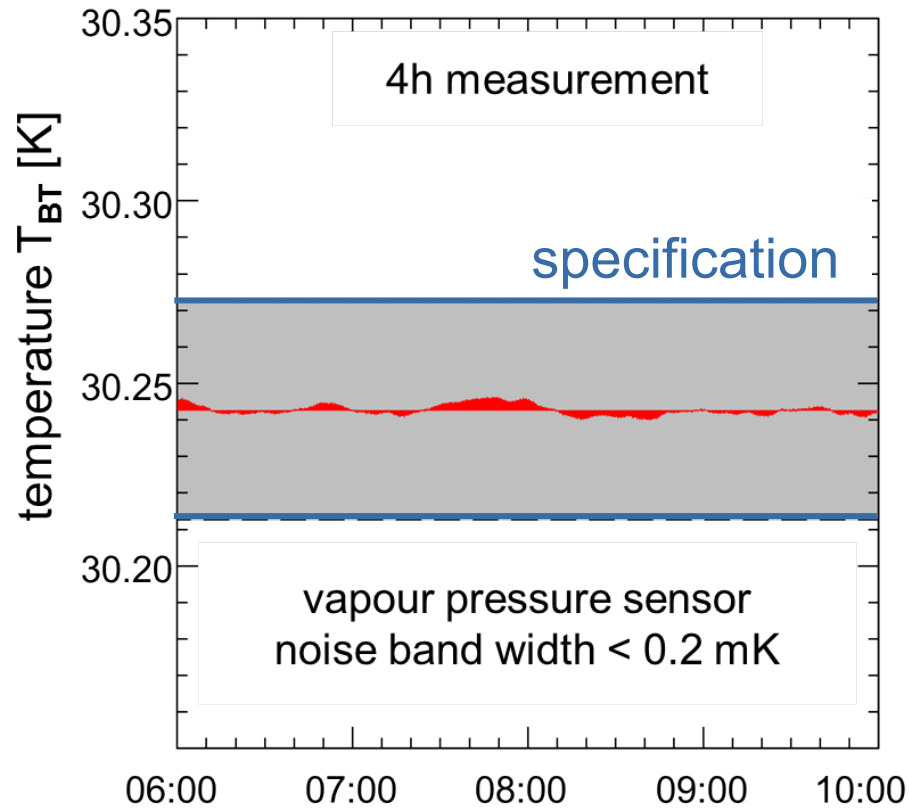


# Experimental challenges ... and solutions



Source demonstrator:  $\Delta T/T \sim 10^{-4}$

HV post-regulation:  $\Delta U \sim 1$  ppm



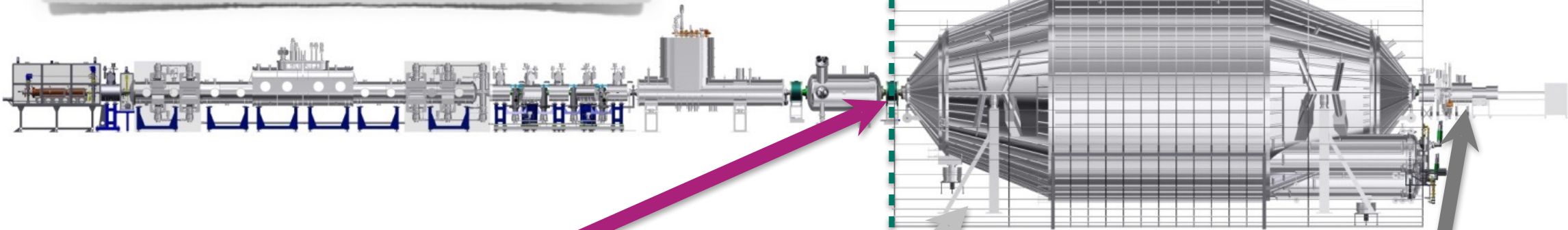
### III. Status of KATRIN

## & route towards start of measurements

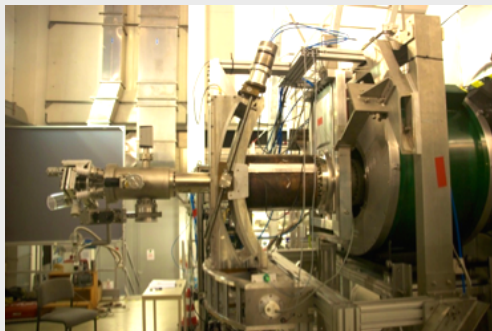


# System integration and commissioning

Spectrometer & detector:  
successful commissioning  
2013-2015



**electron gun**  
well-defined  
energy and angle



**magnetic fields**

- s.c. magnets
- field-shaping  
air coil systems

**precision high voltage**

vessel + wire  
electrodes

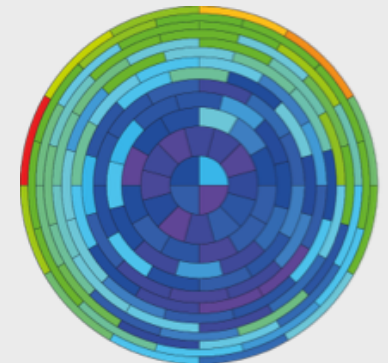


**vacuum system**

- 1250 m<sup>3</sup> vacuum recipient
- TMPs + 3 km getter strips  
(+ LN<sub>2</sub>-cooled baffles)



**148-pix detector**  
spatial & timing info



# Spectrometer & detector commissioning

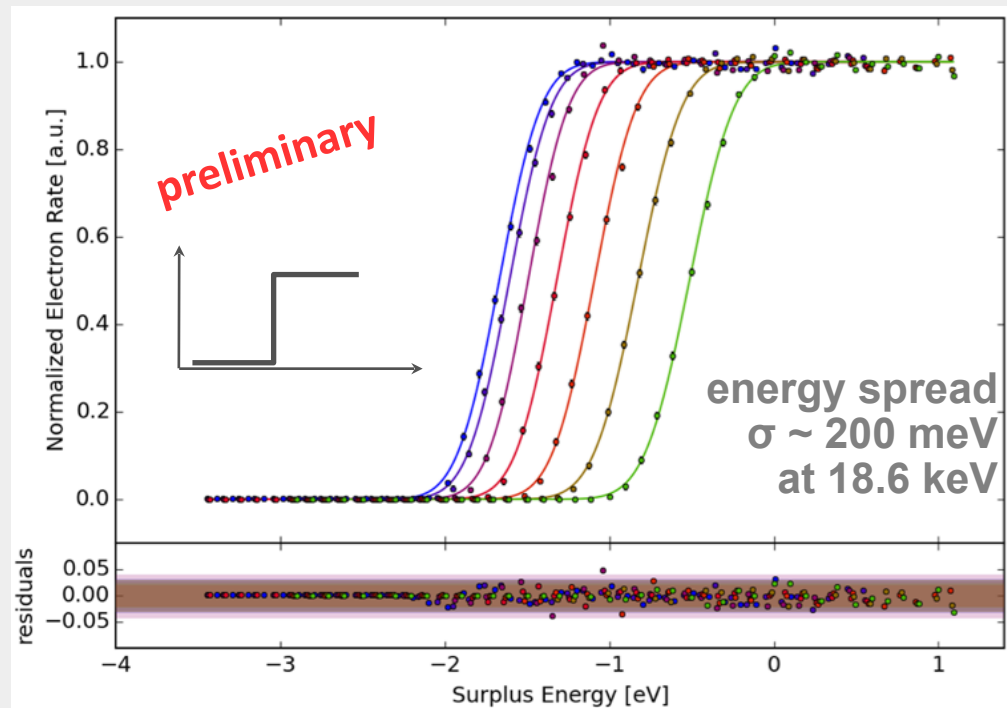
## Characterisation of spectrometer transmission

using precision electron source:  
energy- & angle-selective, point-like

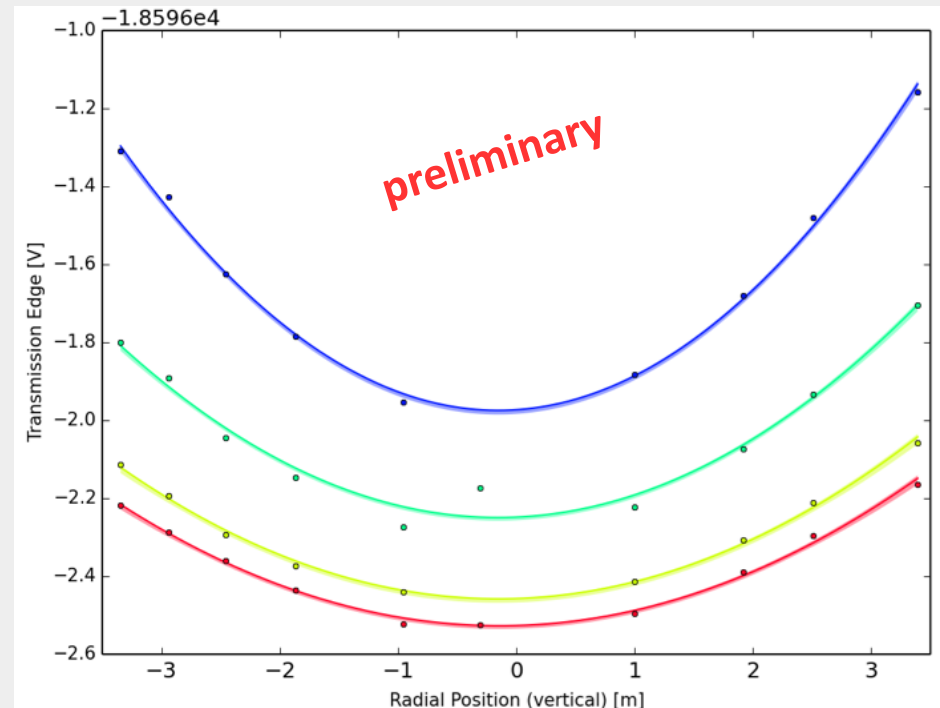
AG  
Weinheimer  
inside



Transmission characteristics of main spec.  
as expected (limited by e-gun systematics ...)



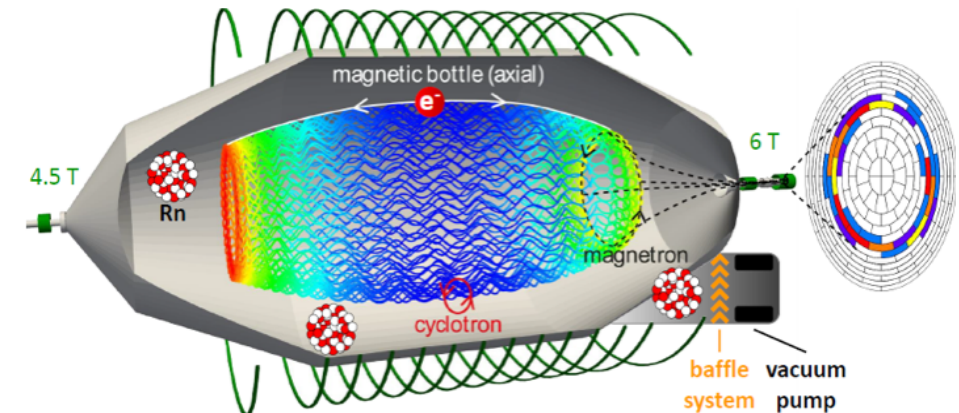
Radial dependence of retardation potential  
as expected (precision mapping by e-gun)



# Spectrometer & detector commissioning

## Characterisation of backgrounds

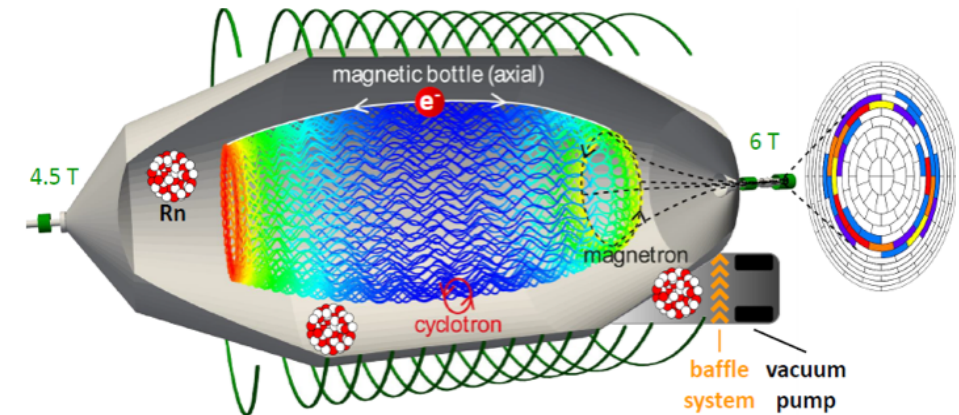
- Very efficient magnetic & electrostatic shielding, but only for charged particles ( $e^-$  and  $H^-$ )
- Neutral, unstable atoms ( $^{219}, ^{220}\text{Rn}$ ,  $H^*$ ) can penetrate into inner flux tube  
→ further measures required, e.g. passive shielding against Rn-induced secondaries



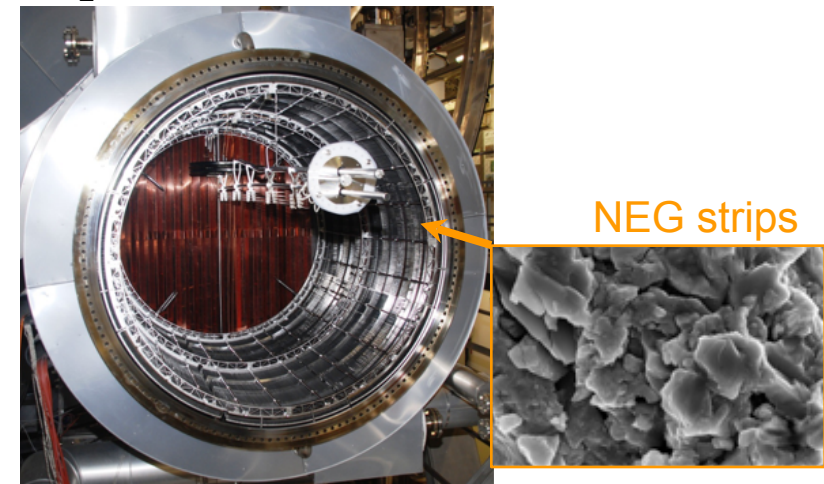
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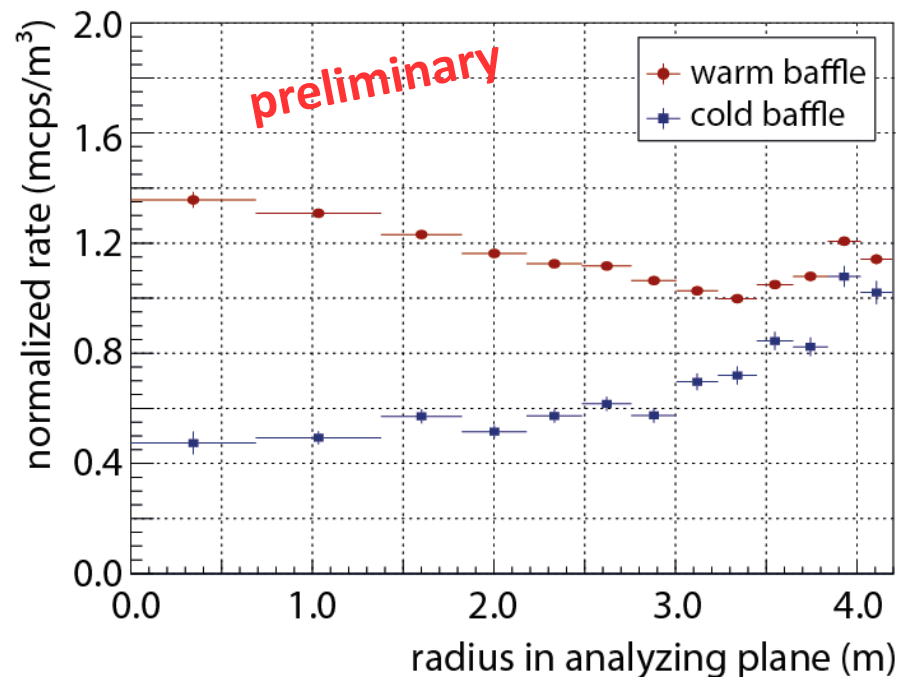
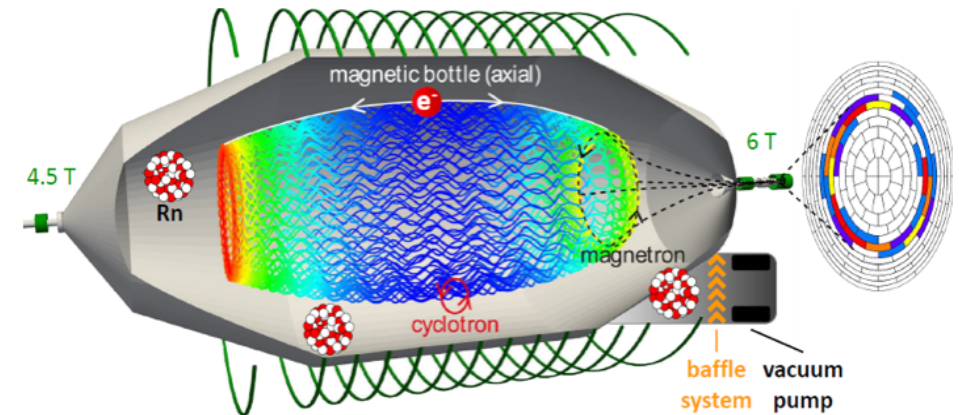
LN<sub>2</sub>-cooled baffles



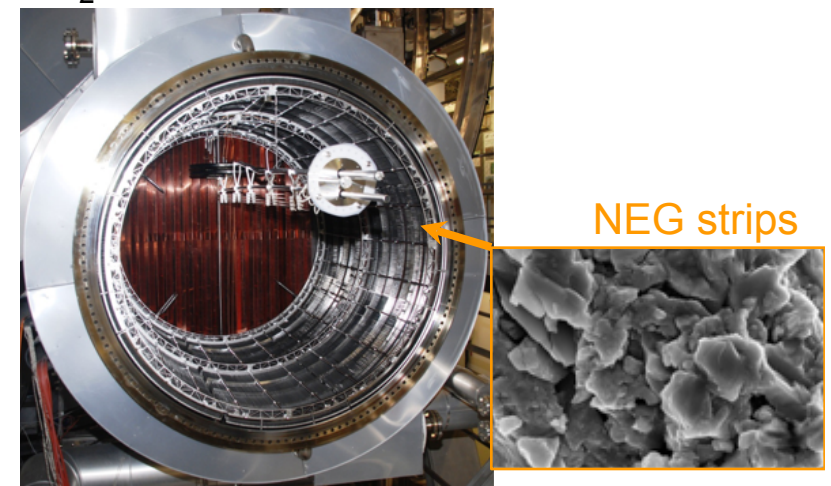
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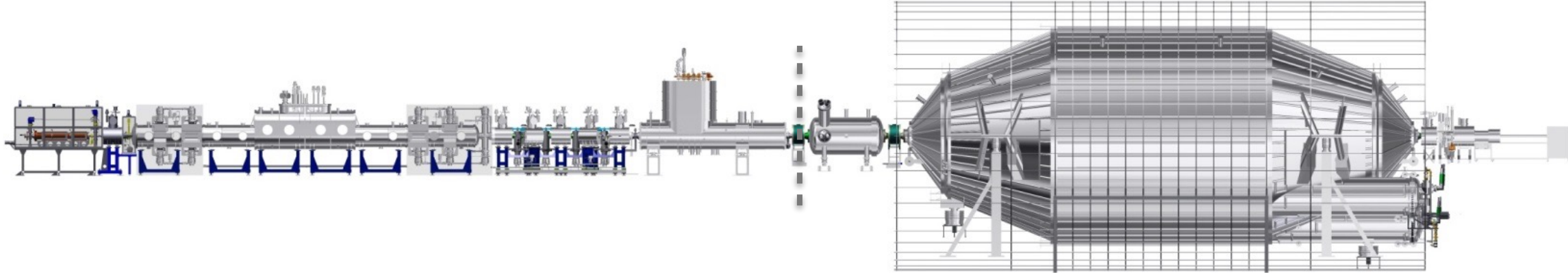
LN<sub>2</sub>-cooled baffles



preliminary

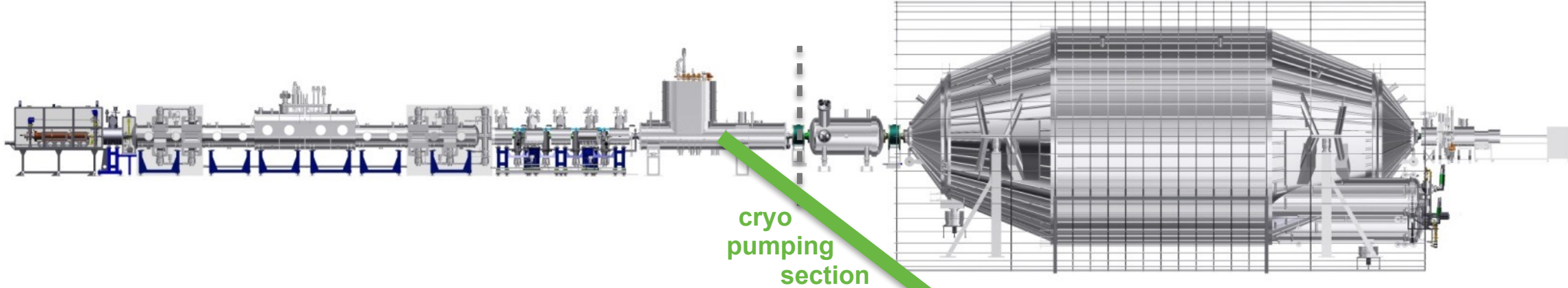
477 ± 3 mcps background level achieved

# System integration and commissioning





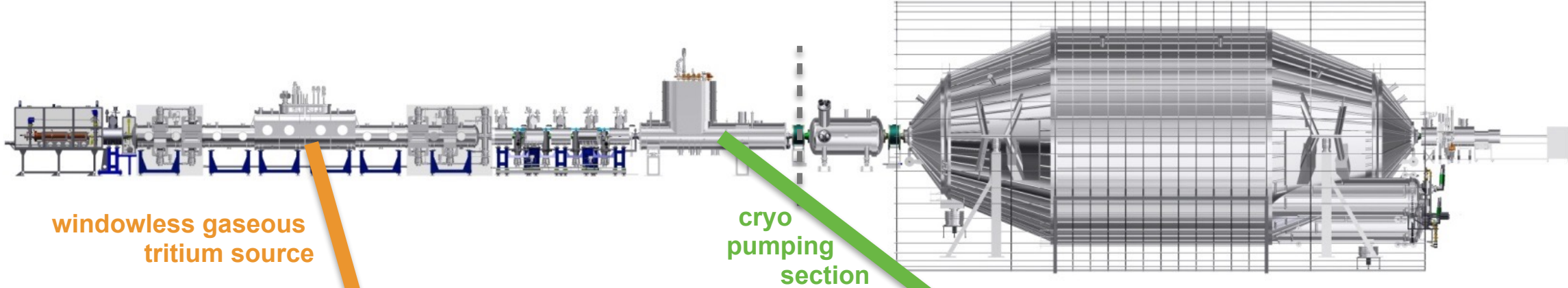
# System integration and commissioning



cryo  
pumping  
section



# System integration and commissioning



windowless gaseous  
tritium source

cryo  
pumping  
section

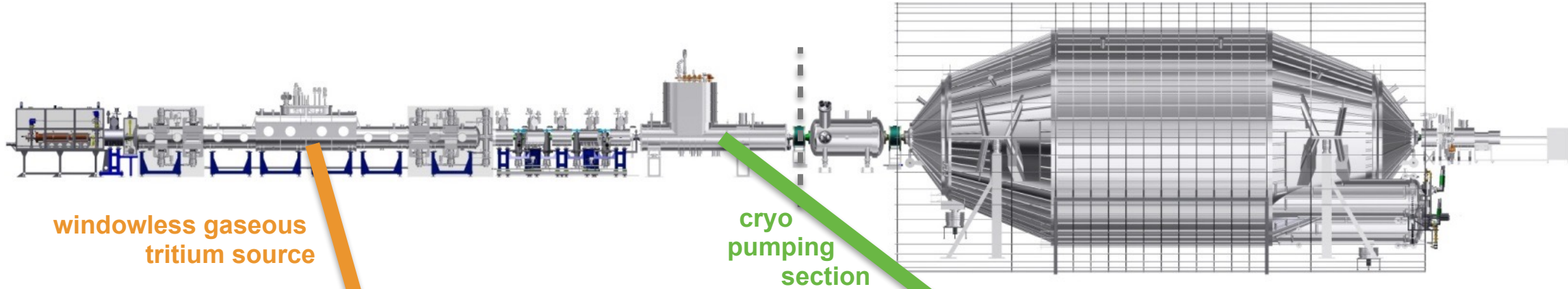
Sept 10, 2015



July 30, 2015



# System integration and commissioning



windowless gaseous tritium source

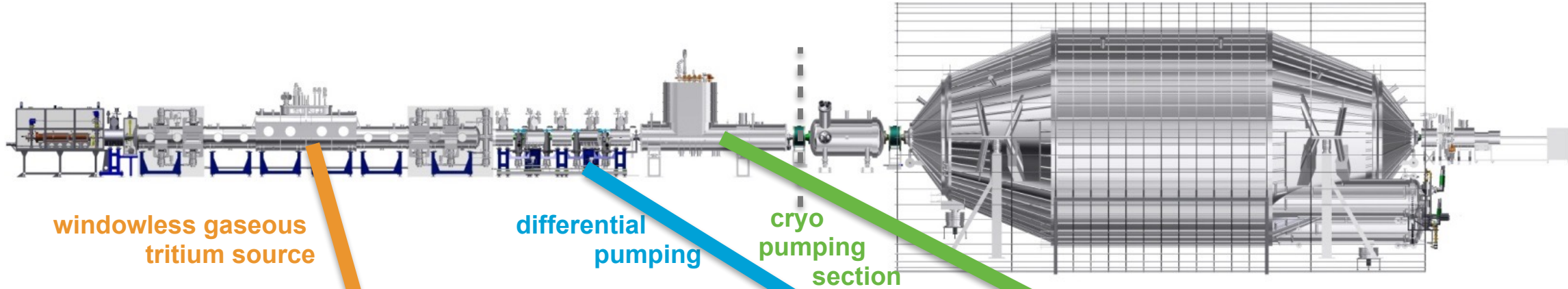
cryo pumping section



Summer 2015:  
Arrival of **last two** major system components on site



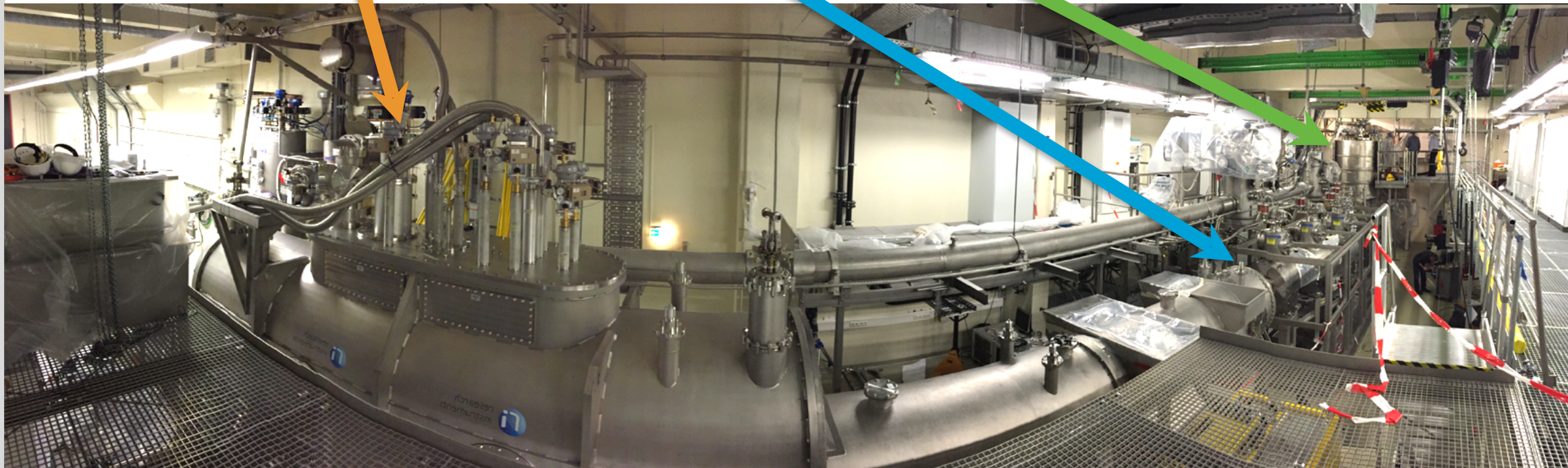
# System integration and commissioning



windowless gaseous  
tritium source

differential  
pumping

cryo  
pumping  
section

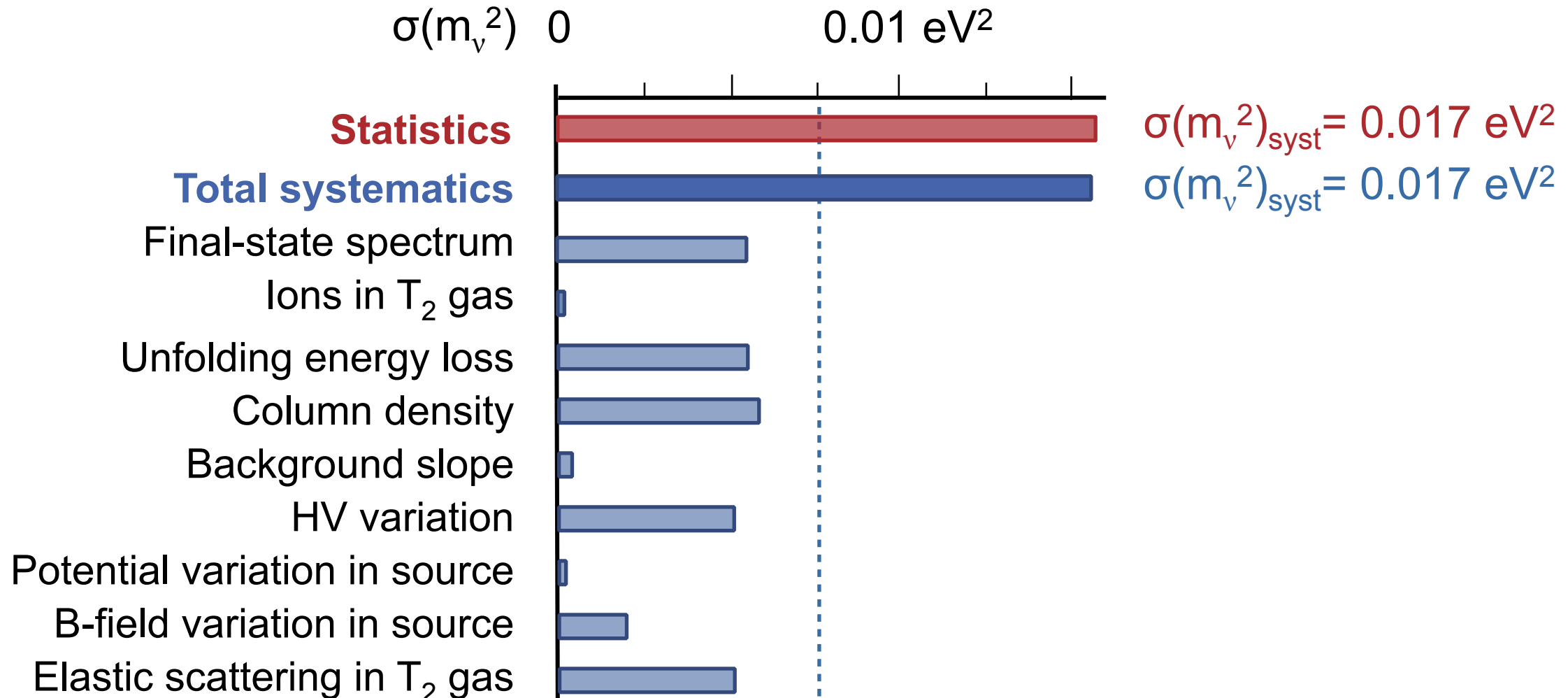


Sept. 2015: **Major milestone achieved**  
full Source and Transport Section in place



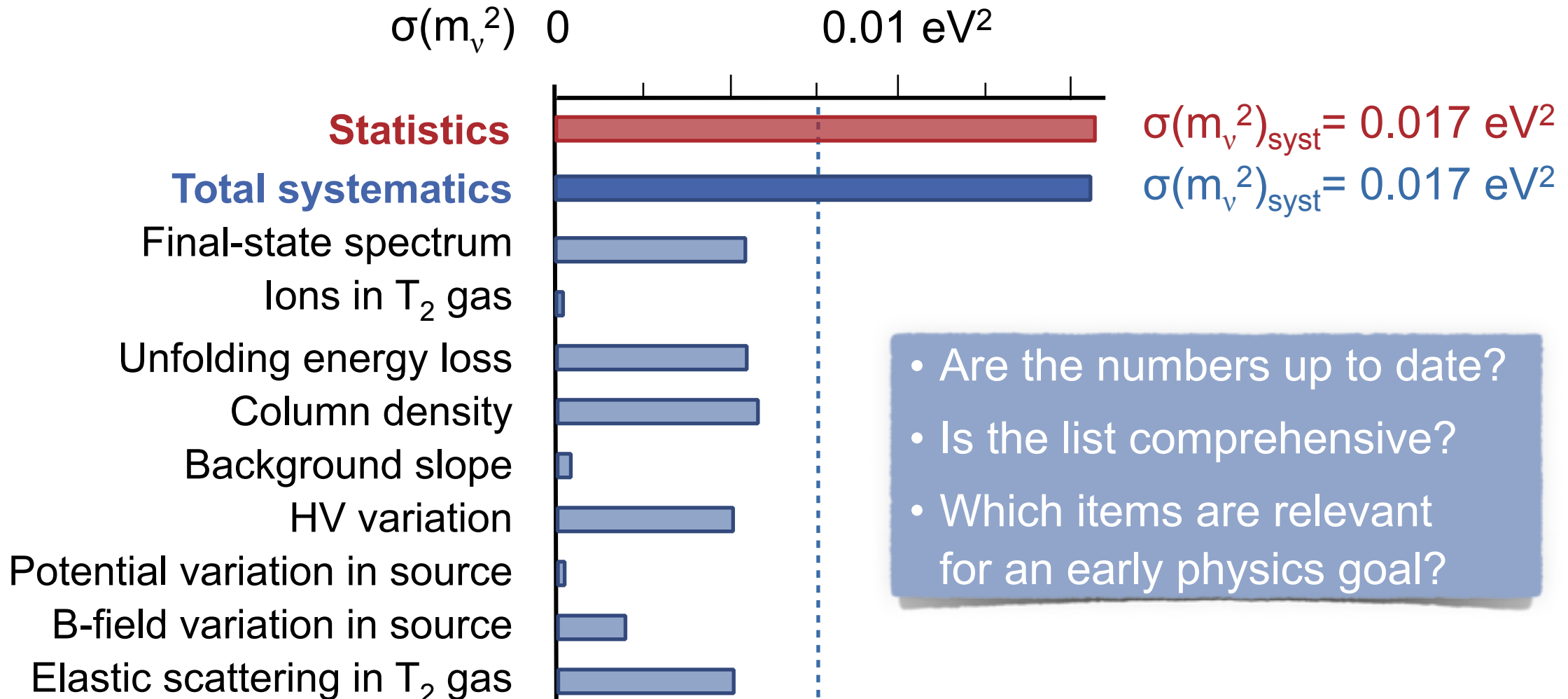
# KATRIN's systematic uncertainty "budget"

- Careful, conservative evaluation in KATRIN Design Report (2004)
- Dominant contributions by source-related effects



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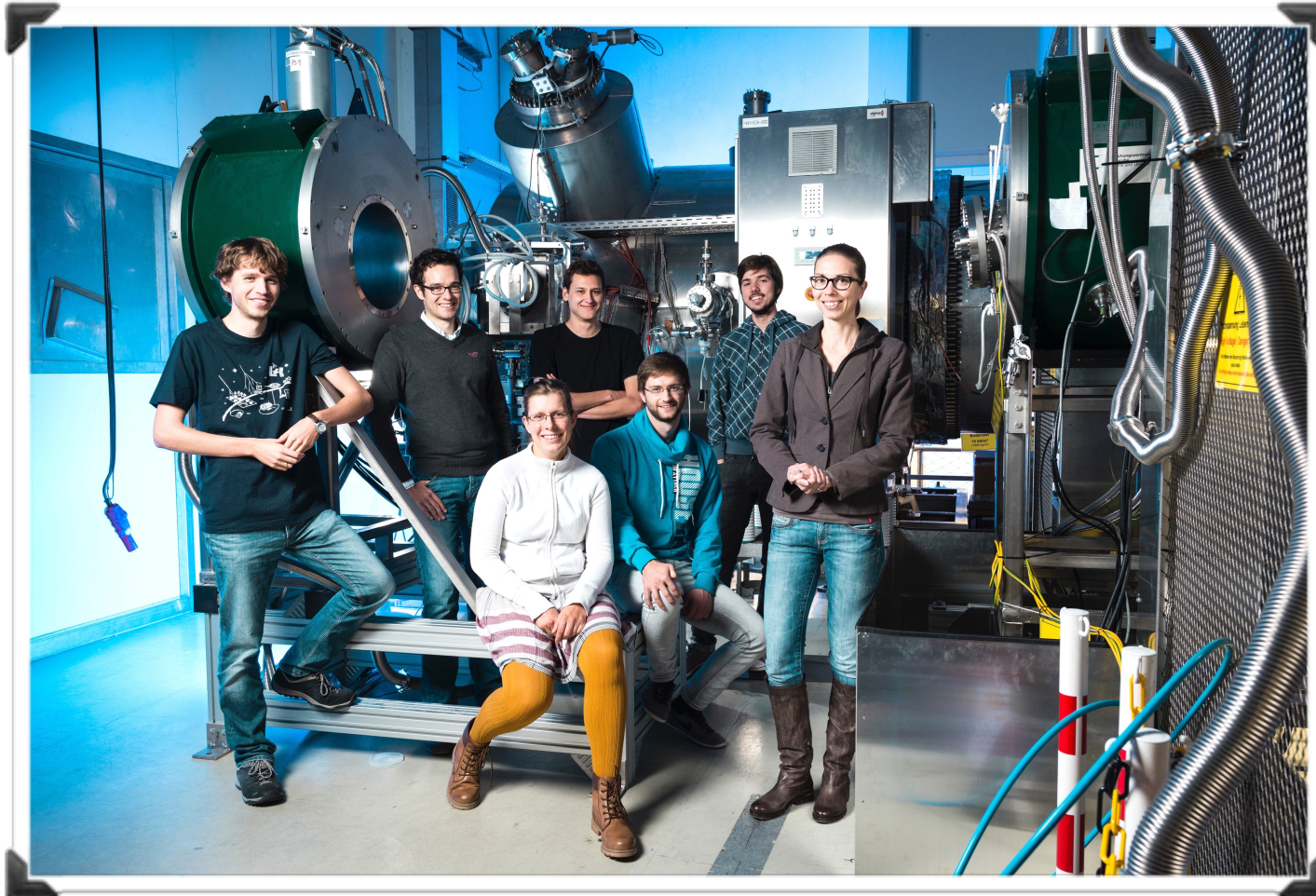
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- Are the numbers up to date?
- Is the list comprehensive?
- Which items are relevant for an early physics goal?

# Helmholtz-University Young Investigators Group (est. 2014):

*“Analysis of KATRIN data to measure the neutrino mass and search for New Physics”*



## Group members

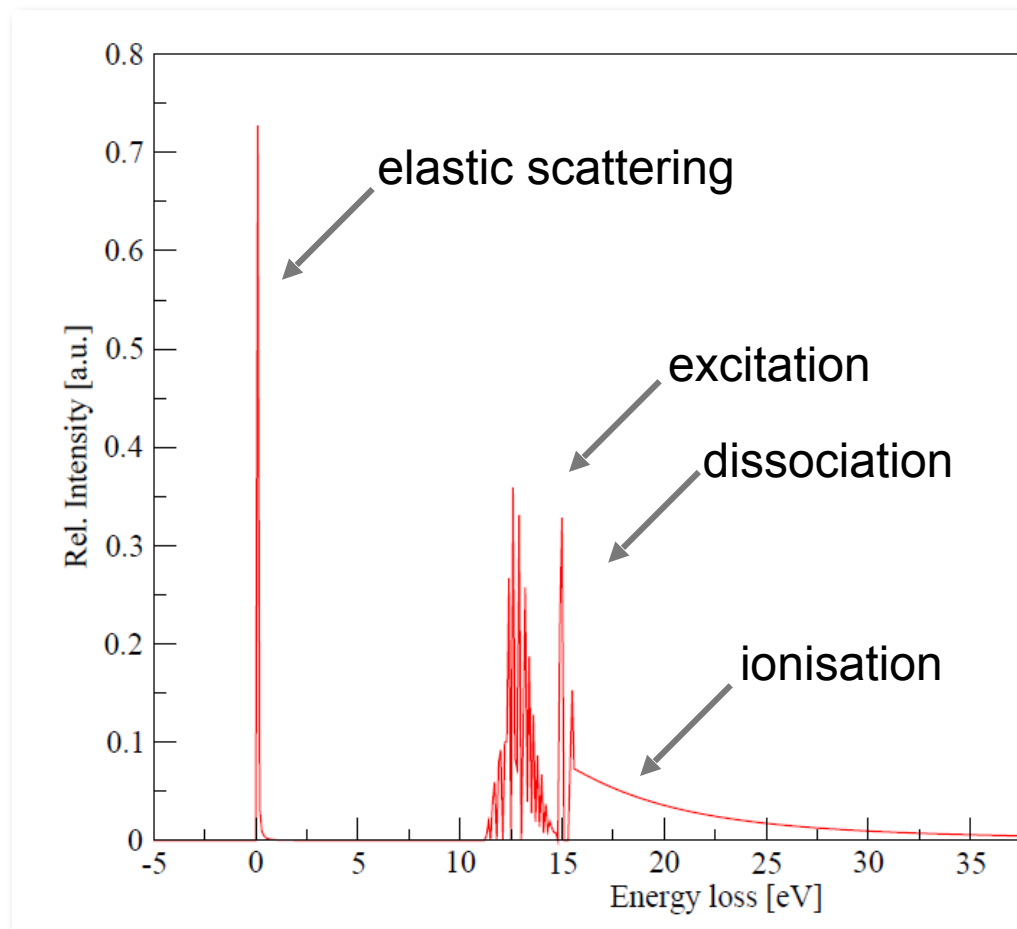
- *K. Valerius*  
group leader
- *M. Kleesiek*  
postdoc
- *L. Kuckert*  
PhD student
- *H. Seitz-Moskaliuk*  
PhD student
- *F. Heizmann*  
PhD student
- *M. Klein*  
PhD student
- *M. Machatschek*  
Master's student

## Former group members

- *M. Babutzka*  
postdoc
- *R. Combe*  
Master's student
- *J. Antoni*  
Diploma student

# Example: Energy loss function

18.6 keV electrons undergo energy loss when scattering in gaseous  $T_2$  source

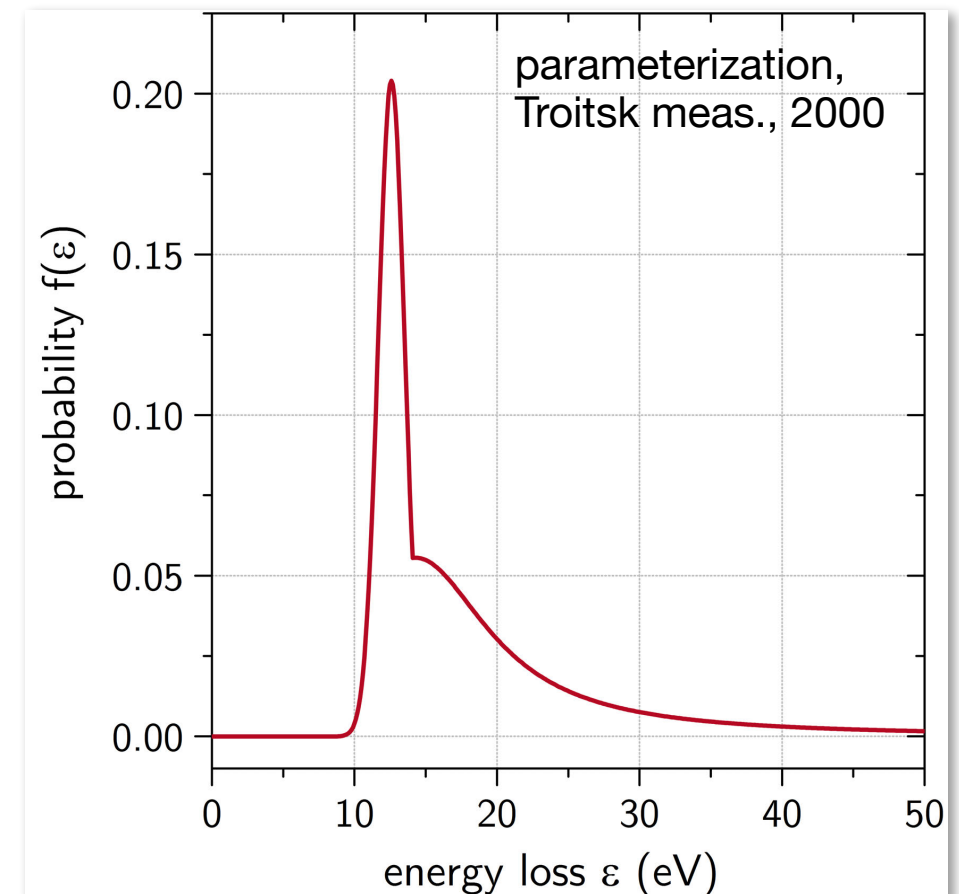
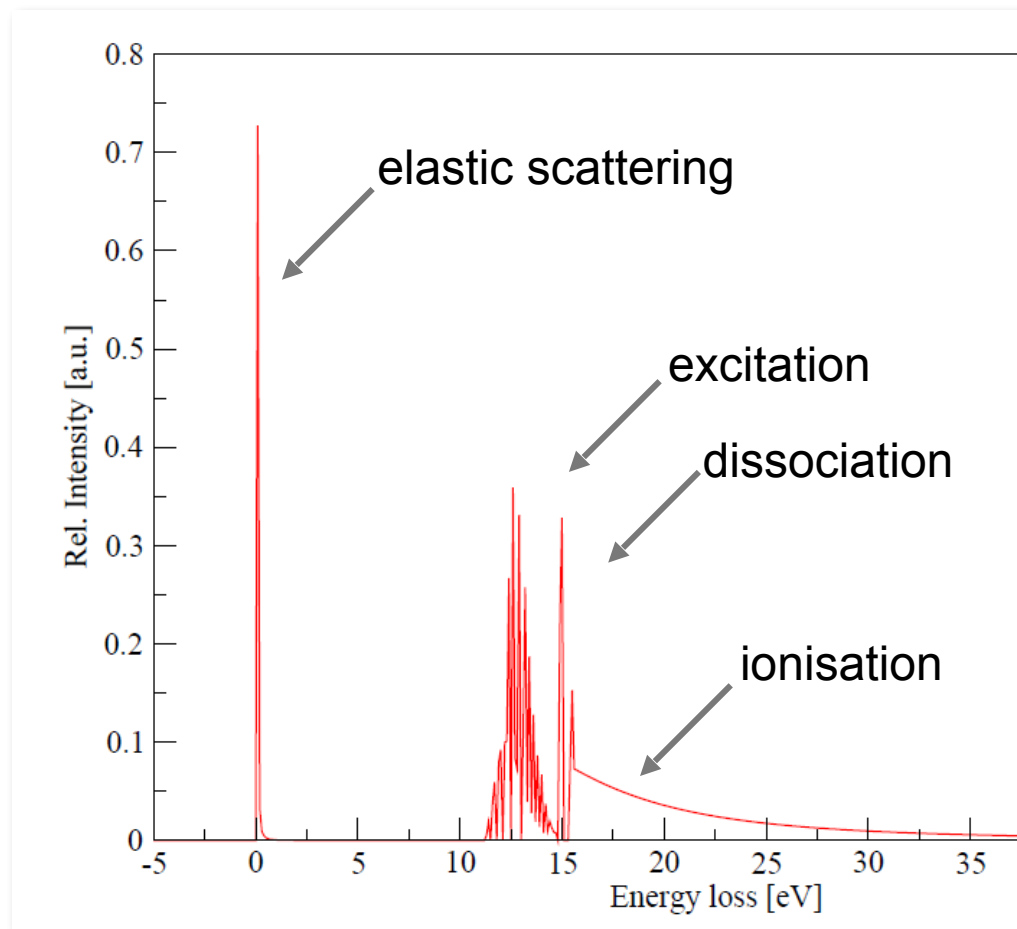




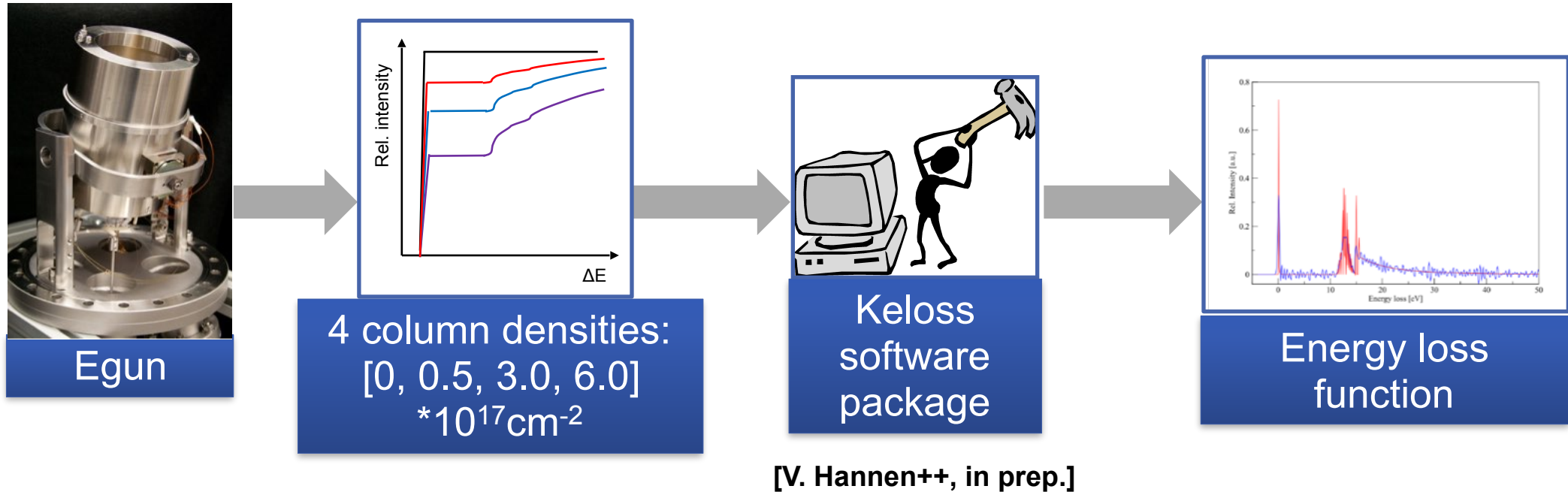
# Example: Energy loss function

18.6 keV electrons undergo energy loss when scattering in gaseous T<sub>2</sub> source

model based on H<sub>2</sub>/D<sub>2</sub> data  
→ *improved* measurement for T<sub>2</sub> necessary



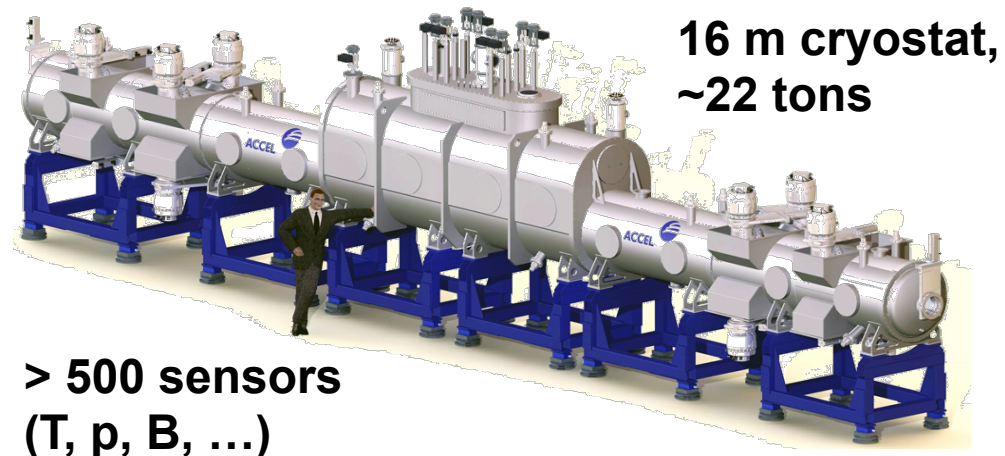
# Energy loss function: measurement



C. Kranz, Diploma thesis 2011

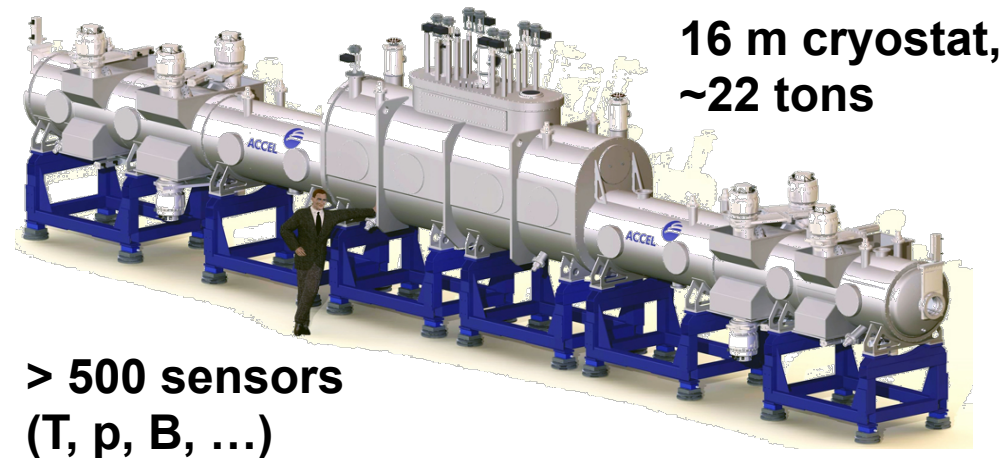
- Work (with V. Hannen) on setting up detailed measurement proposal
- Deconvolution technique accurate enough for KATRIN
- Remaining uncertainties (e.g. column density setting) to be evaluated
- First test with  $\text{D}_2$  suggested to train procedures

# Example: Column density model



- Temperature, pressure, tritium purity to be stabilized at  $10^{-3}$  level
- Small variations of op. parameters lead to fluctuations of column density → syst. influence on  $m^2(v)$

# Example: Column density model

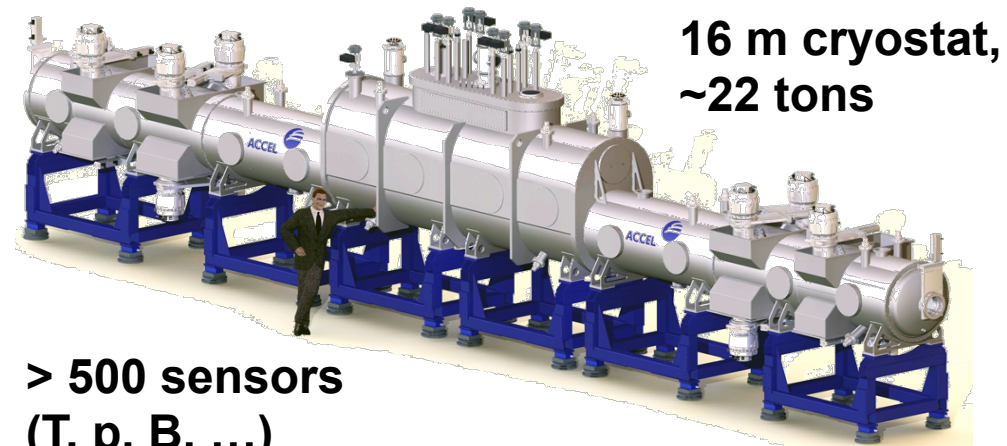


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## Column density monitoring:

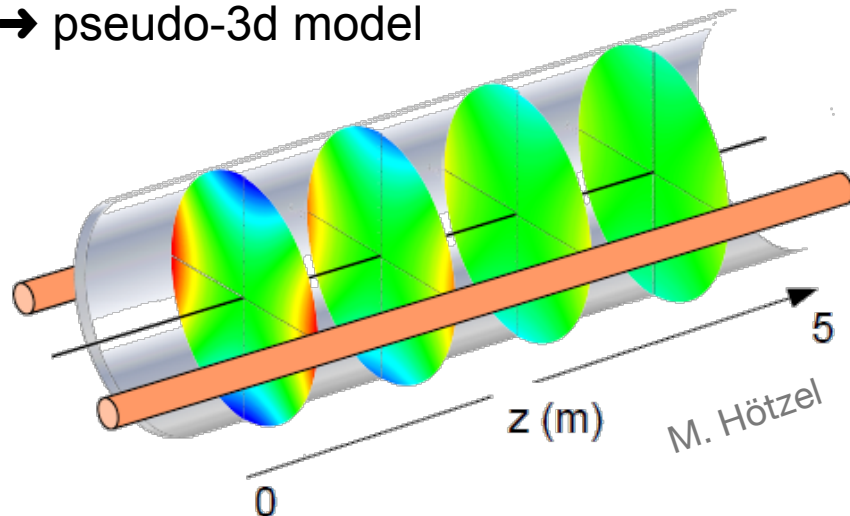
- Small detector in forward direction
- Regular control meas. with e-Gun

# Example: Column density model



> 500 sensors  
(T, p, B, ...)

Gas dynamical model:  
longitudinal profile + 2d slices  
→ pseudo-3d model



- Temperature, pressure, tritium purity to be stabilized at  $10^{-3}$  level
- Small variations of op. parameters lead to fluctuations of column density → syst. influence on  $m^2(v)$

Column density monitoring:

- Small detector in forward direction
- Regular control meas. with e-Gun

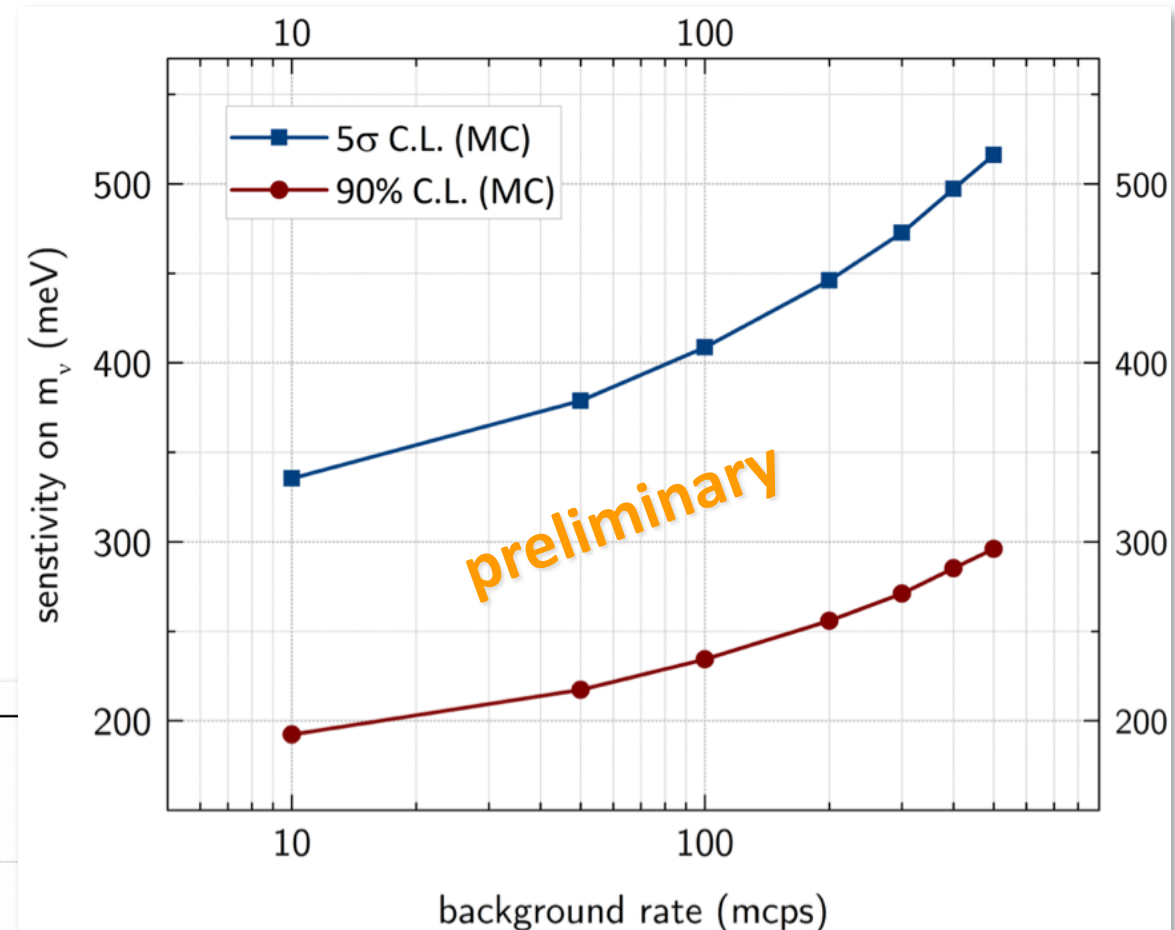
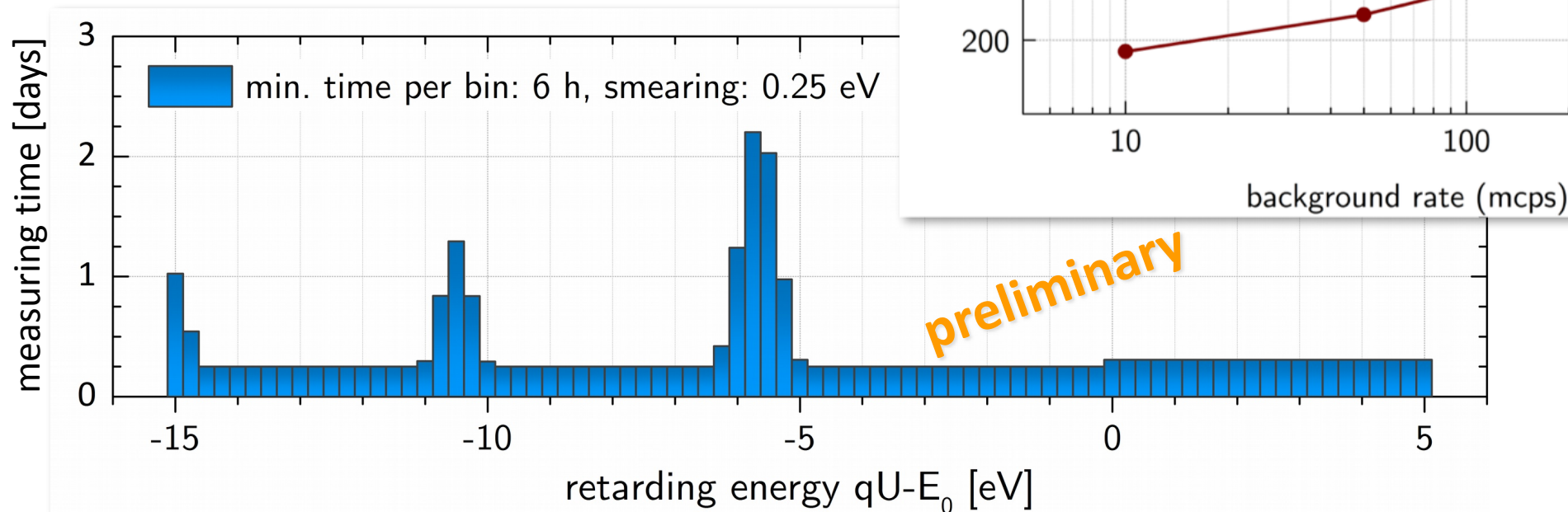
+

- Detailed modeling of **gas dynamics** and resulting spectrum
- Temporal and spatial variations of operational parameters → **sensors**

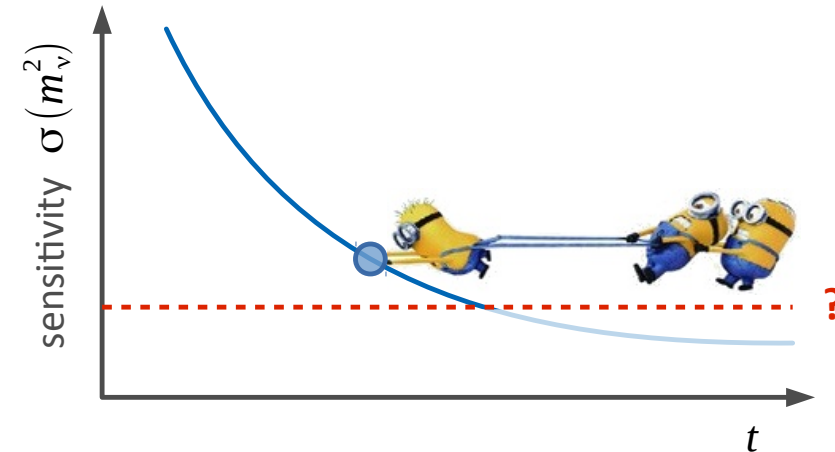
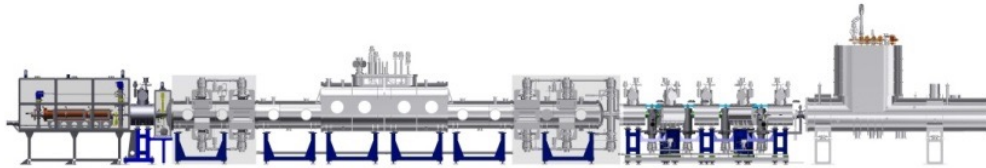
# Example: Sensitivity and background level

Development of  
MC-based tools for  
sensitivity estimates and  
meas. time optimization

Illustration: proposal for **first month**  
of running KATRIN



# Further projects



## Source-related systematics

- Descriptions of el. potential in source
- Plasma effects?
- Efficient retention of tritium ions?

Space charges and el. potential inhomogeneities probed by dispersing  $^{83m}\text{Kr}$  in tritium gas

→ simulation study ongoing

## Technical/Analysis

- Development of high- and medium-level analysis tools
- Planning of commissioning tests during system integration
- ... towards first physics runs with KATRIN!

**Extra:**

# Exploring KATRIN's physics potential beyond neutrino masses





# KATRIN: $\nu$ -mass sensitivity ... and more:

Explore physics potential

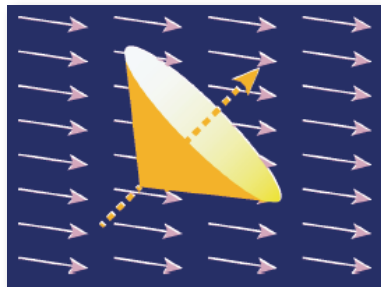
- close to the **spectral endpoint  $E_0$** :

## RH currents

Bonn et al. (2011),  
Barry, Heek, Rodejohann (2014)

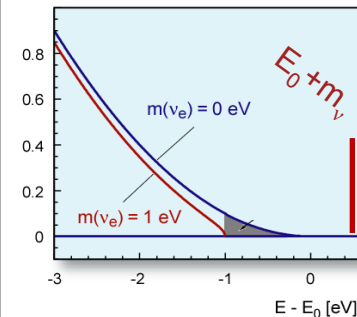
## Violation of Lorentz symmetry

e.g. Diaz, Kostelecky & Lehnert (2013)



## Constraining local C $\nu$ B

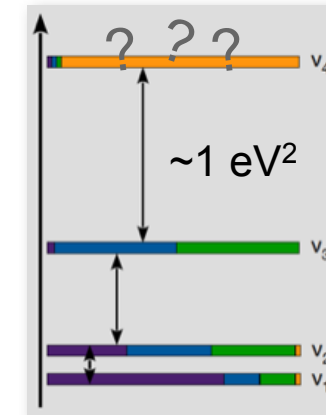
e.g. Kaboth & Formaggio (2010),  
Fässler et al. (2013)



capture of  
relic  $\nu$  on  
 $\beta$ -unstable  
nuclei

## Search for eV-scale sterile $\nu$

e.g. Formaggio & Barrett (2011)



standard operation  
mode for KATRIN

- and **further away from  $E_0$** :

search for keV-mass scale sterile  $\nu$  as WDM candidates

Mertens et al. (2015); Steinbrink et al. (2014)

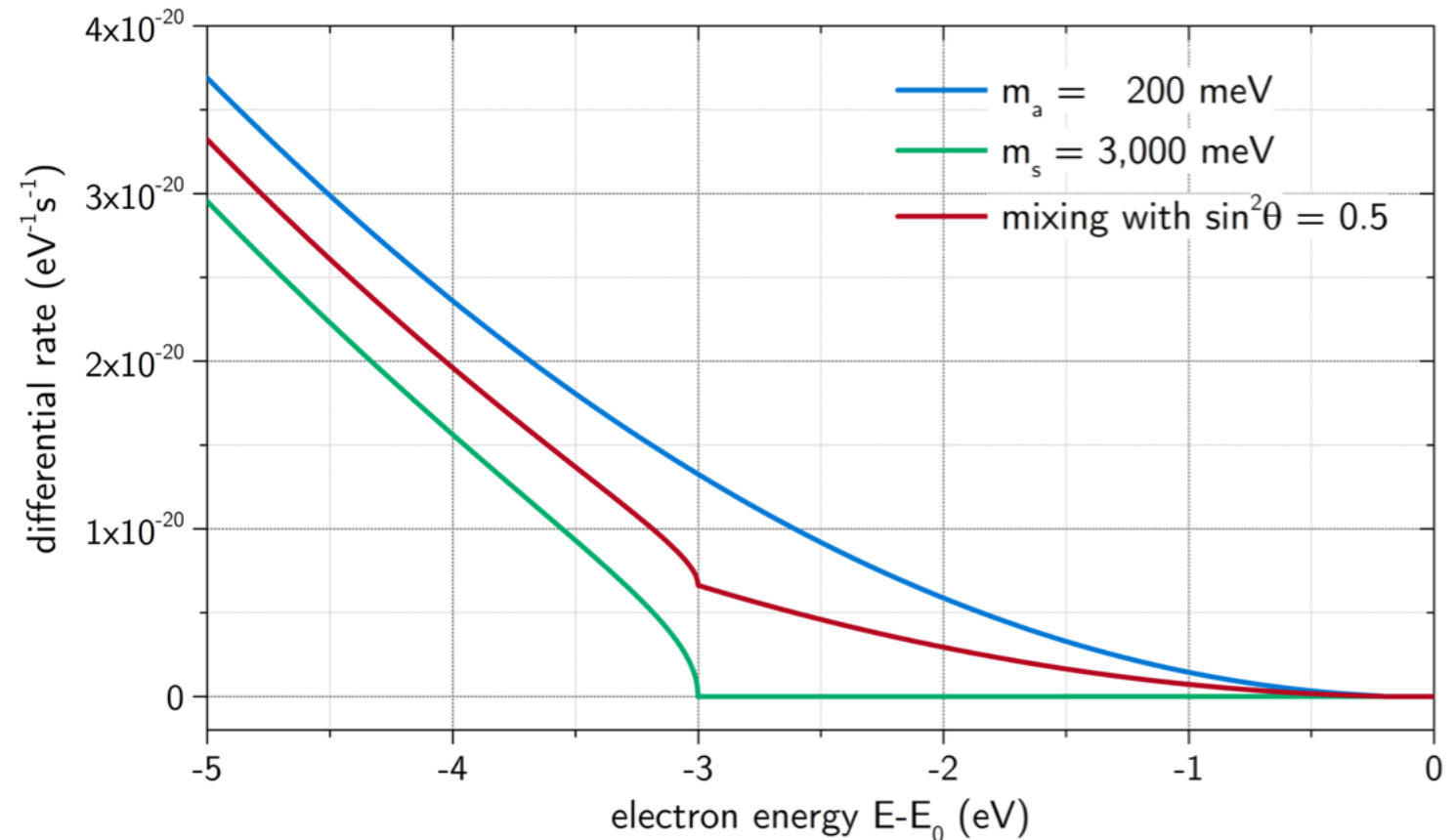
non-standard  
operation, novel  
detector concepts

# Imprint of sterile neutrinos on $\beta$ spectrum

Shape modification below  $E_0$  by active  $(m_a)^2$  and sterile  $(m_s)^2$  neutrinos:

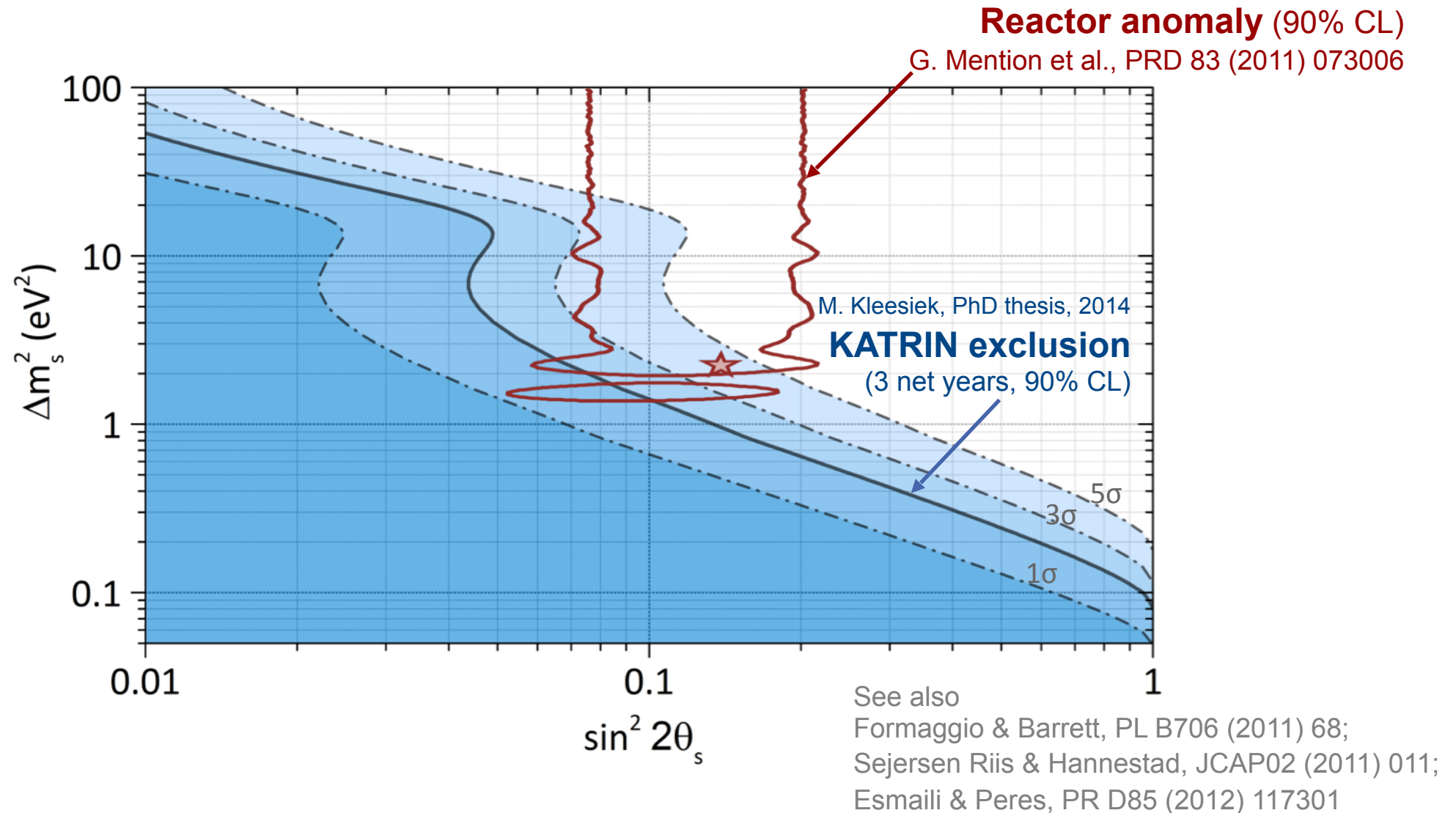
$$\frac{dN}{dE} = \cos^2 \theta_s \frac{dN}{dE}(m_a^2) + \sin^2 \theta_s \frac{dN}{dE}(m_s^2) \quad \longrightarrow \quad \text{additional kink in } \beta \text{ spectrum at } E = E_0 - m_s$$

**example:**  
light sterile  $\nu$   
 $m_s = 3 \text{ eV}$



# Search for eV-scale sterile $\nu$ with KATRIN

- “Reactor antineutrino anomaly”:  $|\Delta m_s^2| > 1.5 \text{ eV}^2$ ,  $\sin^2(2\theta_s) = 0.14 \pm 0.08$  (95% CL)
- Favoured parameter space can be probed by KATRIN:



# Summary & Outlook

- KATRIN sensitivity on  $m(\nu_e)$ : **200 meV** (90% CL, 3 yrs of data)
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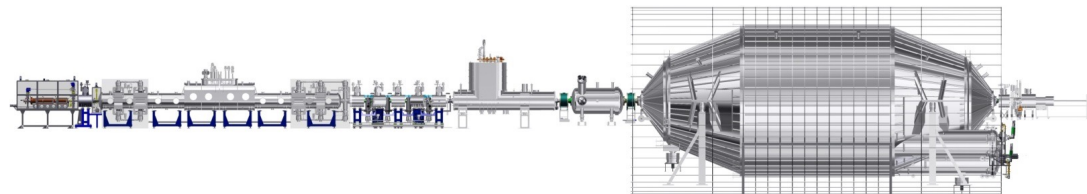
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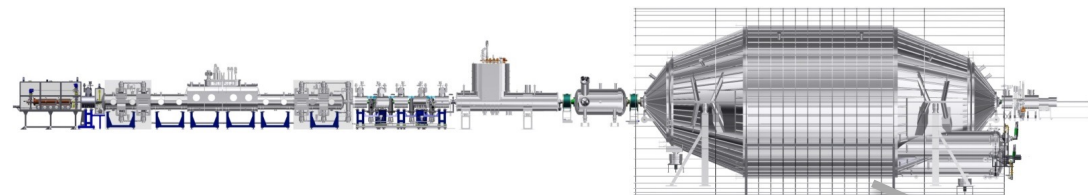
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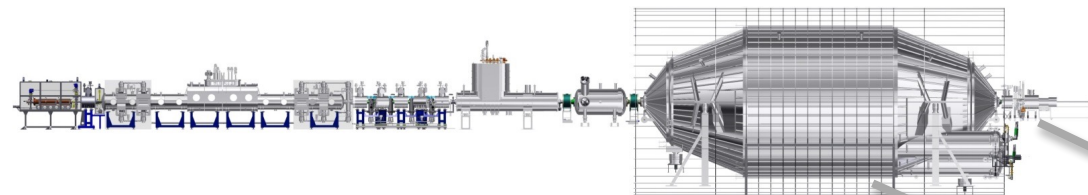
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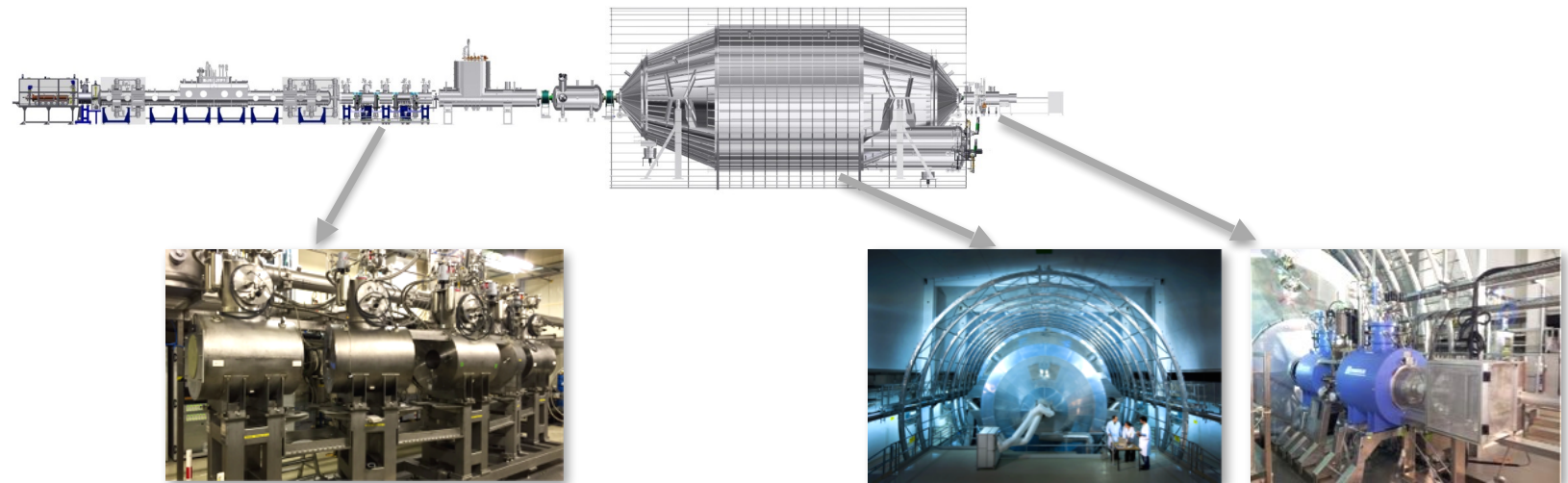
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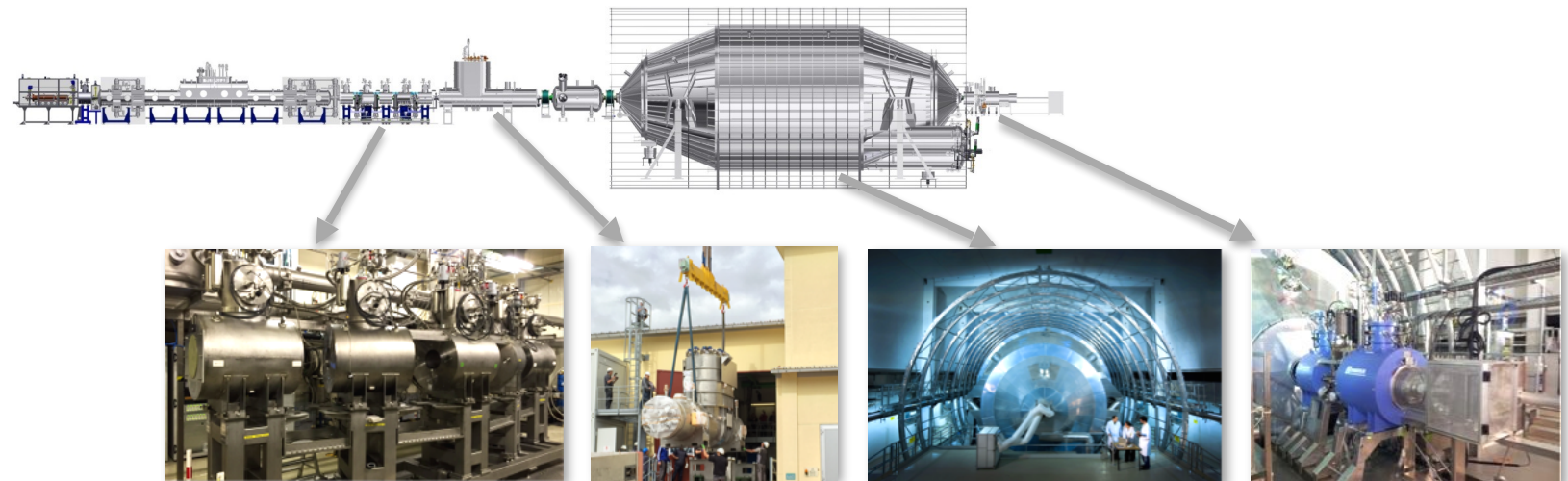
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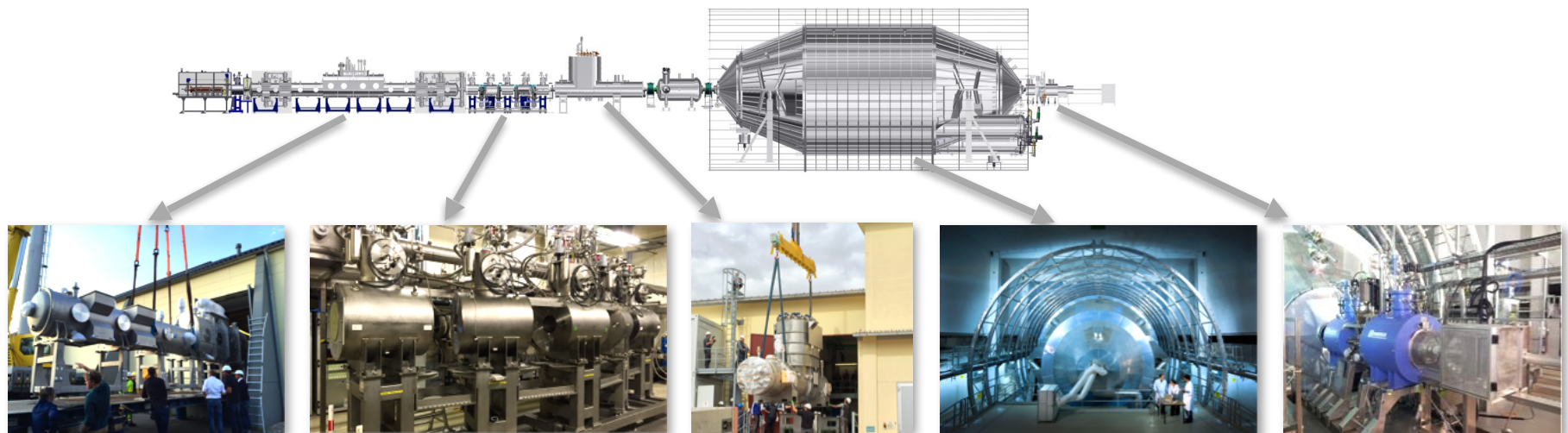
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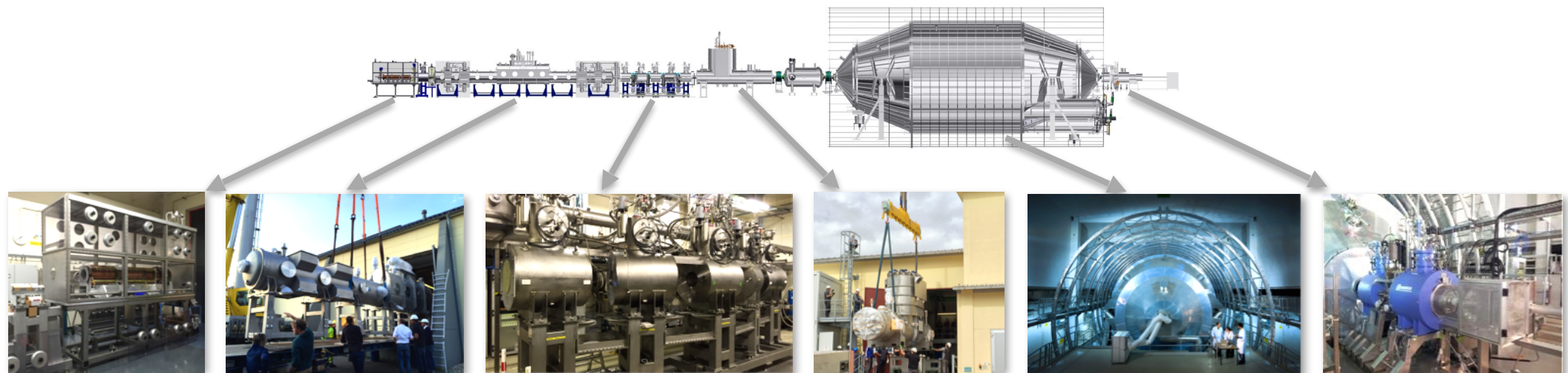
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# Thank you!



# Thank you!



# Supplementing slides

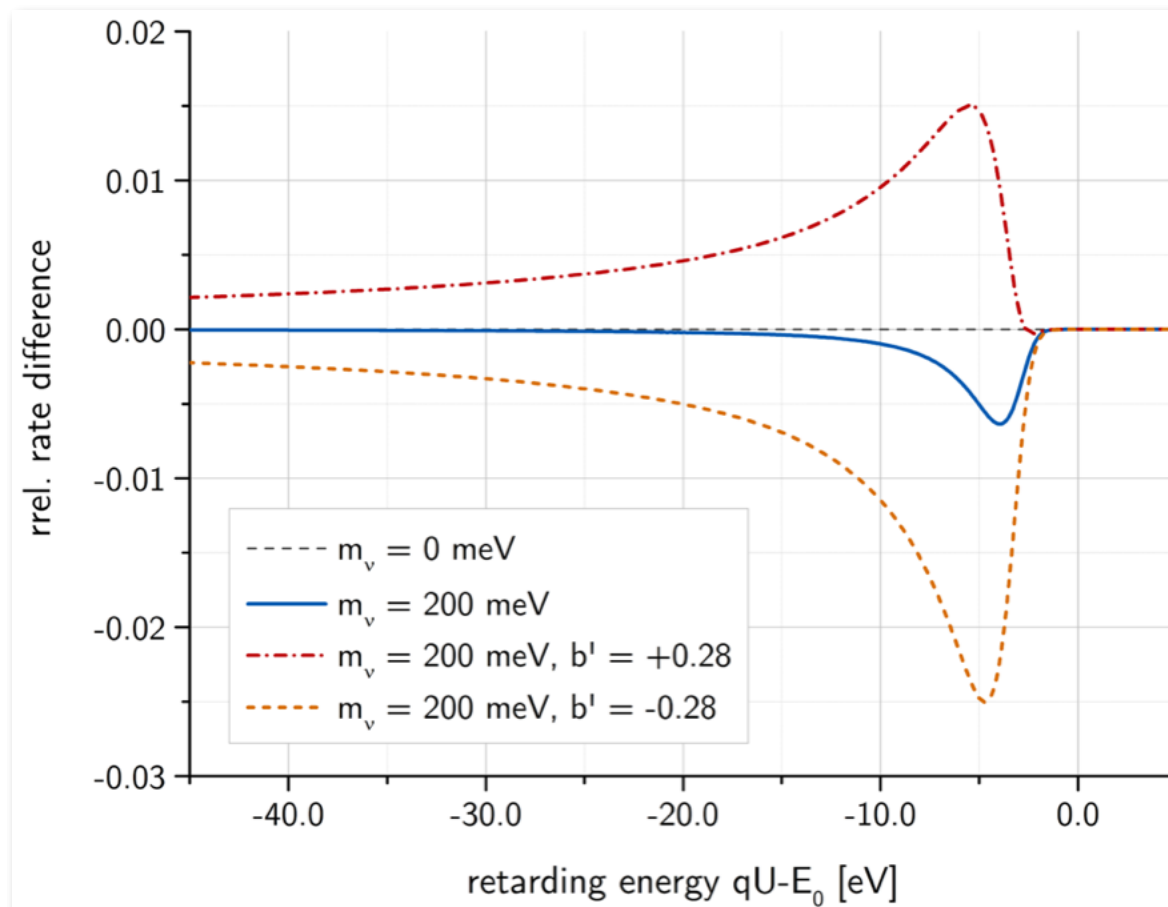
# Effect of RH current contributions

$$\frac{d\Gamma}{dE dt} \propto E_\nu \sqrt{E_\nu^2 - m_\nu^2} \left( 1 + b' \frac{m_\nu}{E_\nu} \right)$$

[J. Bonn et al., *Phys. Lett. B* 703 (2011) 310]

Fierz-like parameter  $b'$  enters differential rate

$$b' \approx -2 \frac{\Re(L_V R_V^* + L_V R_S^*) |\mathcal{M}_F|^2 + \Re(L_A R_A^* + L_A R_T^*) |\mathcal{M}_{GT}|^2}{|L_V|^2 |\mathcal{M}_F|^2 + |L_A|^2 |\mathcal{M}_{GT}|^2}$$



## Imprint on integrated spectrum:

- Only small sensitivity on  $b'$  if endpoint  $E_0$  left free in fit  
 → good for determination of  $m^2(\nu_e)$
- Improvement of present bounds on  $b'$  with KATRIN for small  $m(\nu_e)$  if
  - external  $E_0$  value with accuracy  $< 50$  meV as input\*
  - absolute energy scale in KATRIN  $U_{\text{spec}} - U_{\text{source}}$  known to same accuracy of  $< 50$  meV



# Probing Lorentz invariance in $\beta$ decay

[Kosteletzky & Mewes (2004, 2009)]

## Standard Model Extension (SME) framework:

Neutrinos satisfy Dirac-like equation

$$(i\Gamma^\alpha \partial_\alpha - \mathbf{M}) \psi = 0$$

with  $\Gamma$ ,  $\mathbf{M}$  including momentum-dependent coefficients

[review: Diaz (2014)]

## Experimental searches:

- Neutrino oscillations
- Neutrino velocity (ToF)
- Weak decays

probe oscillation-free parameters

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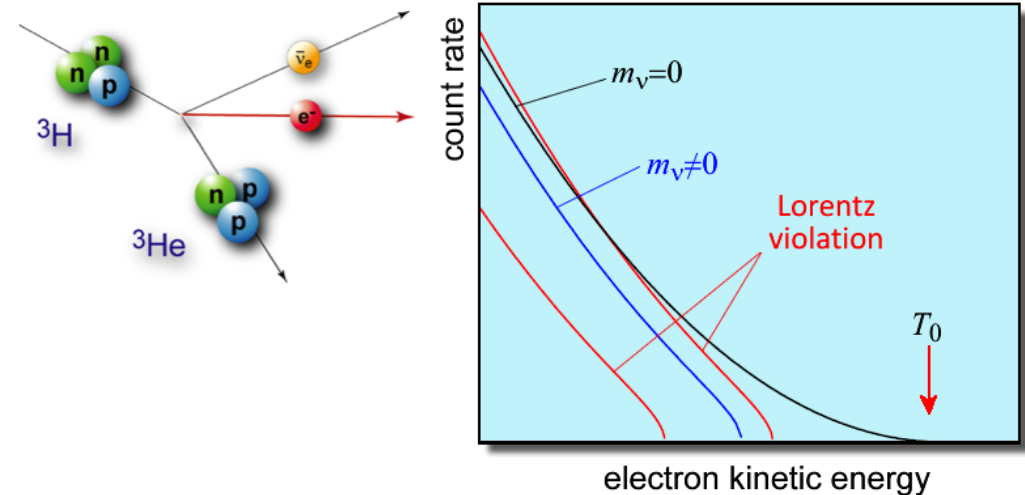
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## Experimental searches:

- Neutrino oscillations
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- Weak decays

probe oscillation-free parameters

## Tritium $\beta$ decay:

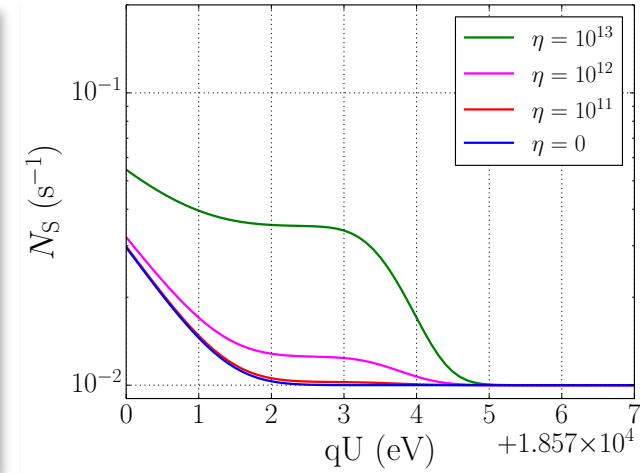
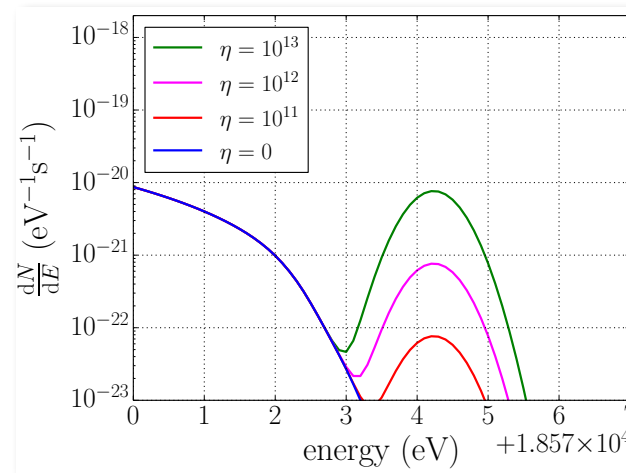
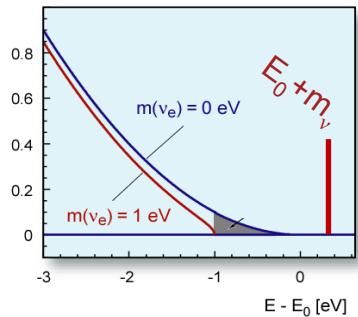
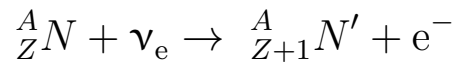


- Modified energy dependence of decay rate
- Spectral shape dependent on sidereal time and experiment orientation
- Effective dim-3 coefficient: osc. shift of endpoint  $T_{0,\text{eff}}$  with  $\omega_{\text{sidereal}}$
- Effective dim-2 coefficient: osc. of  $m^2$  parameter (can mimic tachyonic  $\nu$ )

# Constraining local CvB density with KATRIN

*About every neutrino physicist goes through a phase in his or her career and asks 'There's got to be a way to measure the relic neutrino background' — Peter Fisher*

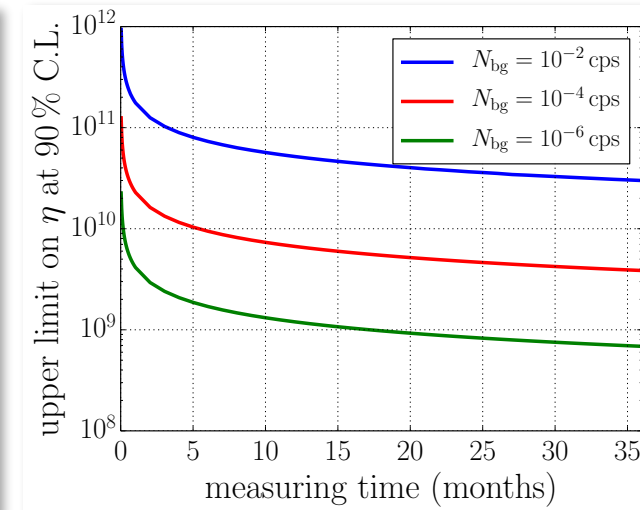
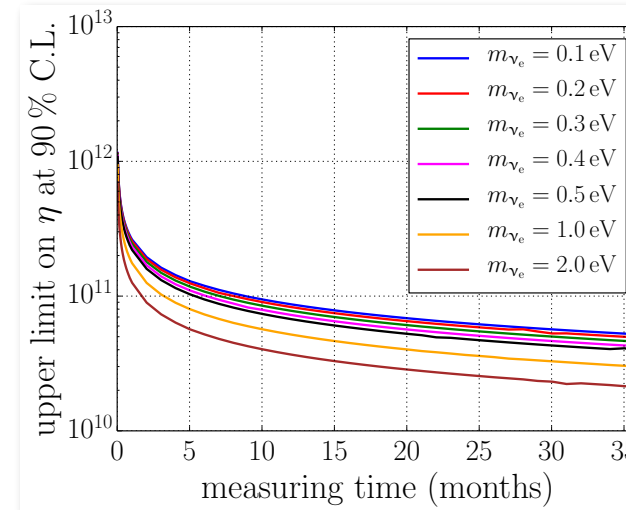
$\nu$  capture on  $\beta$ -instable nuclei:



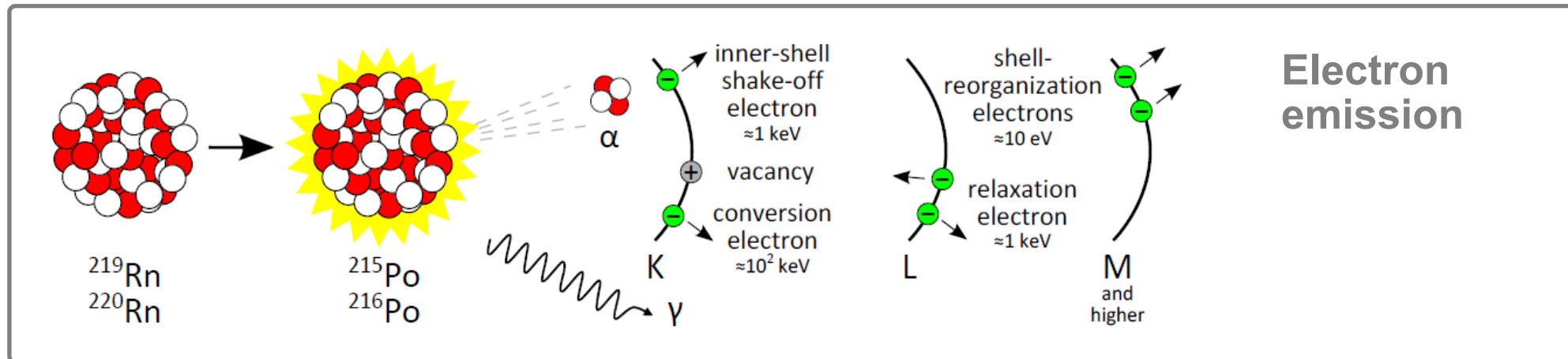
Threshold-free, but tiny cross section  $\sigma(10^{-45} \text{ cm}^2)$

KATRIN reference setup only sensitive to local overdensity:

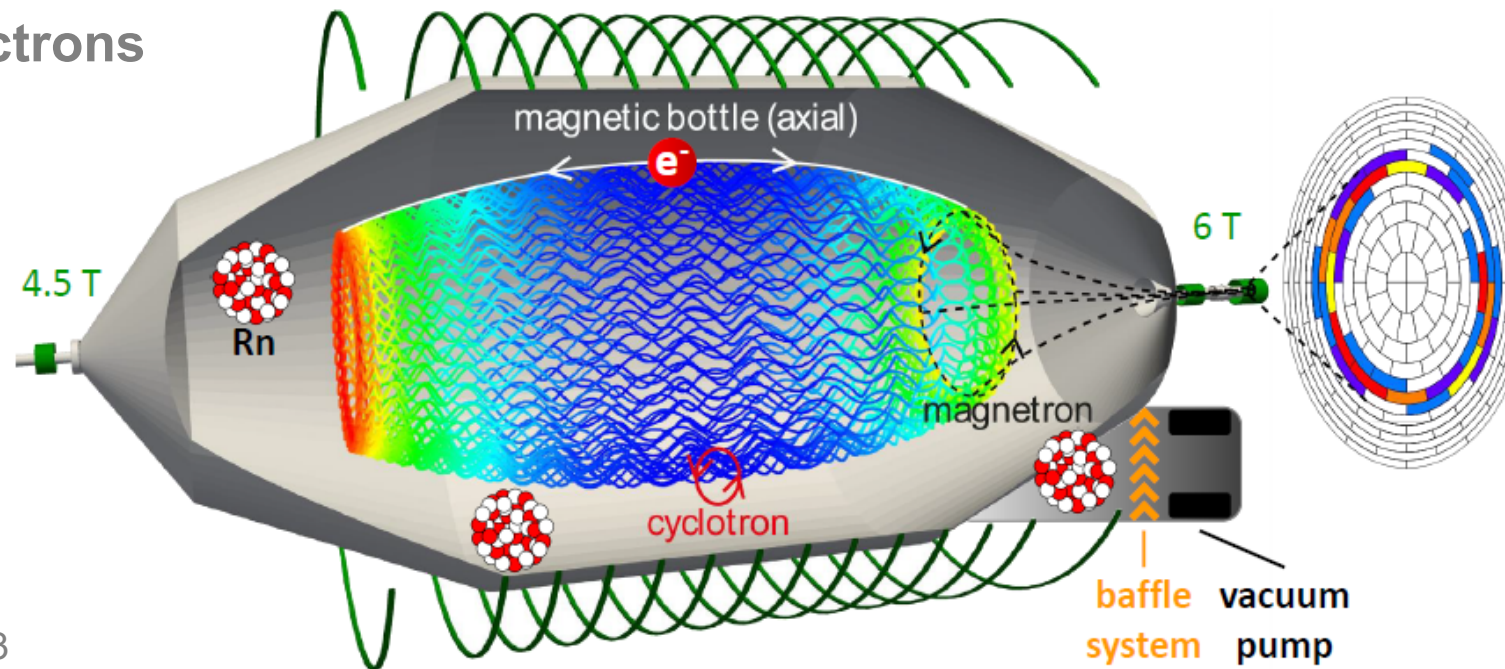
$$\eta = \frac{n_\nu}{\langle n_\nu \rangle}$$



# Radon-induced background



## Trapping of electrons & production of secondaries



Fränkle et al. 2011  
 Mertens et al. 2012  
 Wandkowsky et al. 2013

# Background characterization

2015:

2nd phase of commissioning measurements completed

- > Spectrometer works as MAC-E filter
- > LN2-cooled baffles eliminate Radon-induced background with efficiency of  $(97 \pm 2)\%$
- > Remaining background still under investigation

$$B_{\text{total}} = S_{\text{Rn}} + C_{\text{Rn}} + R$$

$$S_{\text{Rn}} = \alpha \cdot C_{\text{Rn}}$$

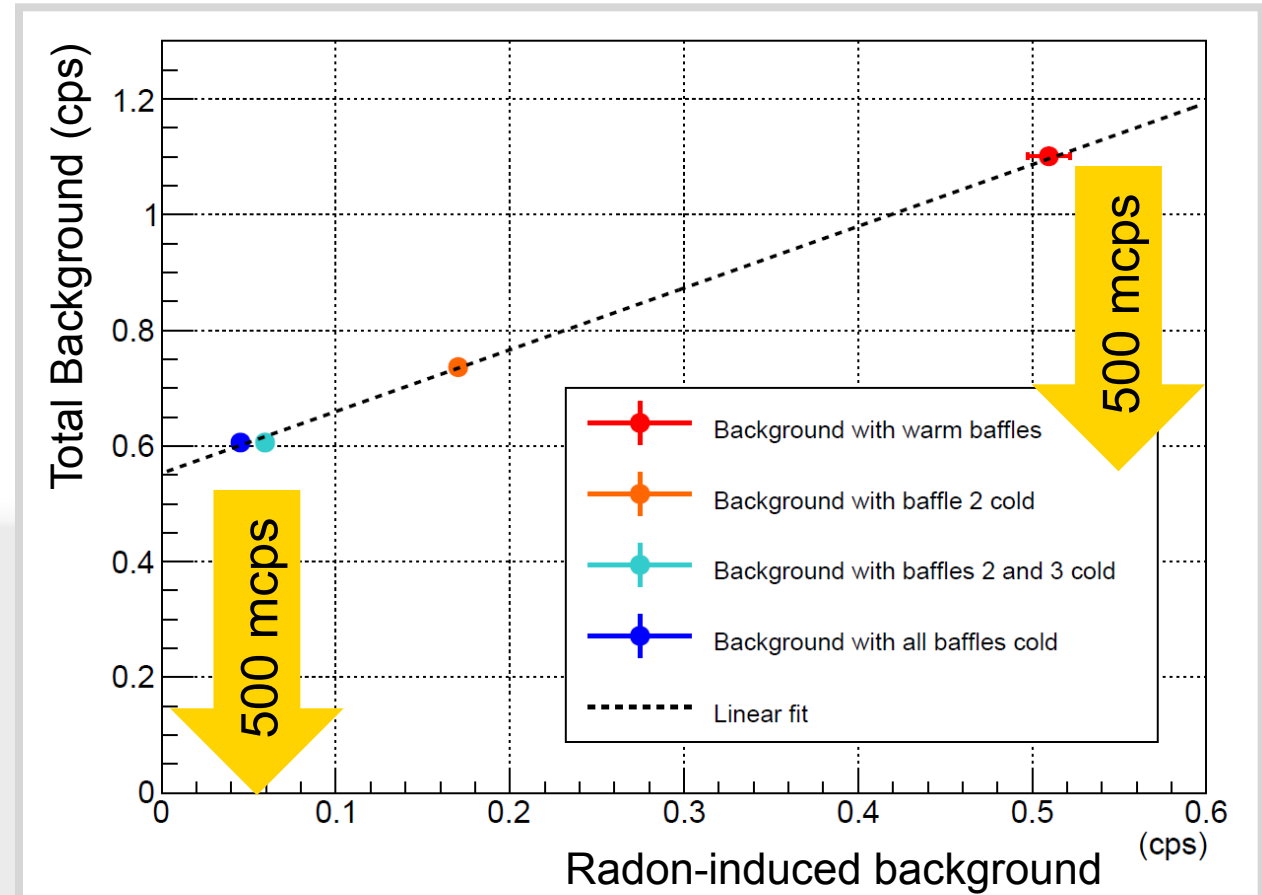
$$B_{\text{total}} = (\alpha + 1) \cdot C_{\text{Rn}} + R$$

$B_{\text{total}}$  : Total background rate

$S_{\text{Rn}}$  : Radon-induced single event rate

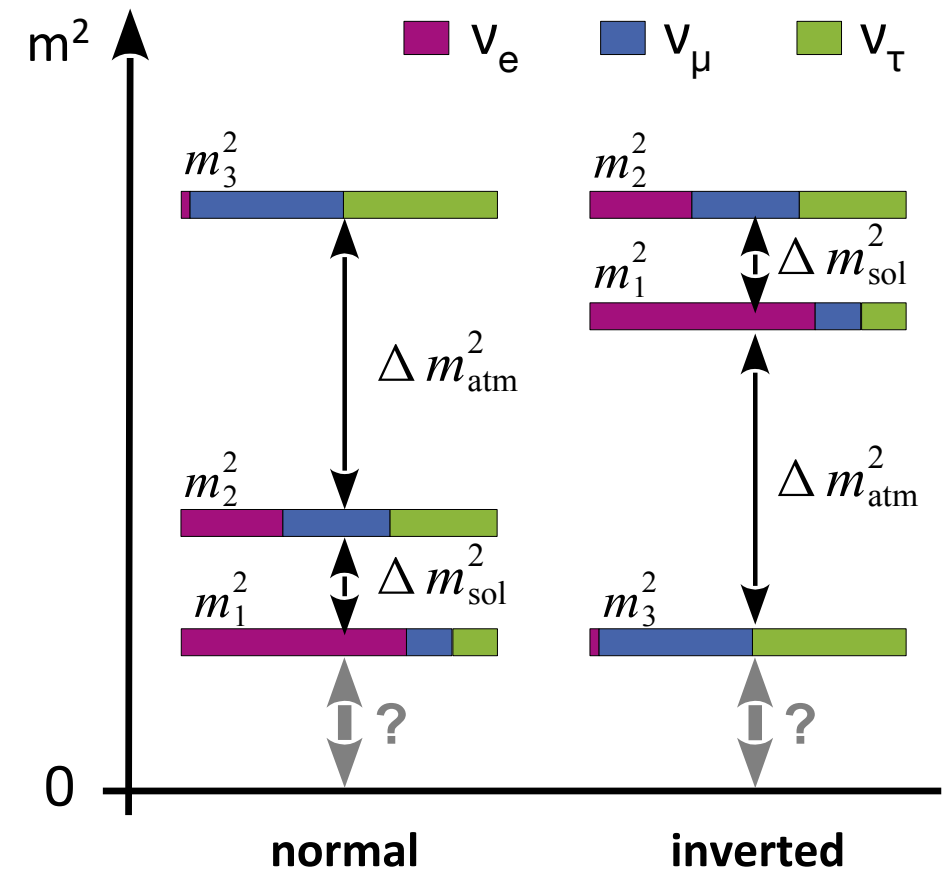
$C_{\text{Rn}}$  : Event rate in Radon-induced clusters

$R$  : Non-Radon-induced bg rate



# Neutrino mixing and mass scheme

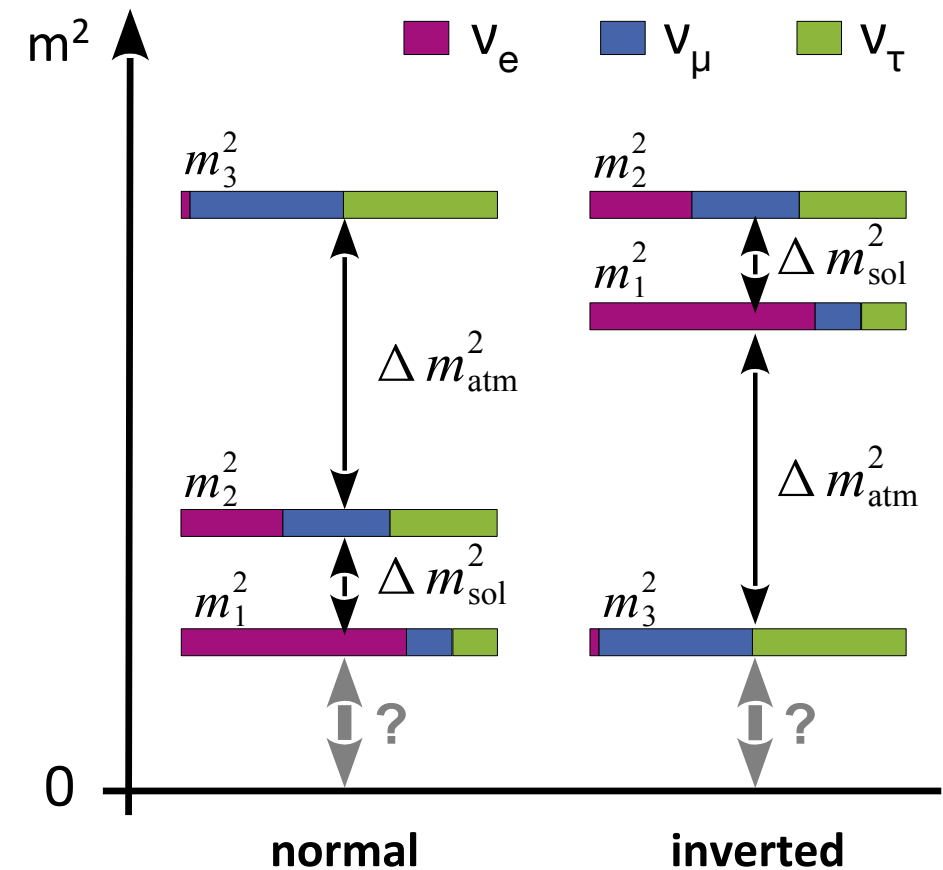
Wealth of  $\nu$  oscillation data:



# Neutrino mixing and mass scheme

## Wealth of $\nu$ oscillation data:

- Large neutrino mixing and tiny neutrino masses  $m(\nu_i) \neq 0$  established

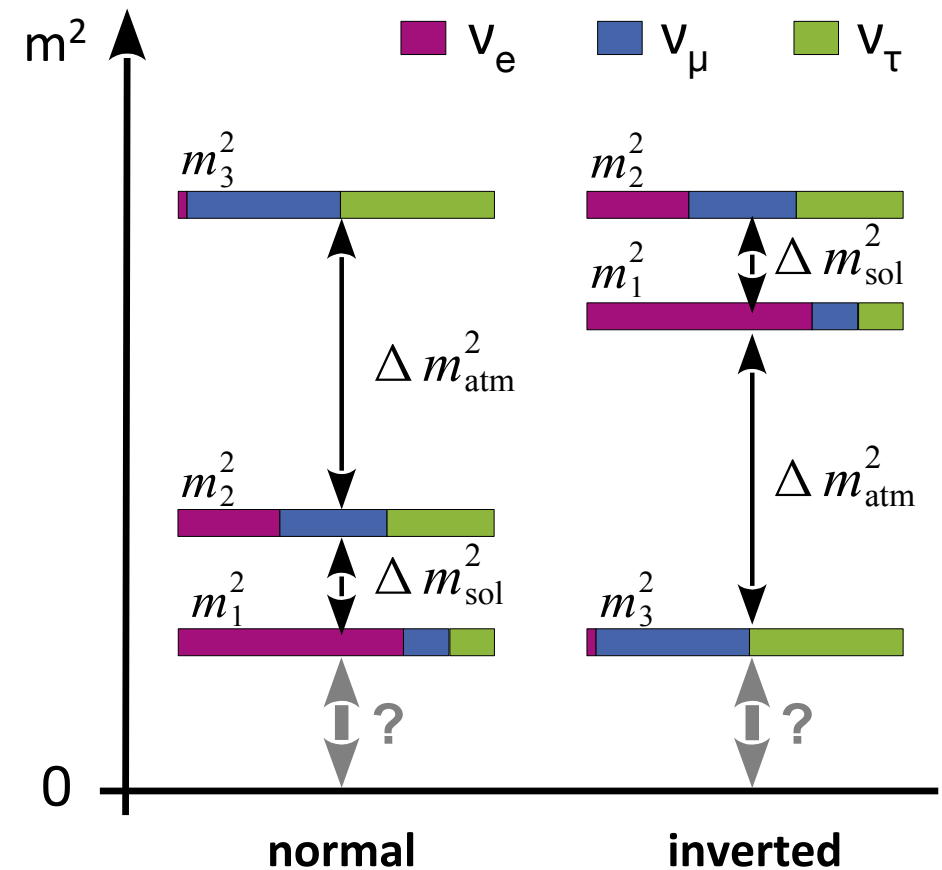


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**New!  
BSM physics!**





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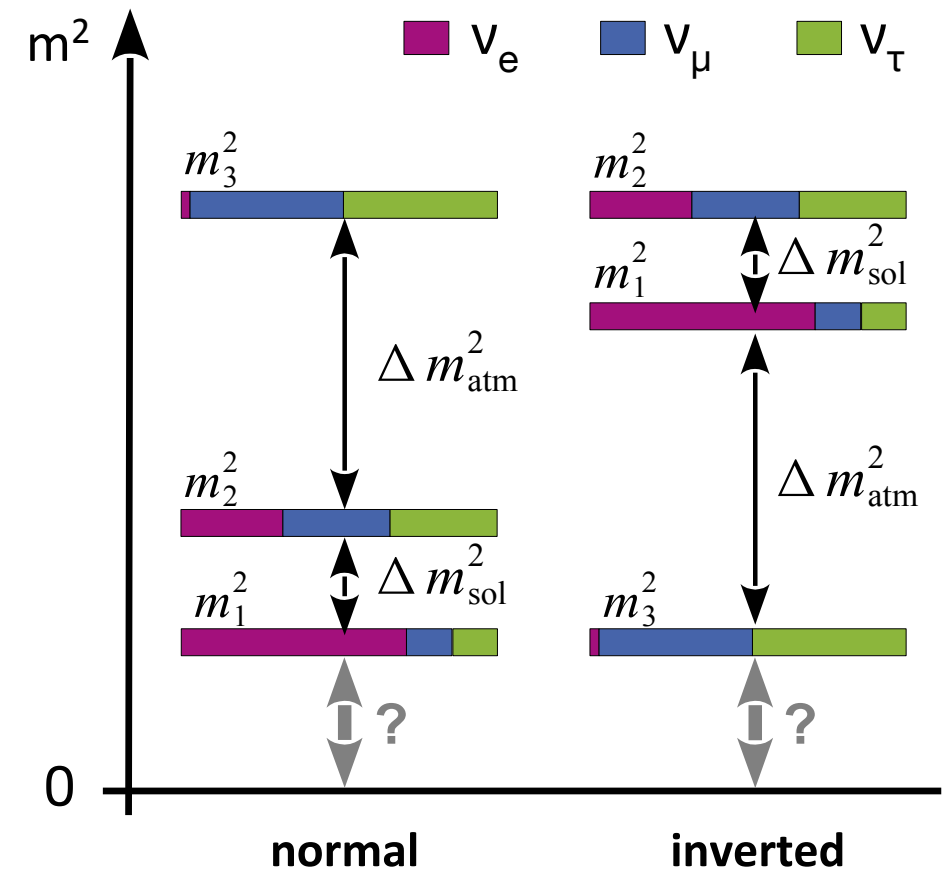
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- Oscillation experiments: only interferometric measurement, no absolute values



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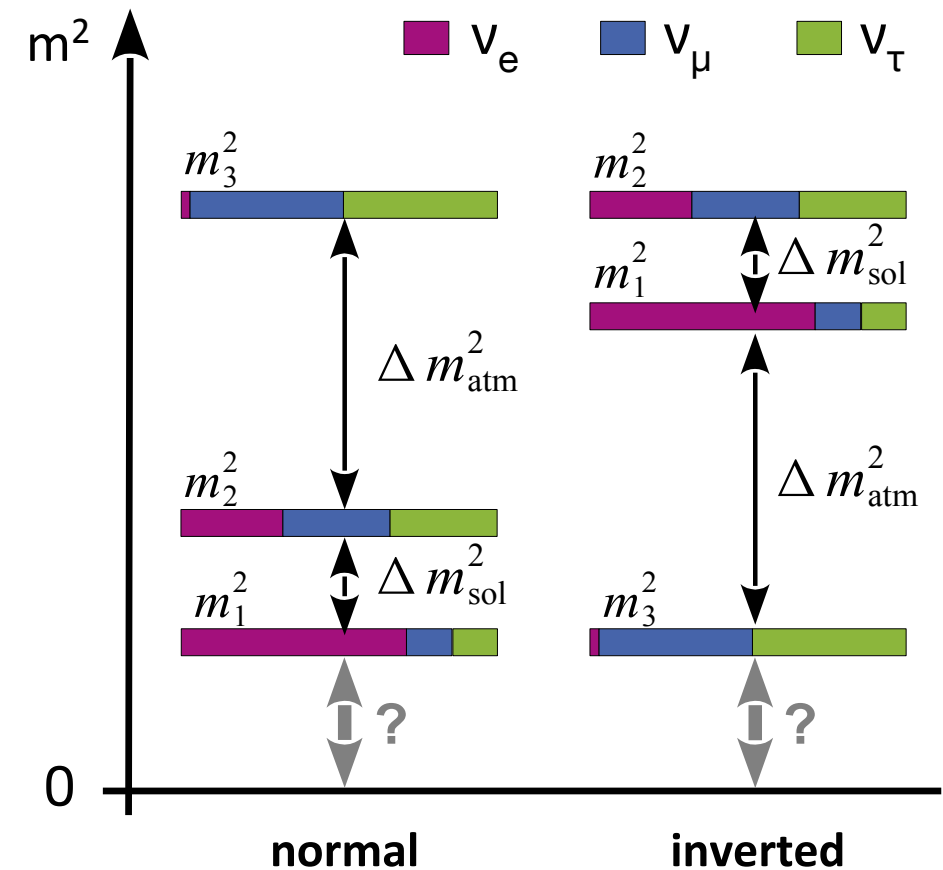
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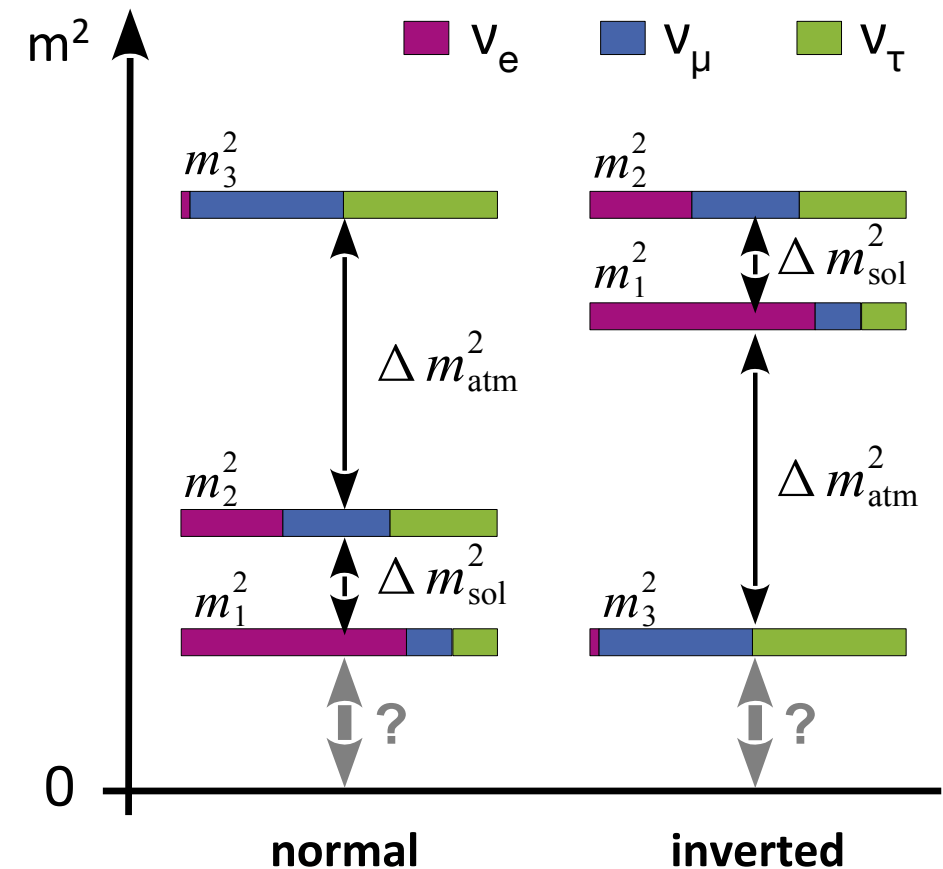
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- What is the absolute  $\nu$  mass scale?



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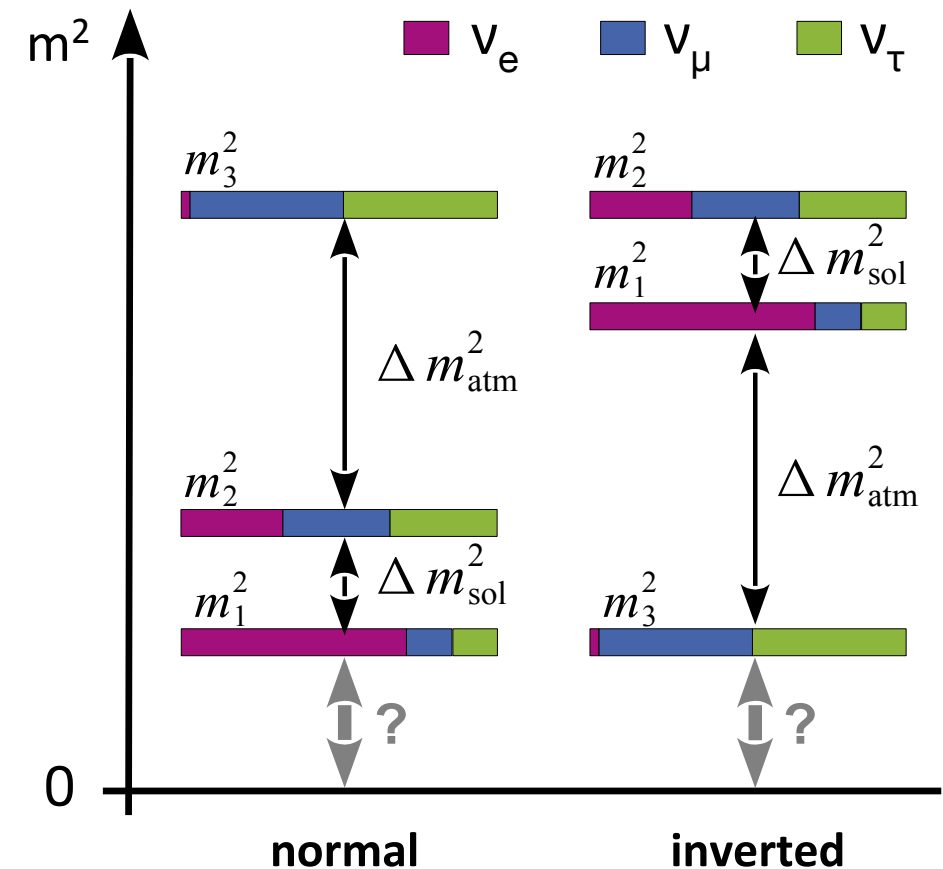
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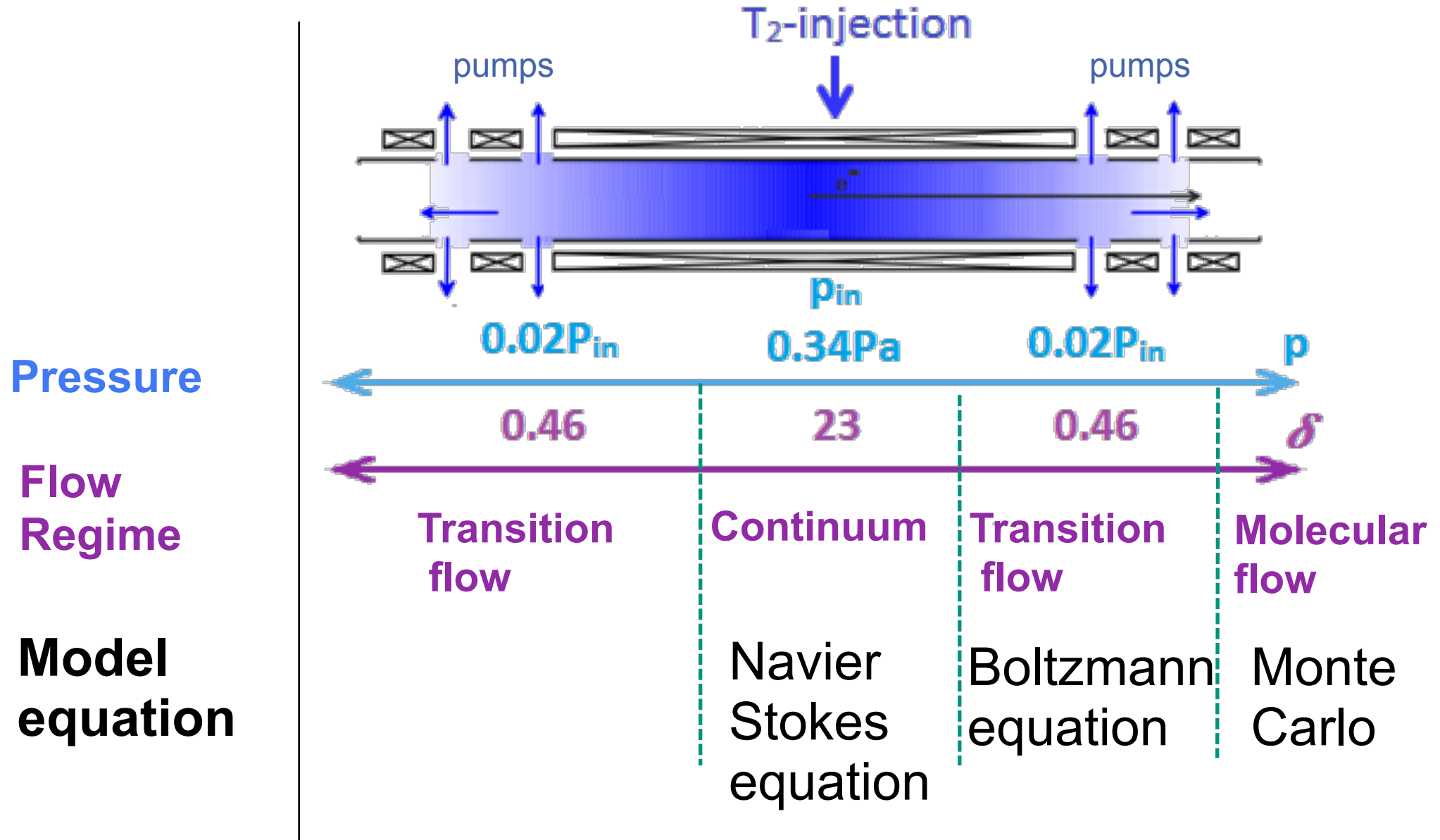


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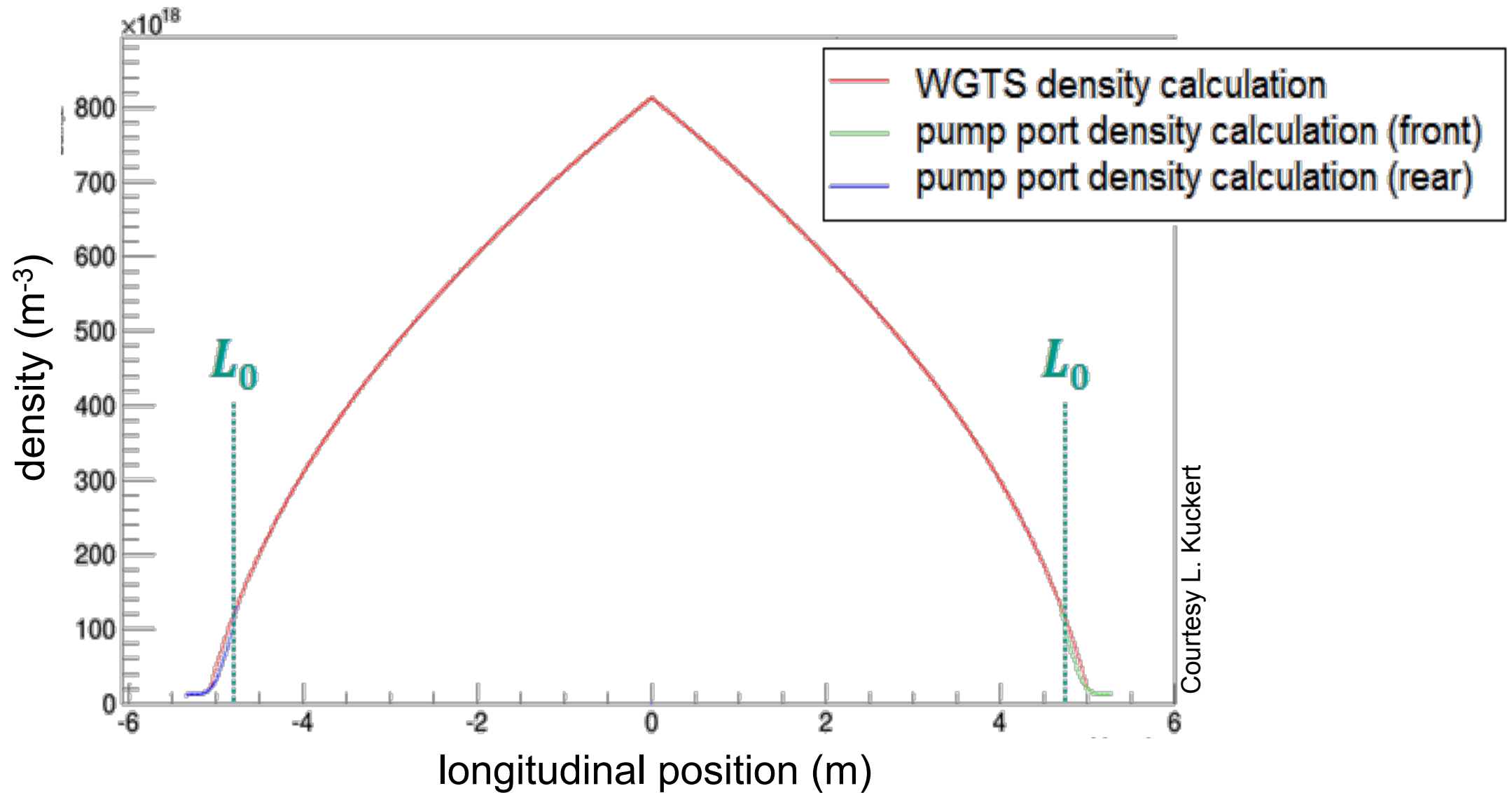


So far: only upper ( $< 2$  eV) and lower bounds ( $>0.01$  resp.  $>0.05$  eV)

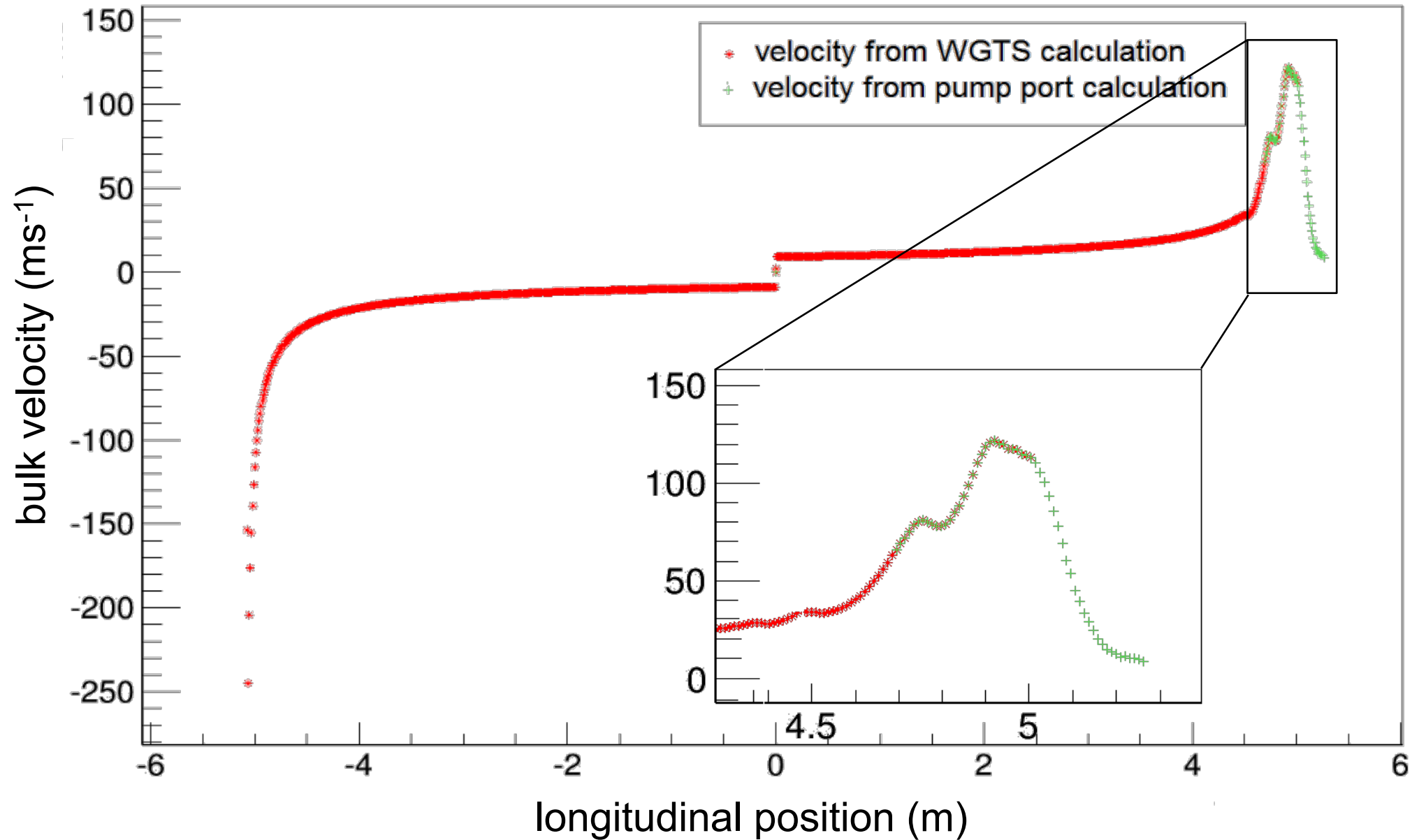
# WGTS gas flow regimes



# 1D tritium density profile

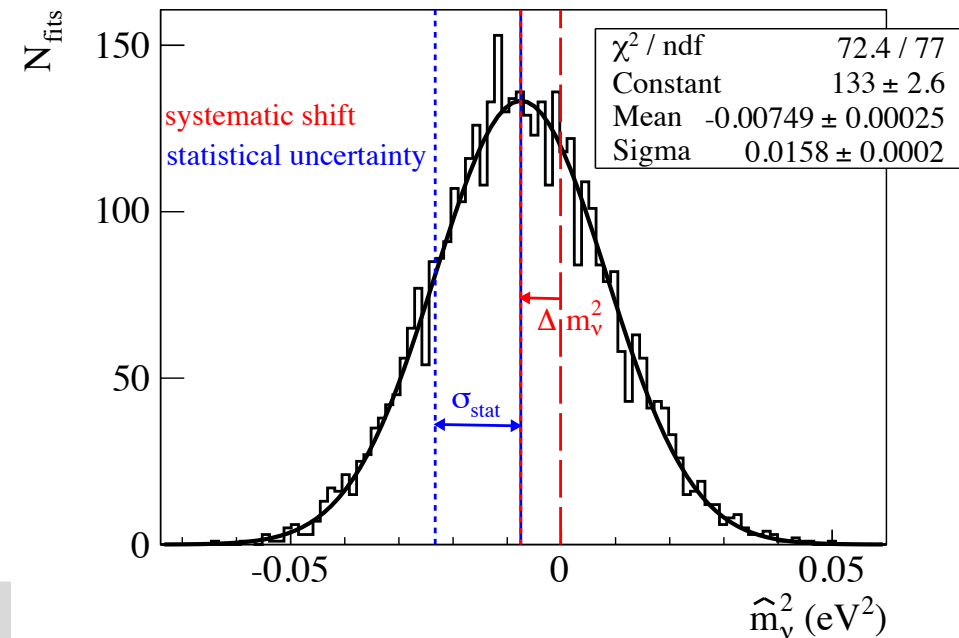
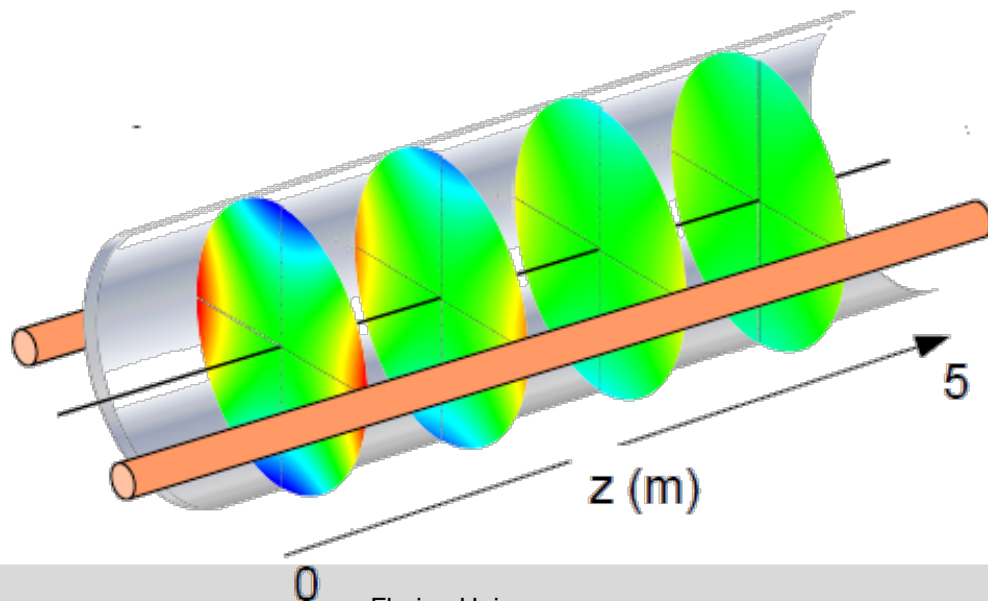
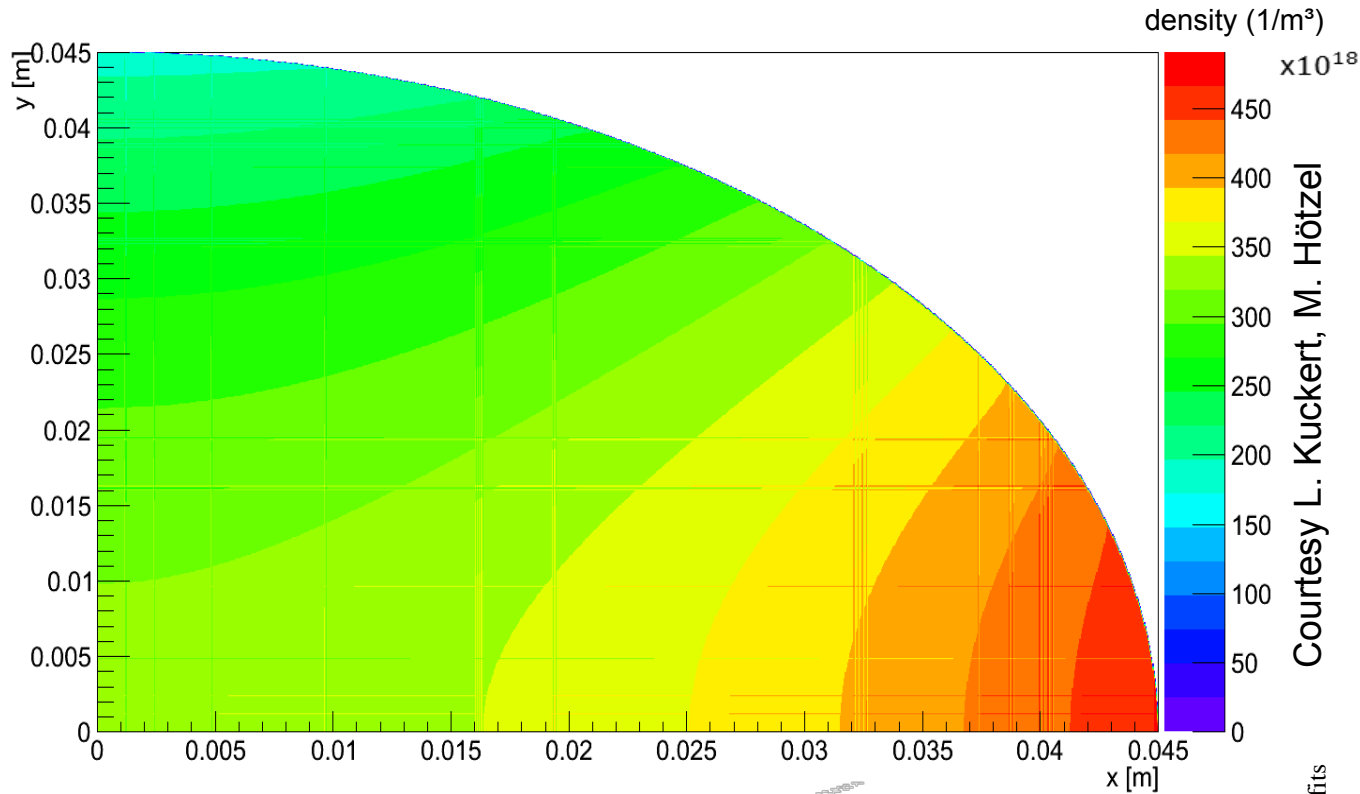


# 1D velocity profile



Courtesy L. Kuckert

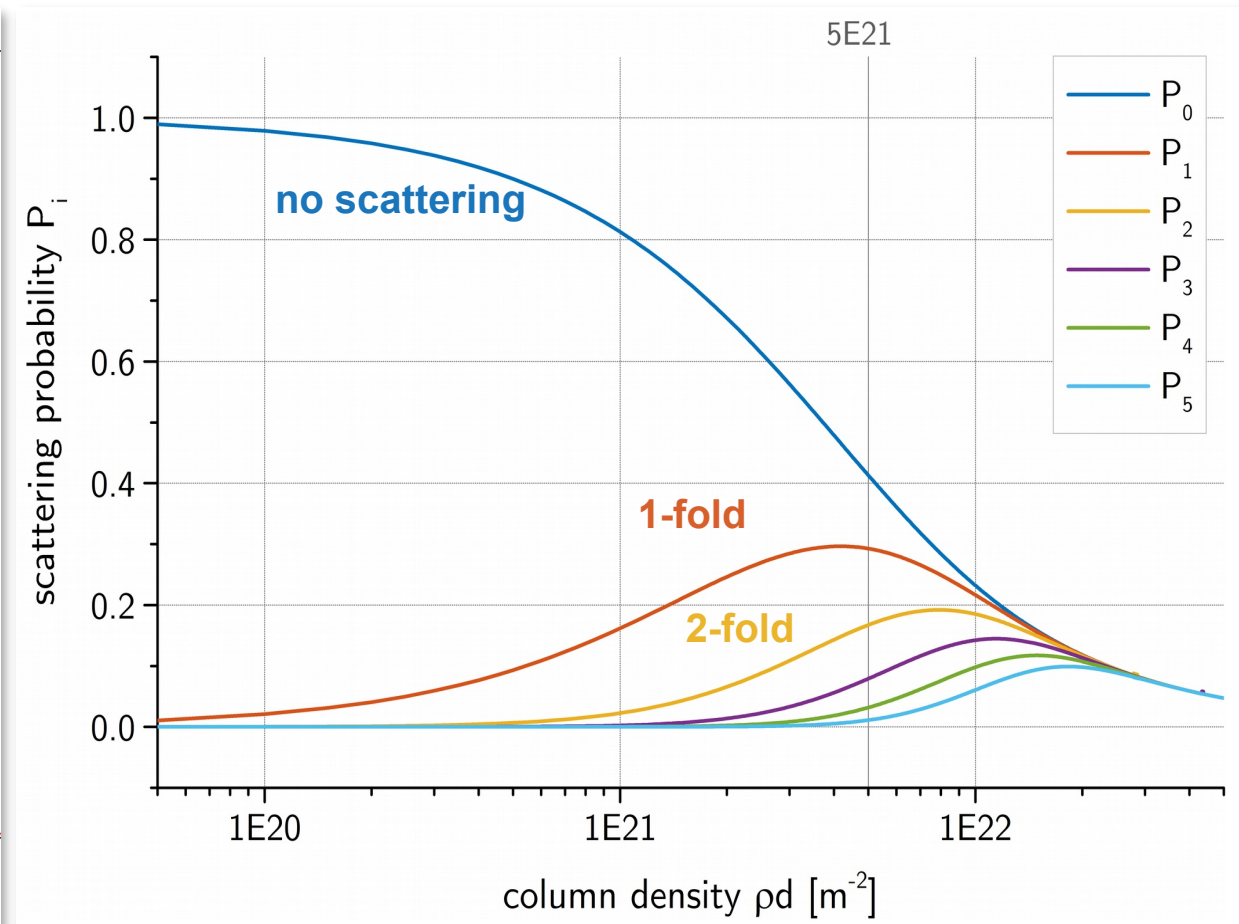
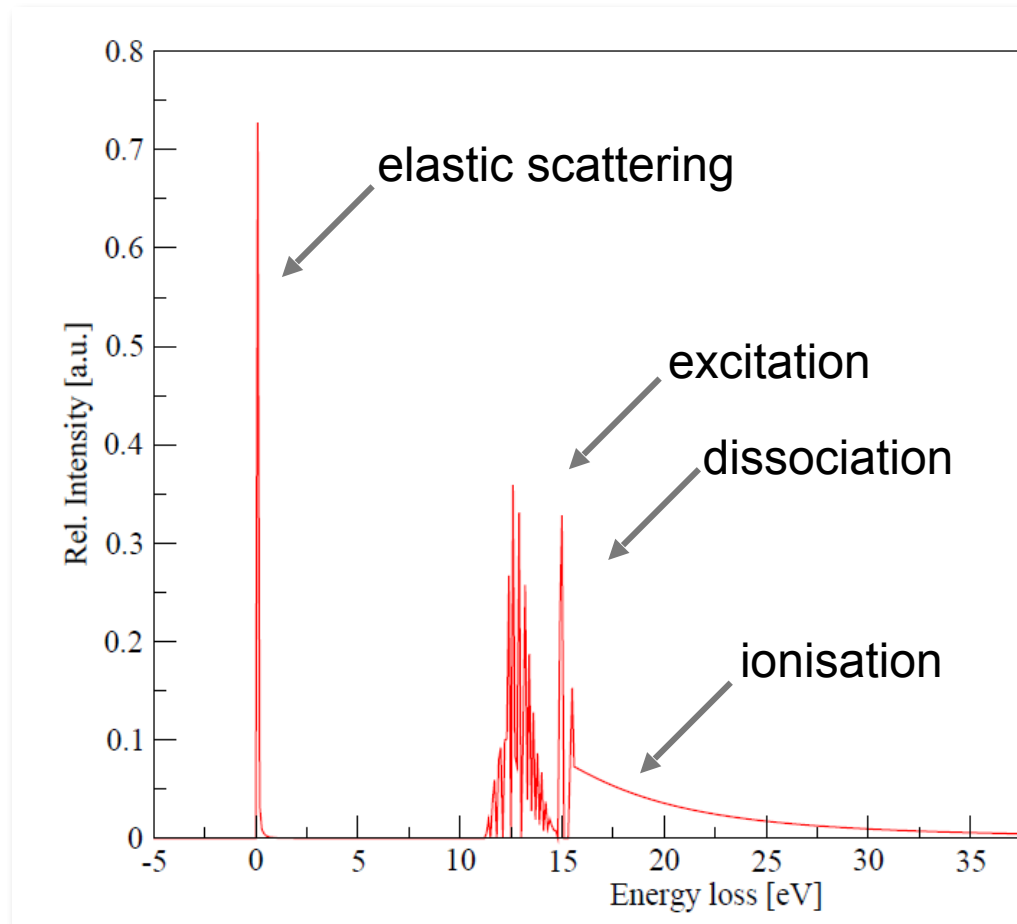
# Pseudo 3D density profile





# Example: Energy loss function

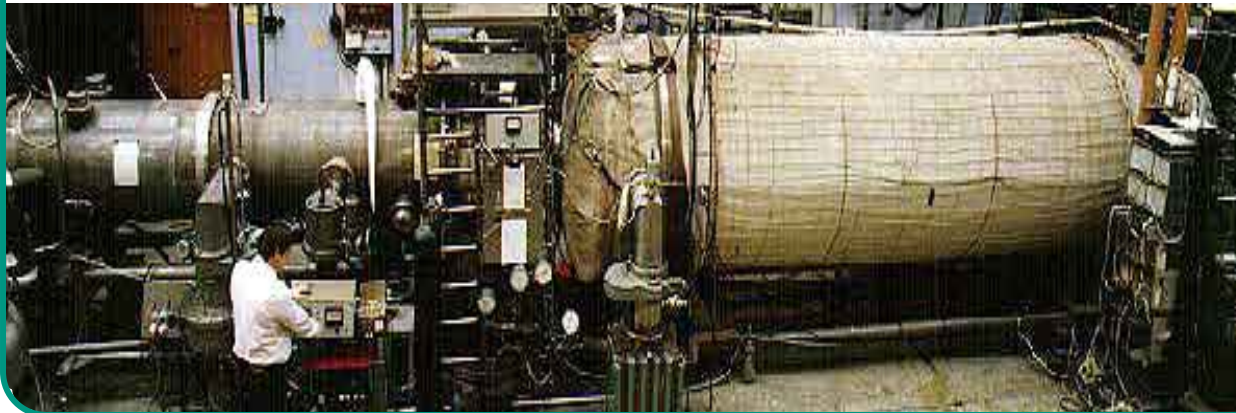
18.6 keV electrons undergo scattering & energy loss when traversing the gaseous T<sub>2</sub> source:



# Troitsk & Mainz experiments

## Troitsk experiment

- windowless gaseous tritium source



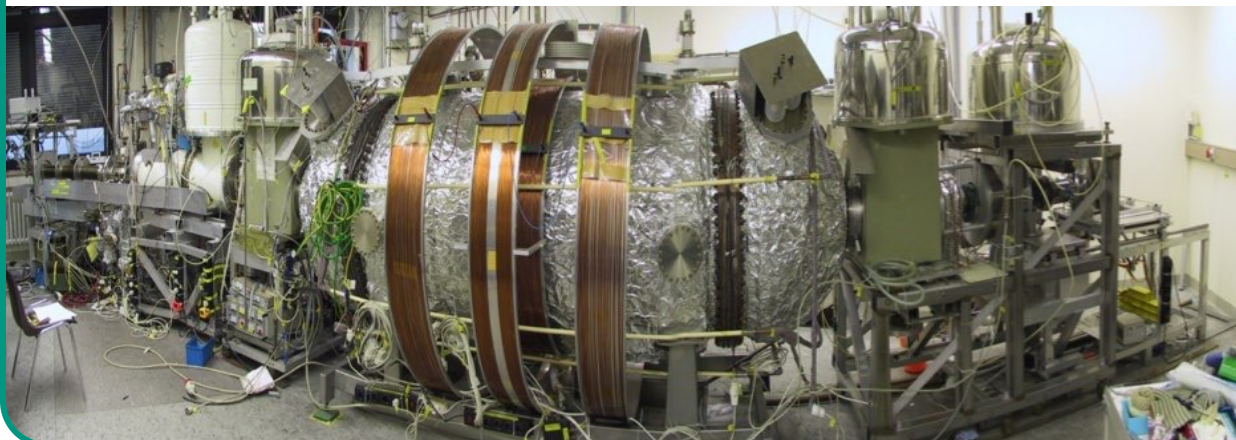
$$m^2(\nu_e) = (-0.67 \pm 1.89 \pm 1.68) eV^2$$

$$m(\nu_e) < 2.05 eV$$

V.N. Aseev et al., Phys. Rev. D 84 (2011) 112003

## Mainz experiment

- quench condensed tritium source

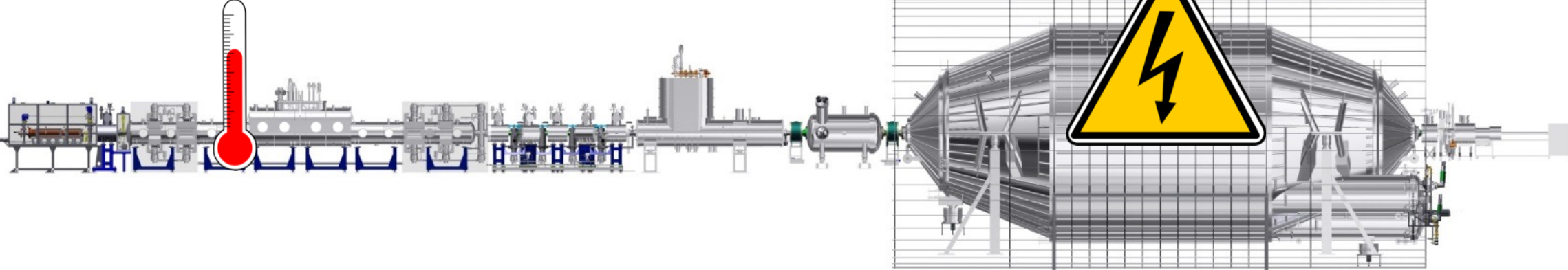


$$m^2(\nu_e) = (-0.6 \pm 2.2 \pm 2.1) eV^2$$

$$m(\nu_e) < 2.3 eV$$

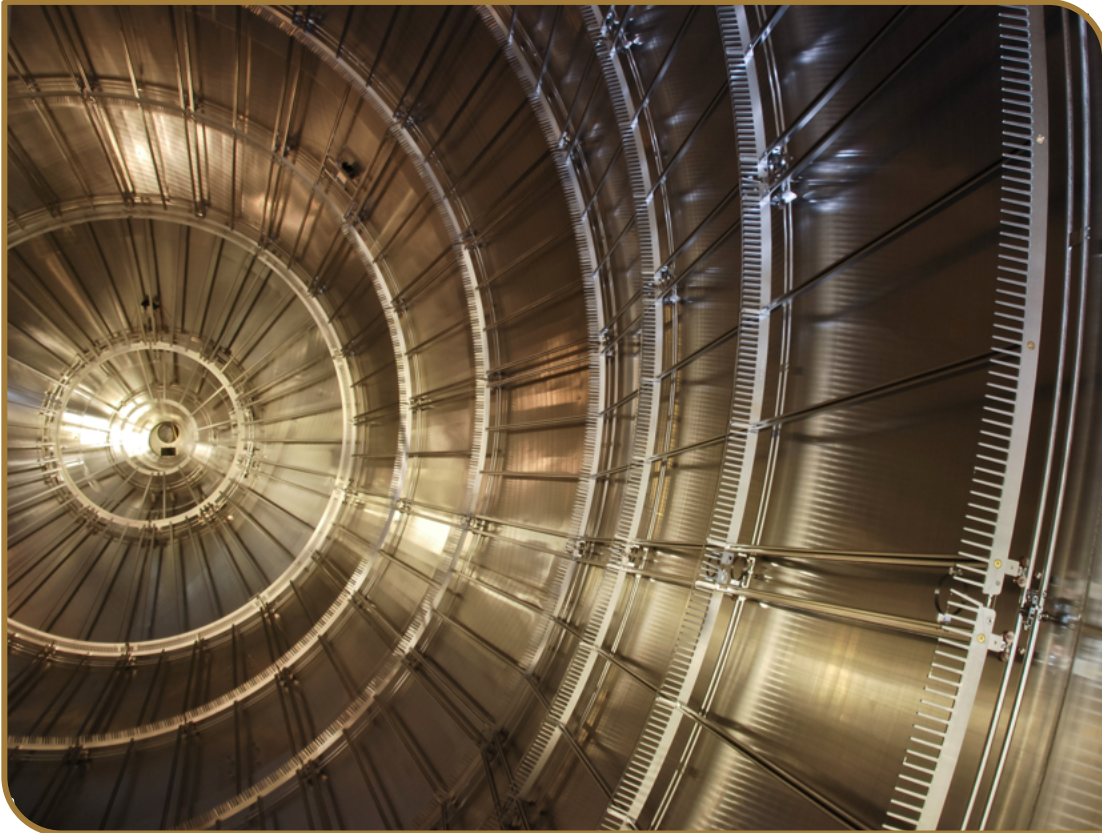
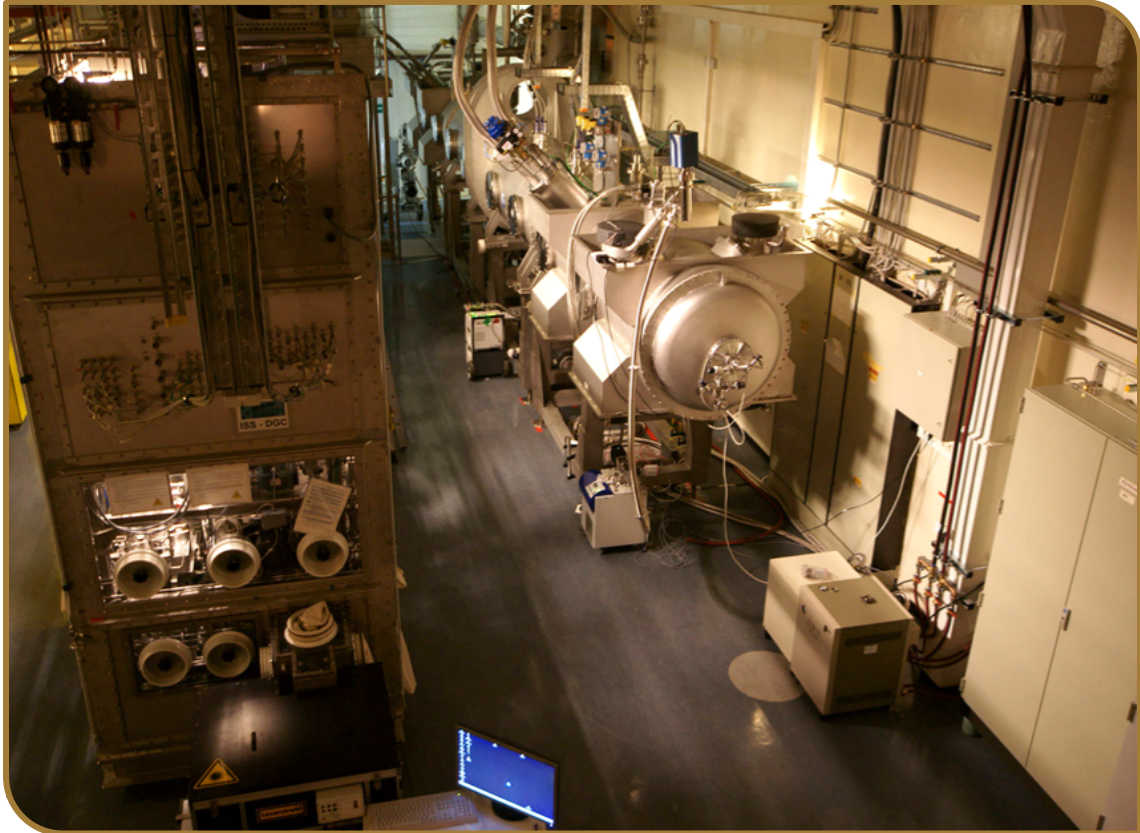
C. Kraus et al., Eur. Phys. J. C 40 (2005) 447

# Experimental challenges ... and solutions



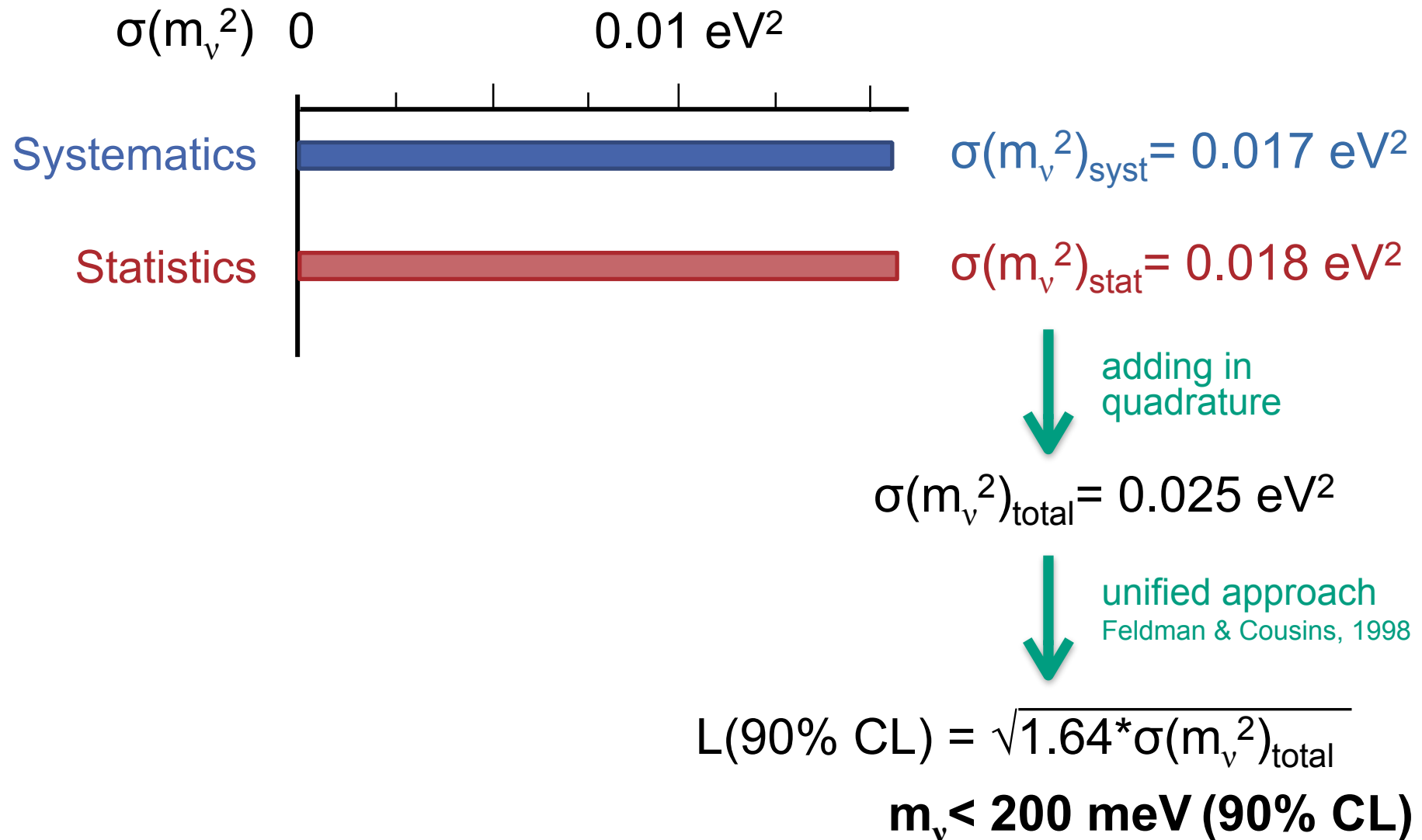
required: source fluctuation:  $\Delta T/T < 10^{-3}$

required: HV-fluctuations:  $\Delta U < 60 \text{ mV}$



# KATRIN sensitivity in a nutshell

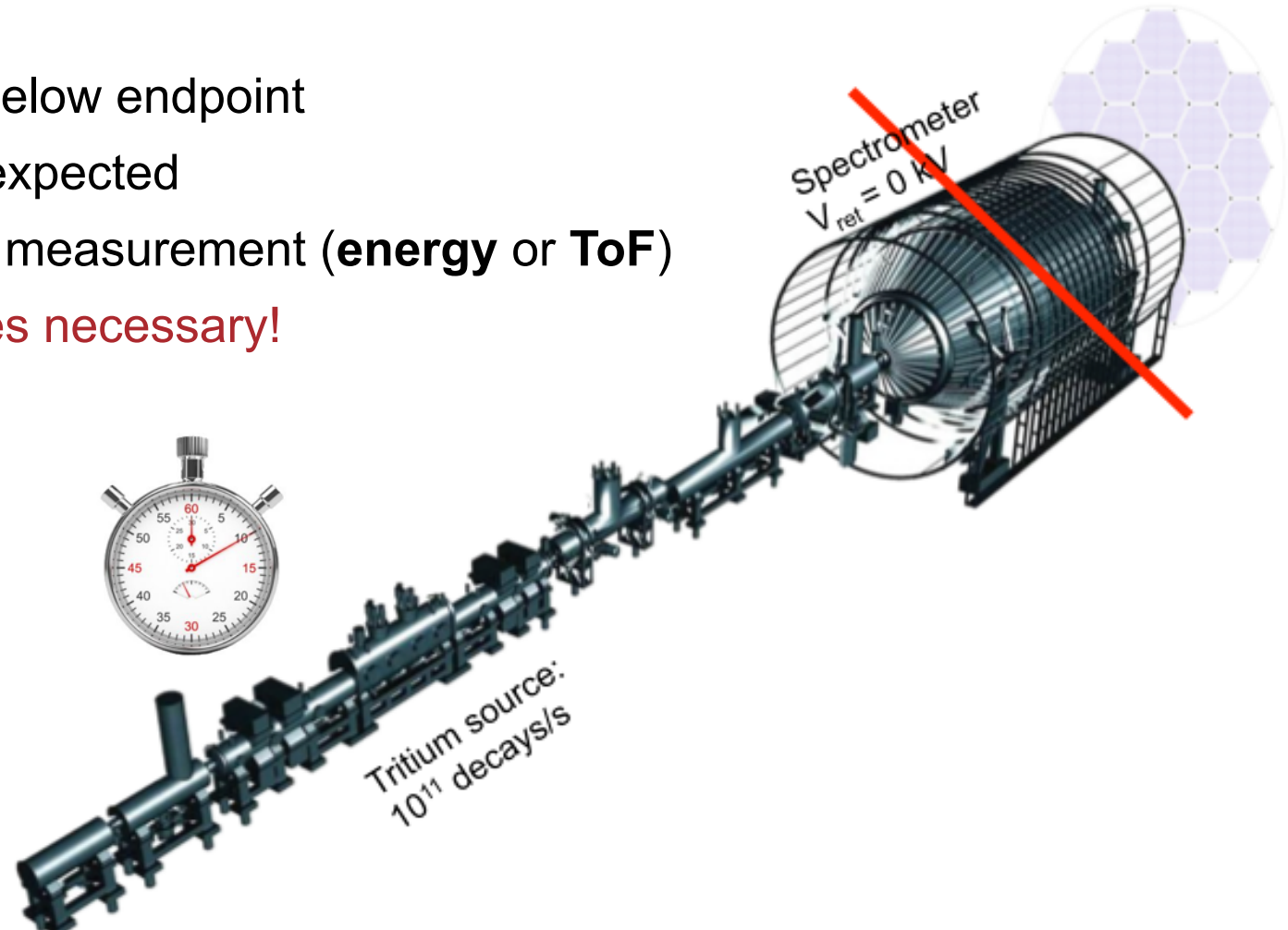
A simple sensitivity estimate from combining (conservative) **systematics** budget with **statistical** uncertainty (3 net years of data):



# Search for keV-scale sterile $\nu$ with KATRIN

## The challenge:

- High count rates at  $\sim$ few keV below endpoint
  - Tiny sterile admixture  $\sin^2(\theta_s)$  expected
  - Best sensitivity for **differential** measurement (**energy** or **ToF**)
- ➔ Development of new techniques necessary!

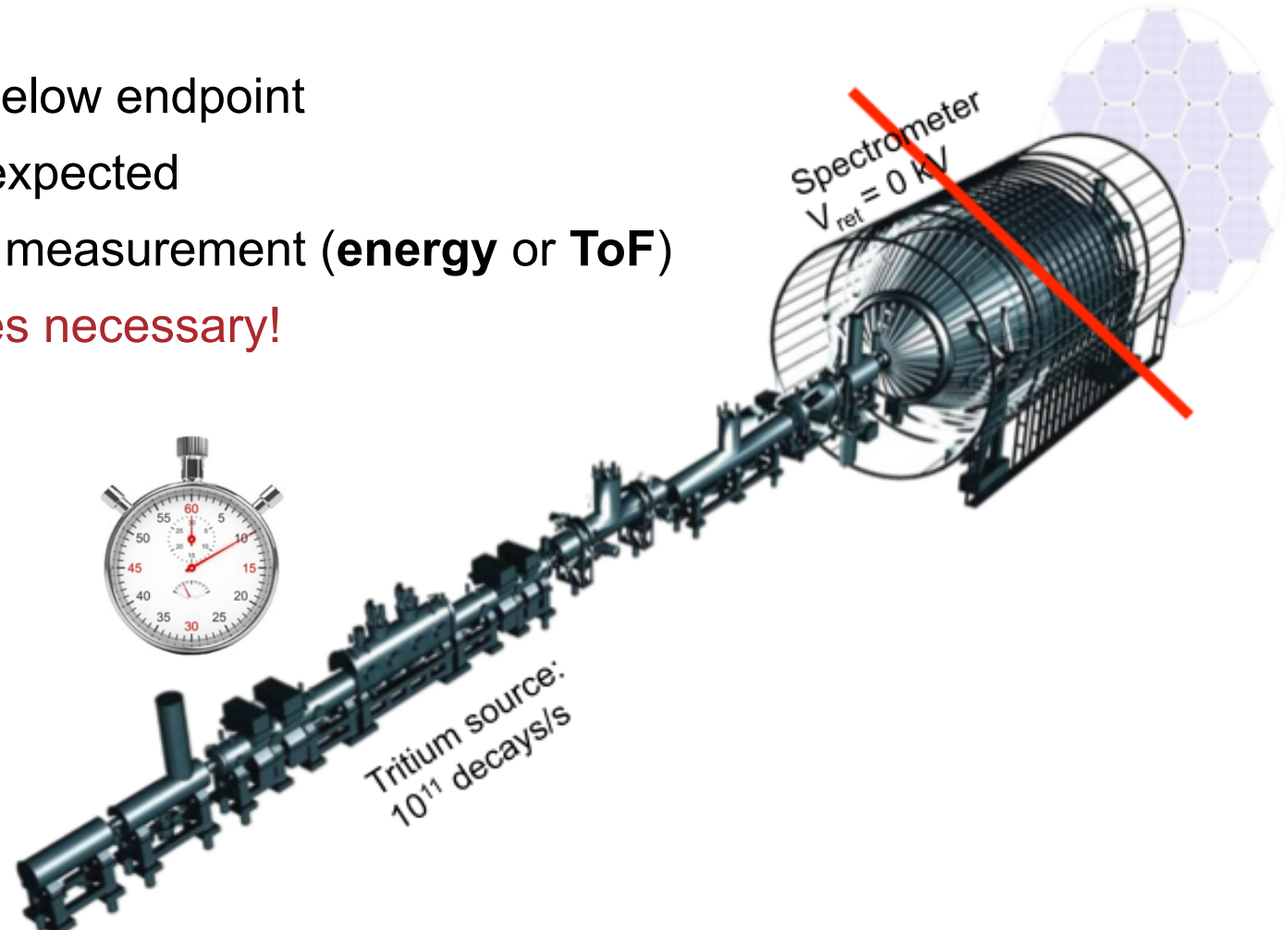
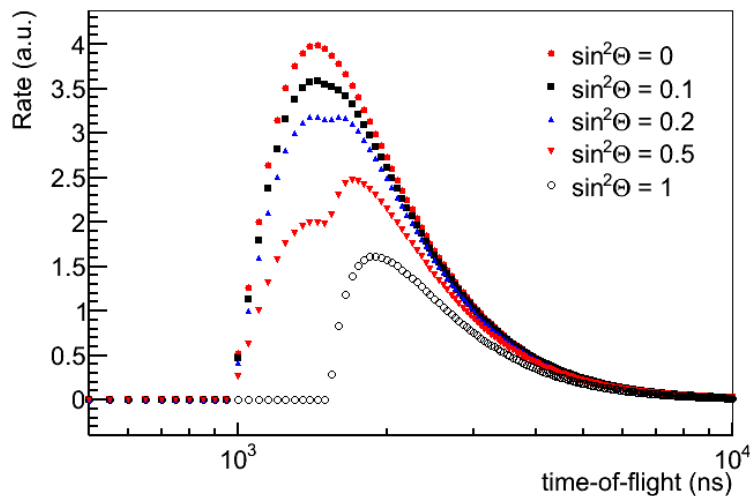


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ToF option:  
electron tagger required



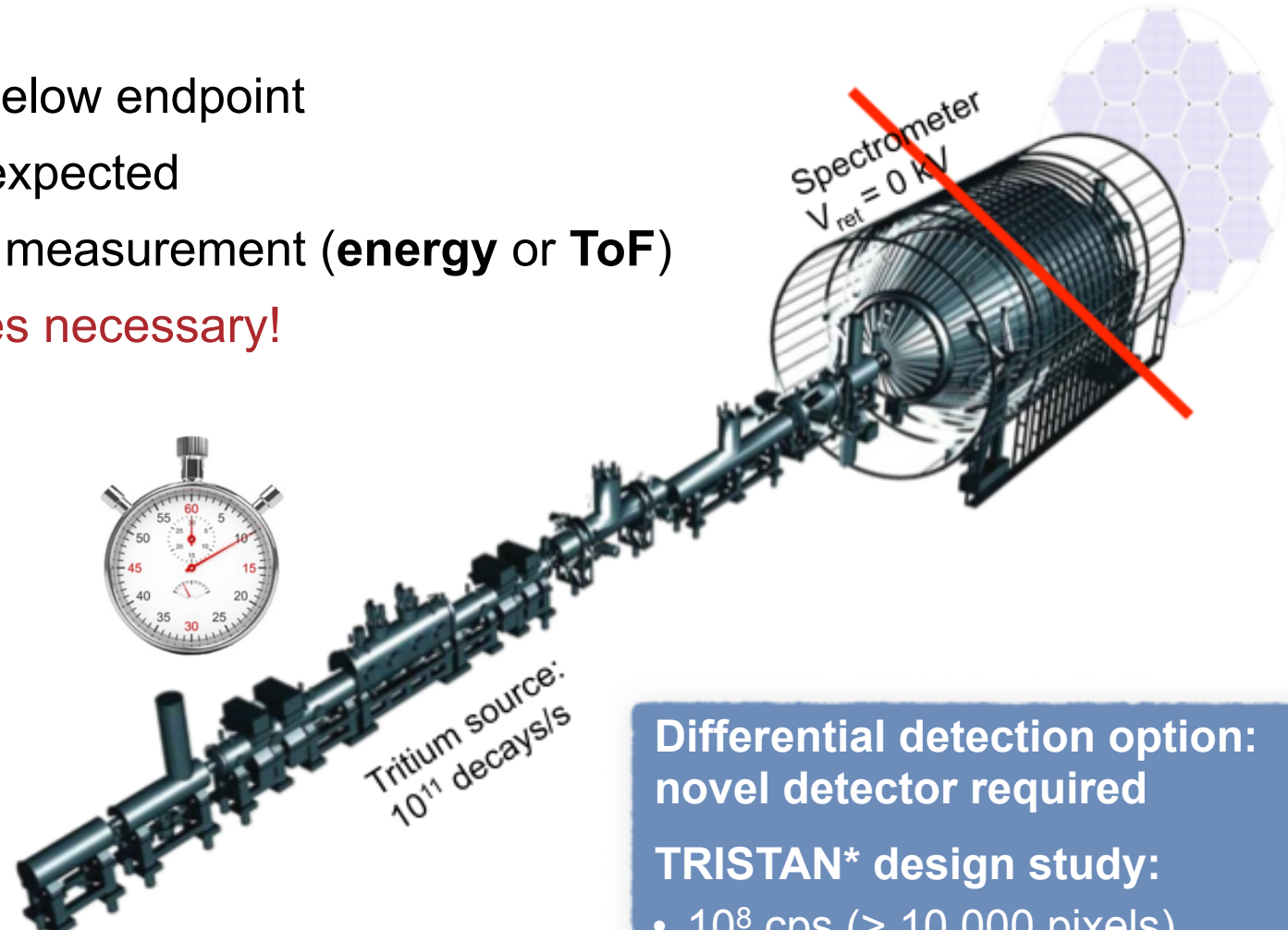
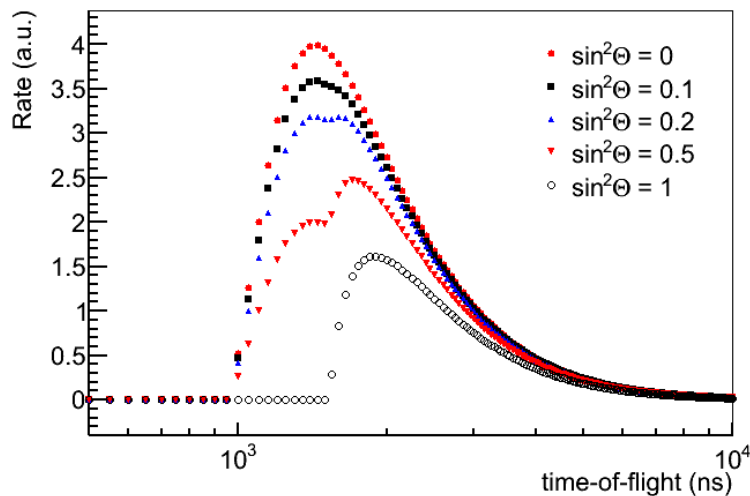
[Steinbrink et al. (2013), Robertson et al. (in prep.)]

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- ➔ Development of new techniques necessary!

## ToF option: electron tagger required



## Differential detection option: novel detector required

### TRISTAN\* design study:

- $10^8$  cps ( $> 10\,000$  pixels)
- FWHM 300 eV @ 20 keV
- $> 20$  cm diameter

[Steinbrink et al. (2013), Robertson et al. (in prep.)]

[Mertens et al. (2015)]

# Search for keV-scale sterile $\nu$ with KATRIN



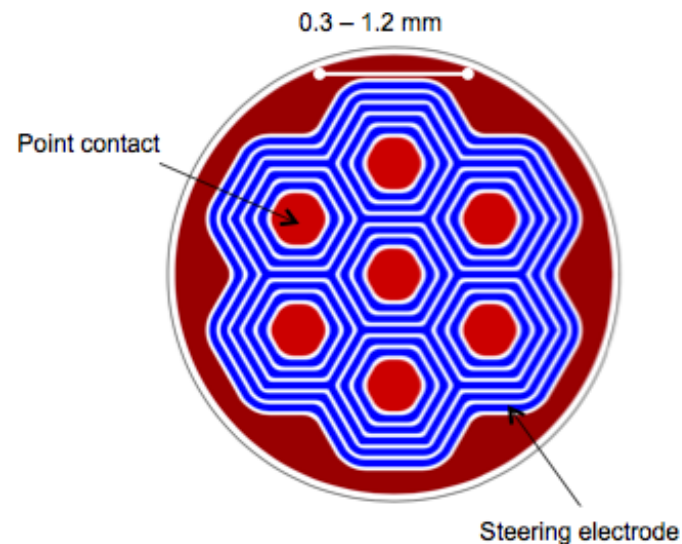
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- Prototyping and sensitivity studies for upgraded detector system under way

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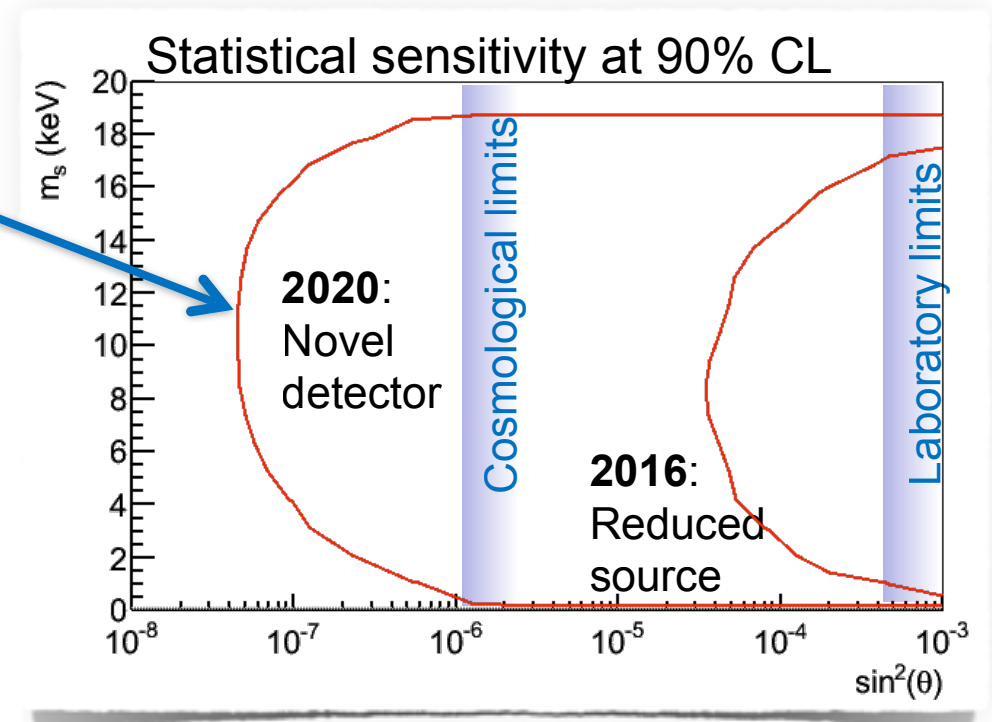
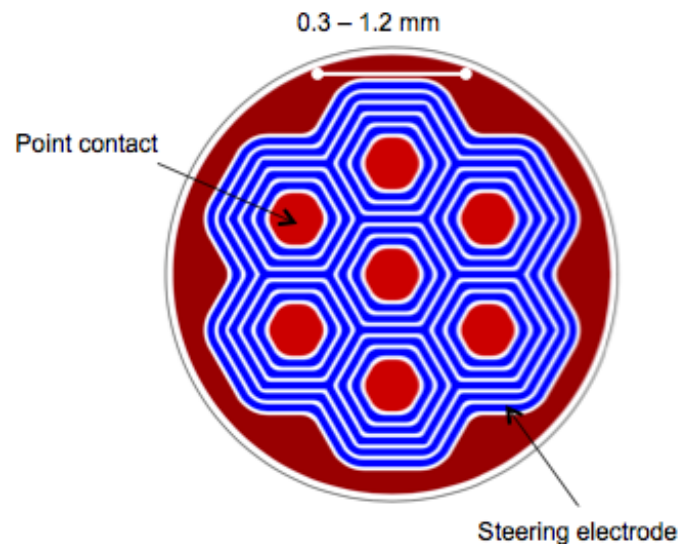
**TRISTAN prototype (10/2015):**  
characterize pile-up, backscattering,  
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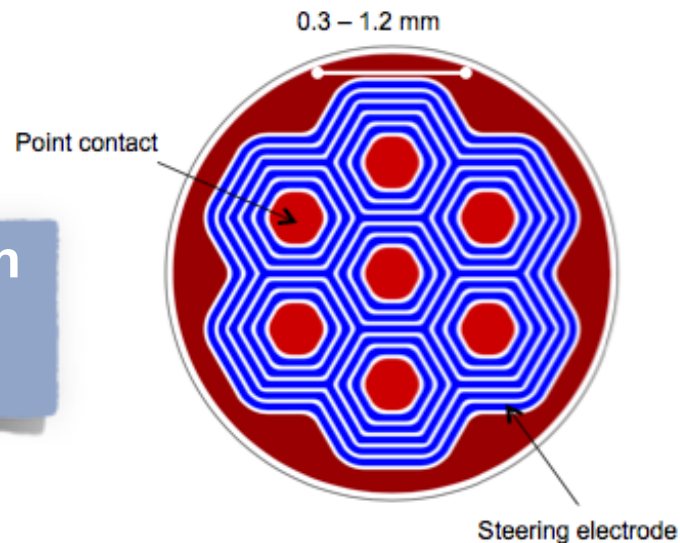


courtesy S. Mertens

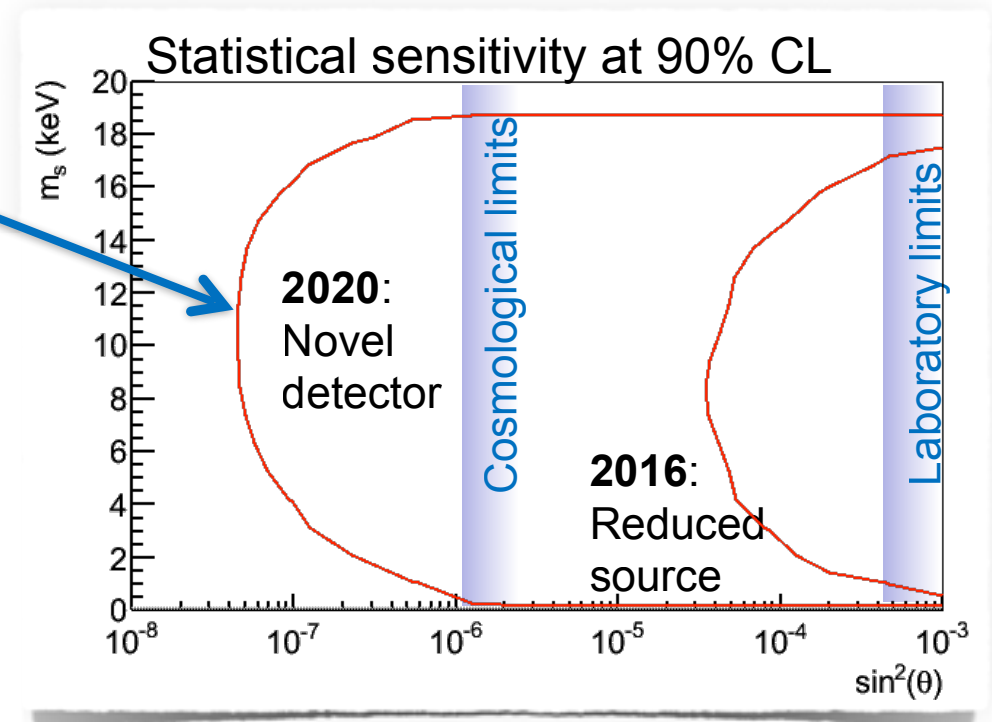
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Collaboration with  
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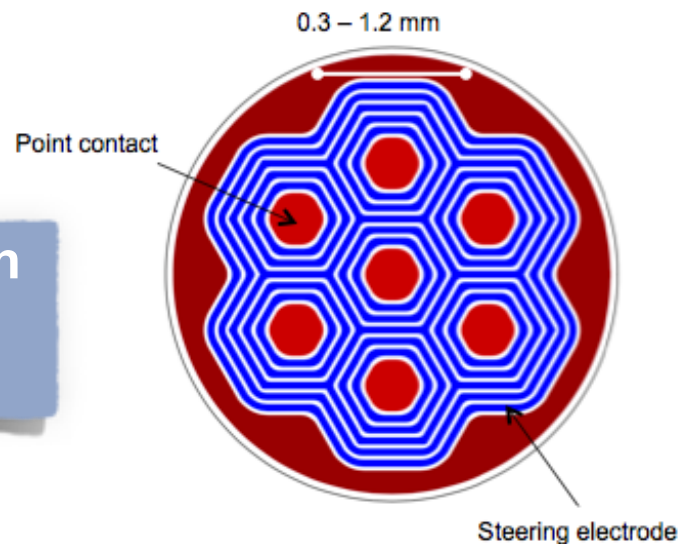


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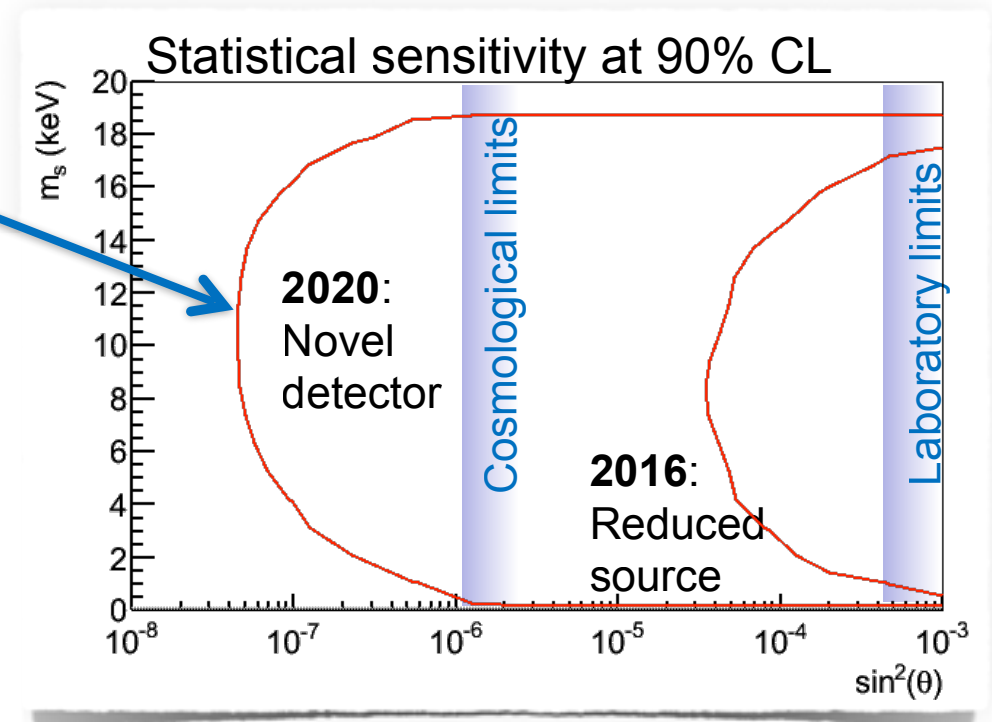
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courtesy S. Mertens

→ High-sensitivity keV sterile  $\nu$  search probing cosmologically allowed parameter space after the  $\nu$ -mass measurement with KATRIN

# Rear Section – design and assembly

Major importance for systematics:

- **Precision  $e^-$  source:**  
column density monitoring  
and determination of energy loss  
function (scattering)



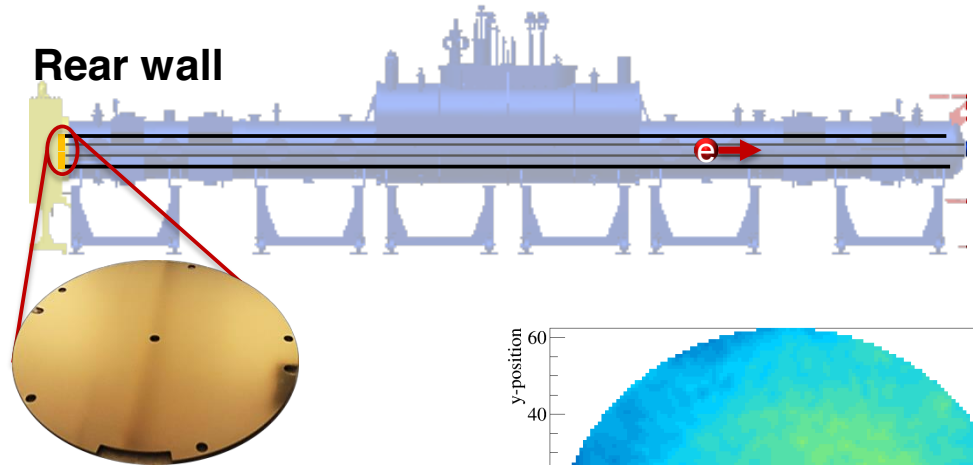
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- **Rear Wall:**  
stable and homogeneous  
electrostatic potential in the  
source plasma

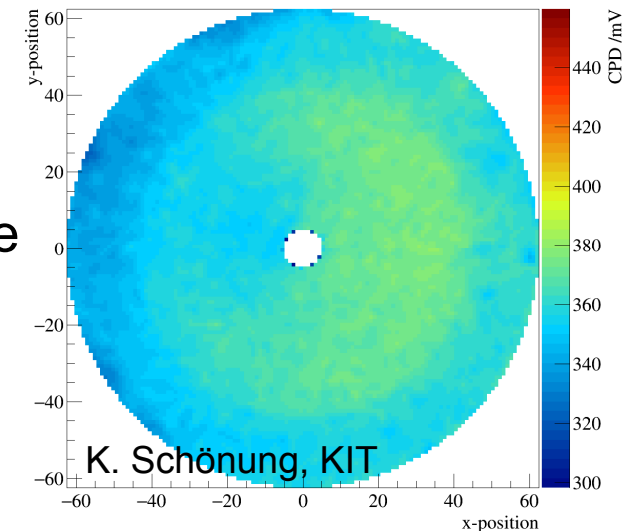


Rear wall



Ø15 cm gold surface  
for homogeneity

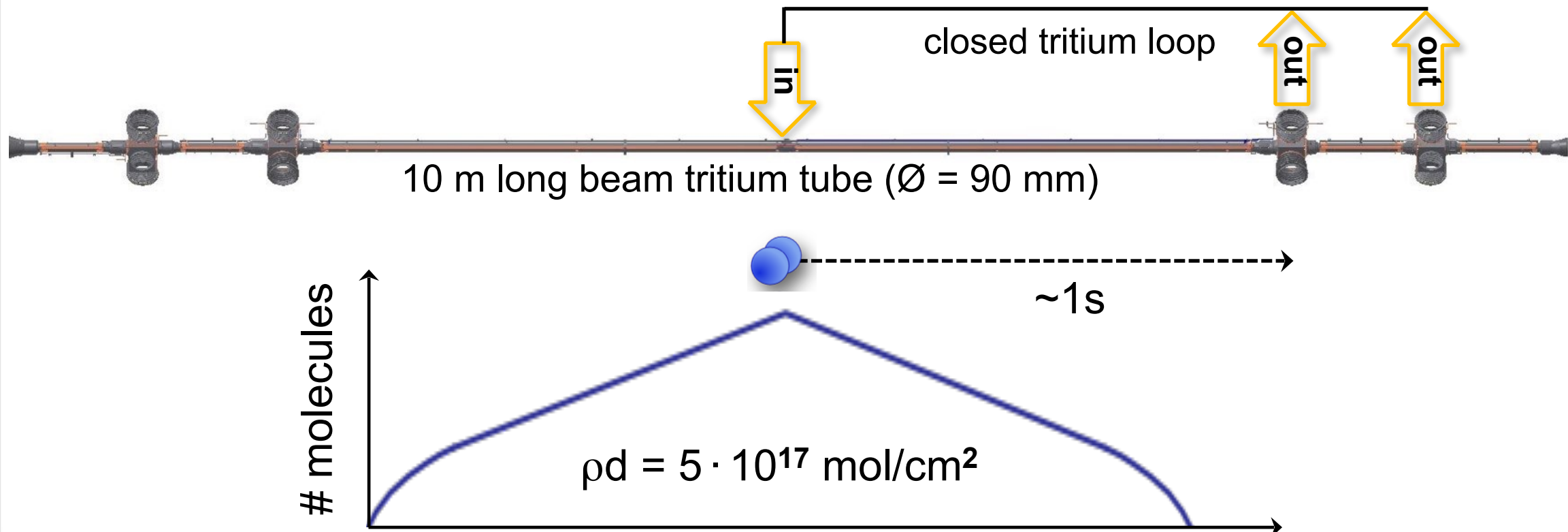
work function test  
of smaller sample



# WGTS – windowless gaseous source

## Closed-loop processing of molecular $T_2$ :

- isotopic purity > 90%
- $10^{11}$   $\beta$  decays / s
- 40% no-loss electrons
- stability at level  $10^{-3}$
- extensive control of systematics



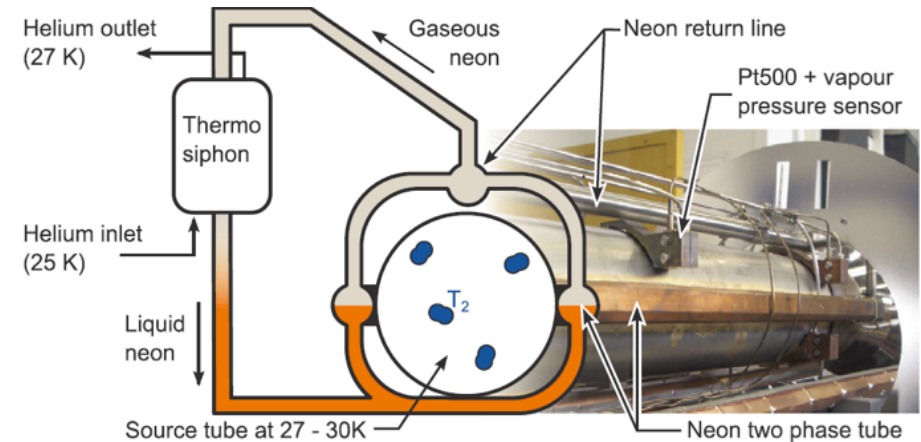


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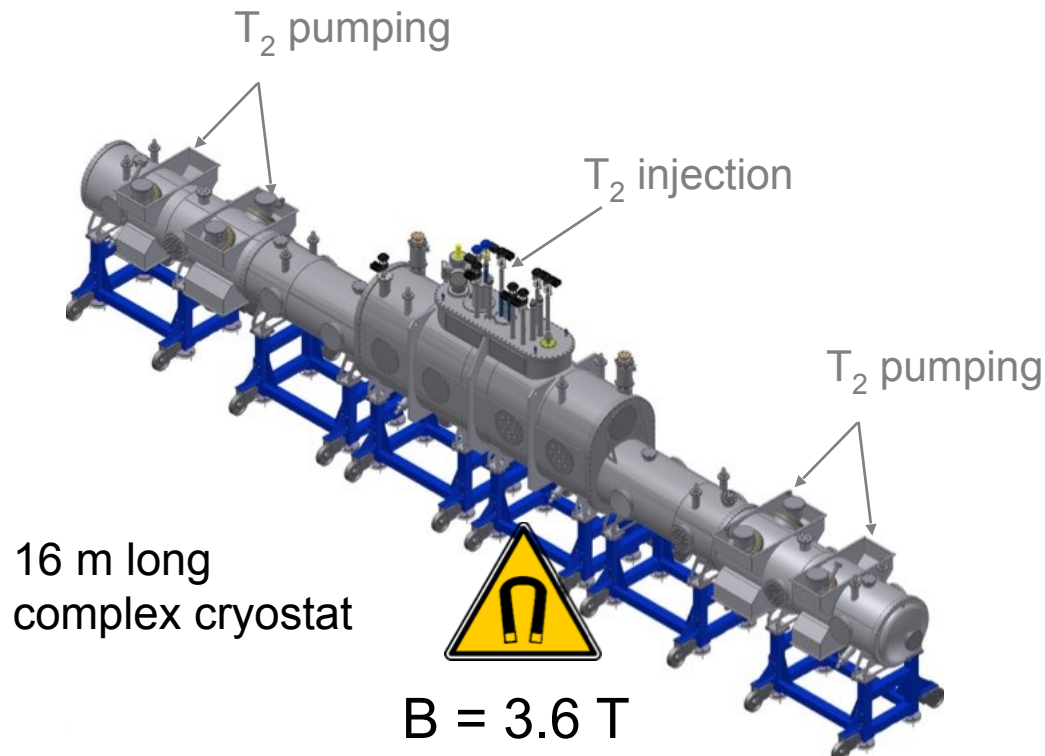
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## novel 2-phase neon cooling concept



[S. Grohmann et al., Cryogenics 55–56 (2013) 5]

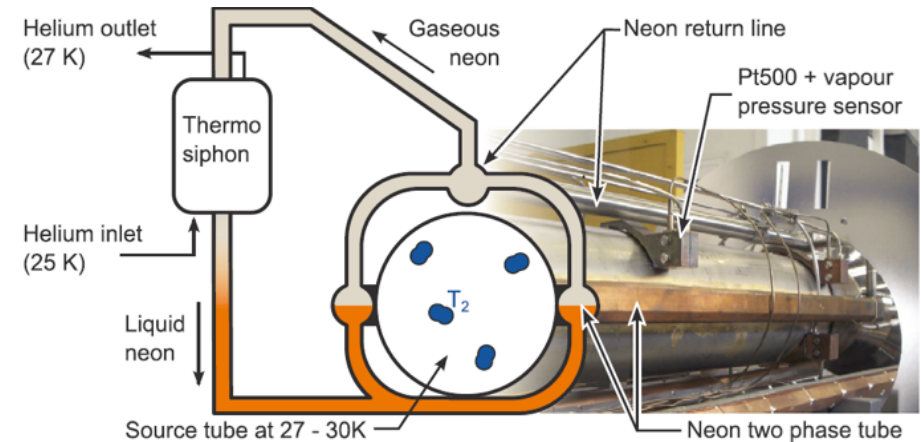


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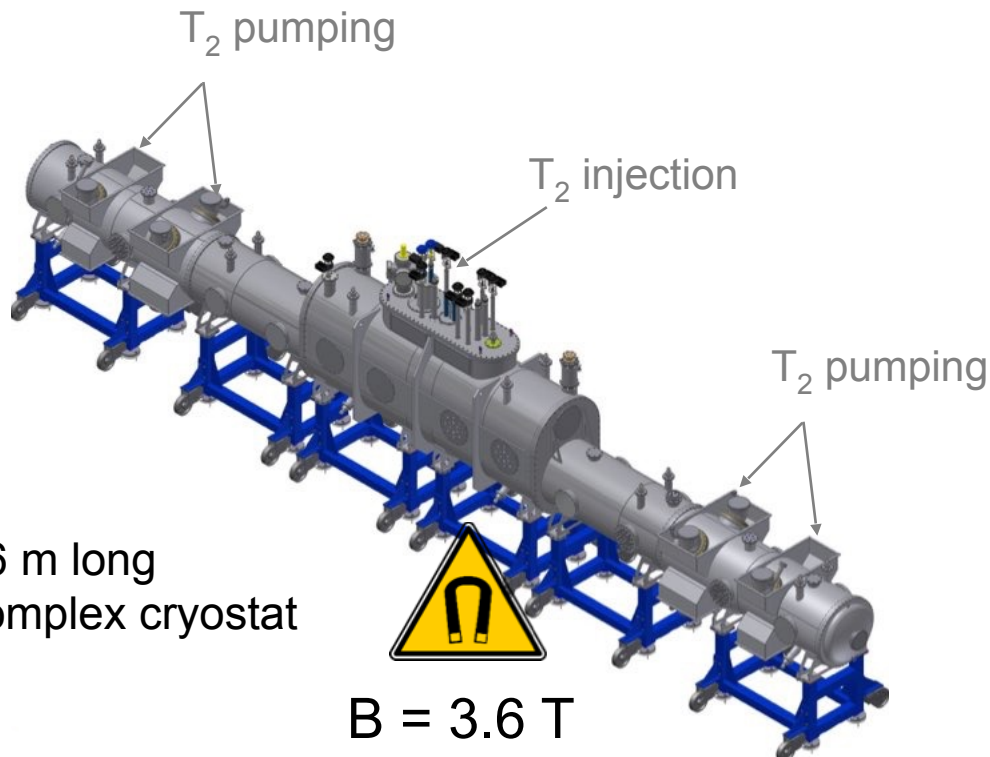
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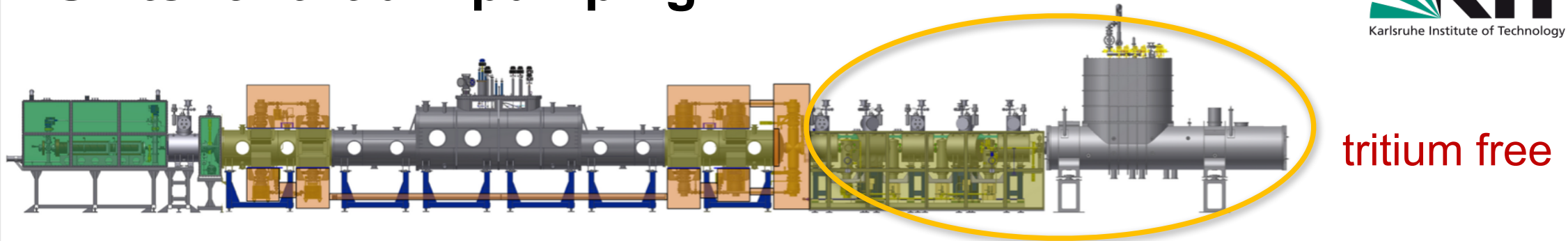
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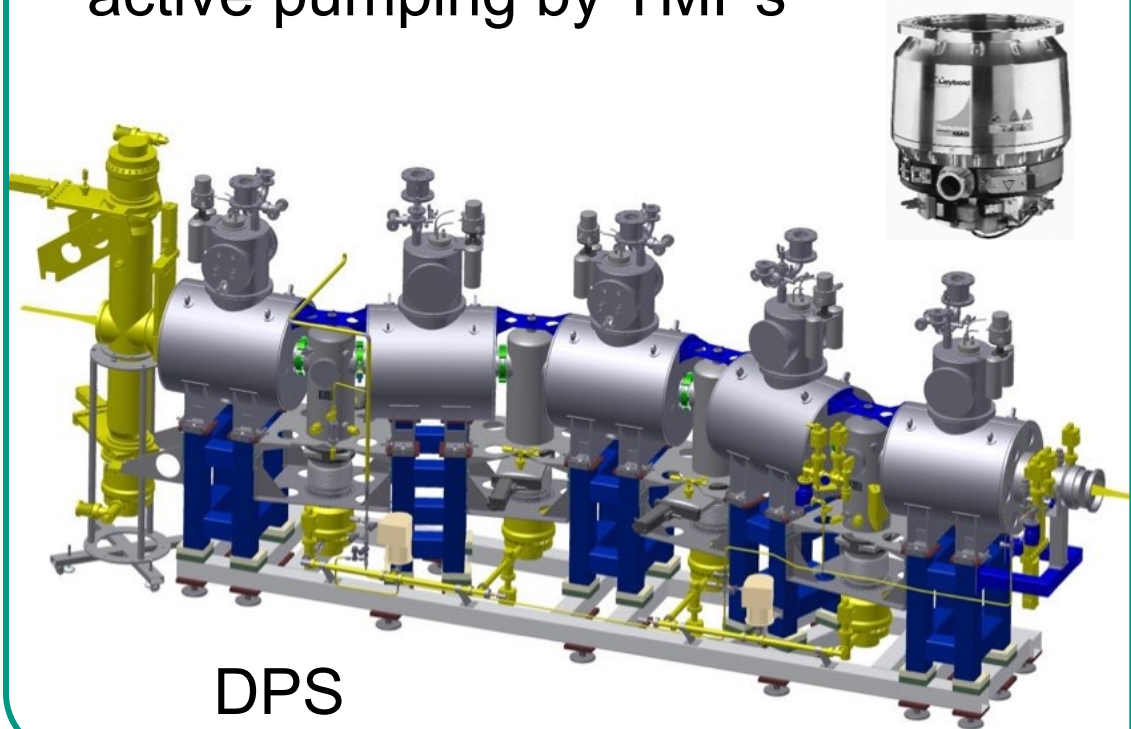


# Units for tritium pumping

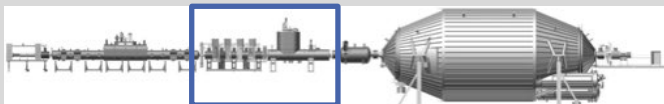
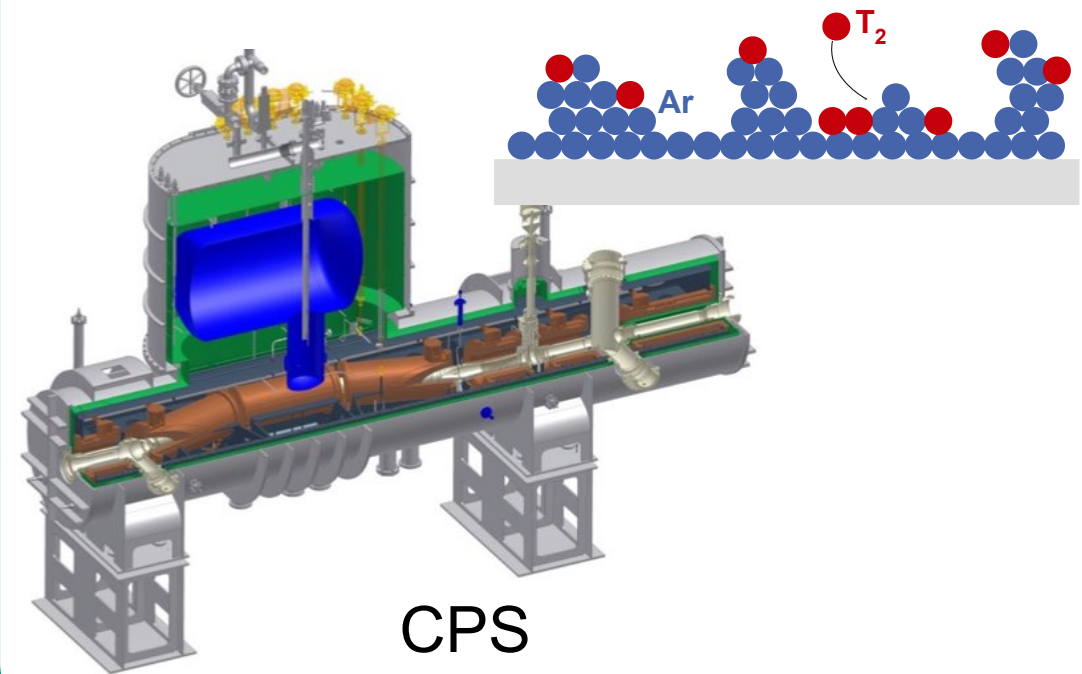


Two large cryostat systems for overall tritium retention **factor**  $> 10^{14}$

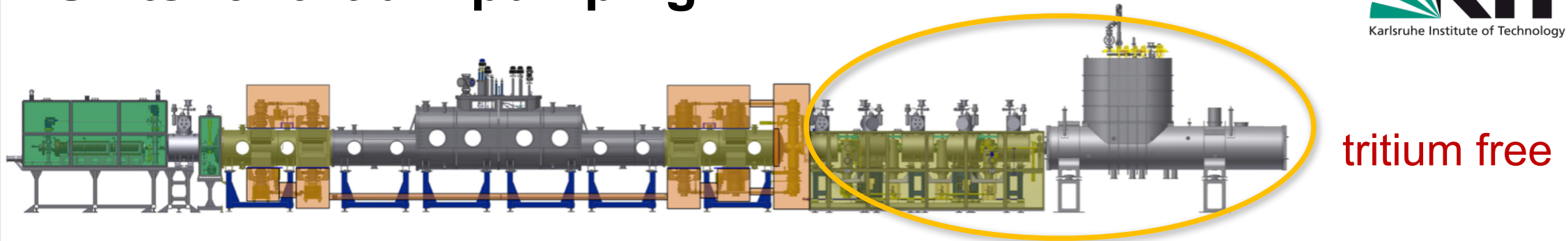
- **Differential Pumping Section DPS**  
active pumping by TMPs



- **Cryogenic Pumping Section CPS**  
cryosorption on Ar-frost at 3-4 K



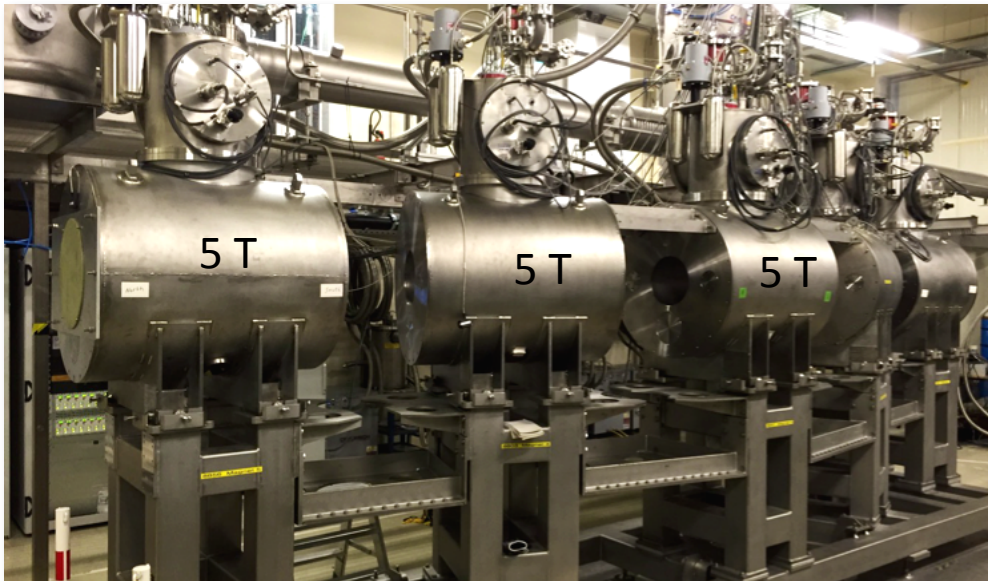
# Units for tritium pumping



tritium free

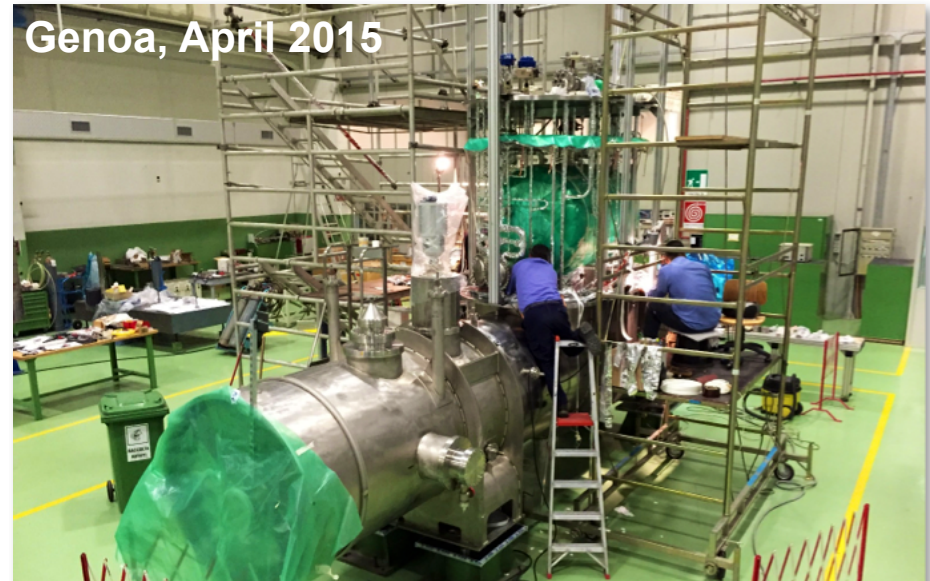
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## ■ Differential Pumping Section DPS



DPS site acceptance tests at KIT almost completed

## ■ Cryogenic Pumping Section CPS



Delivered to KIT 07/2015  
Installation started



**LFCS**

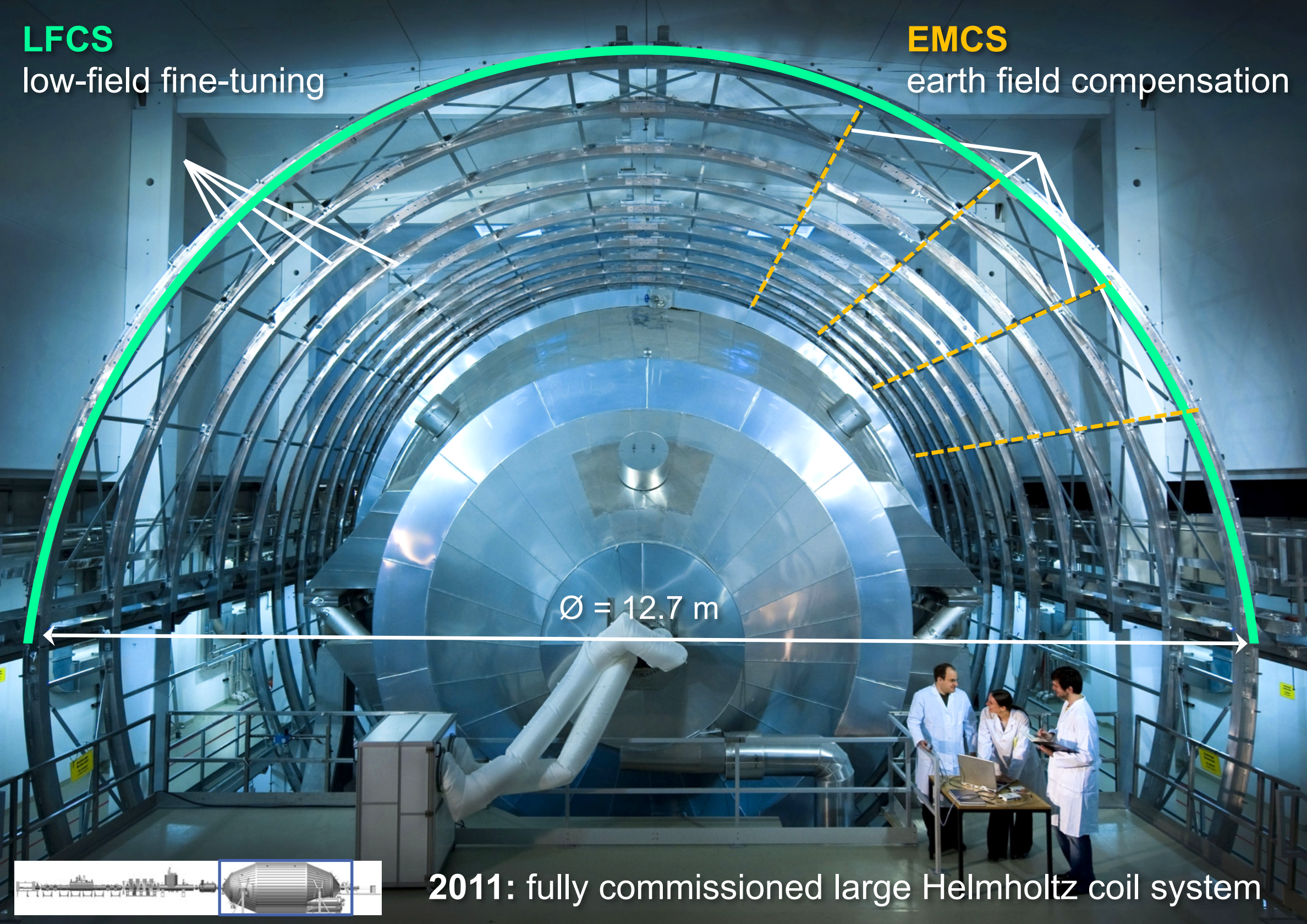
low-field fine-tuning

**EMCS**

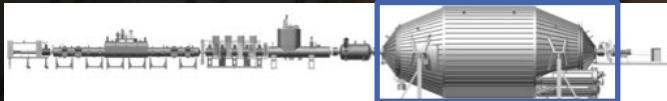
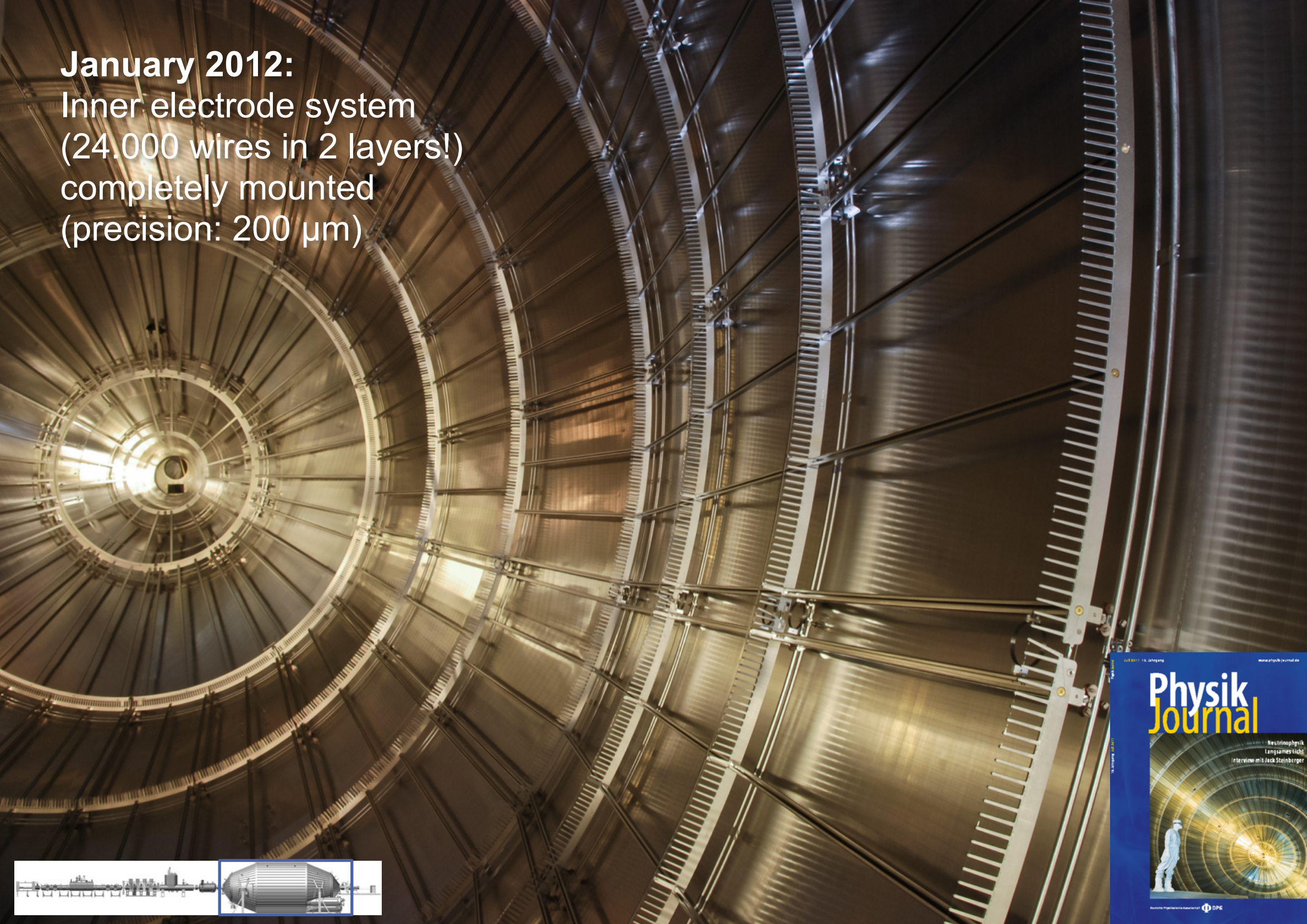
earth field compensation

$\text{Ø} = 12.7 \text{ m}$

2011: fully commissioned large Helmholtz coil system



**January 2012:**  
Inner electrode system  
(24.000 wires in 2 layers!)  
completely mounted  
(precision: 200  $\mu\text{m}$ )



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 $p < 10^{-11}$  mbar**



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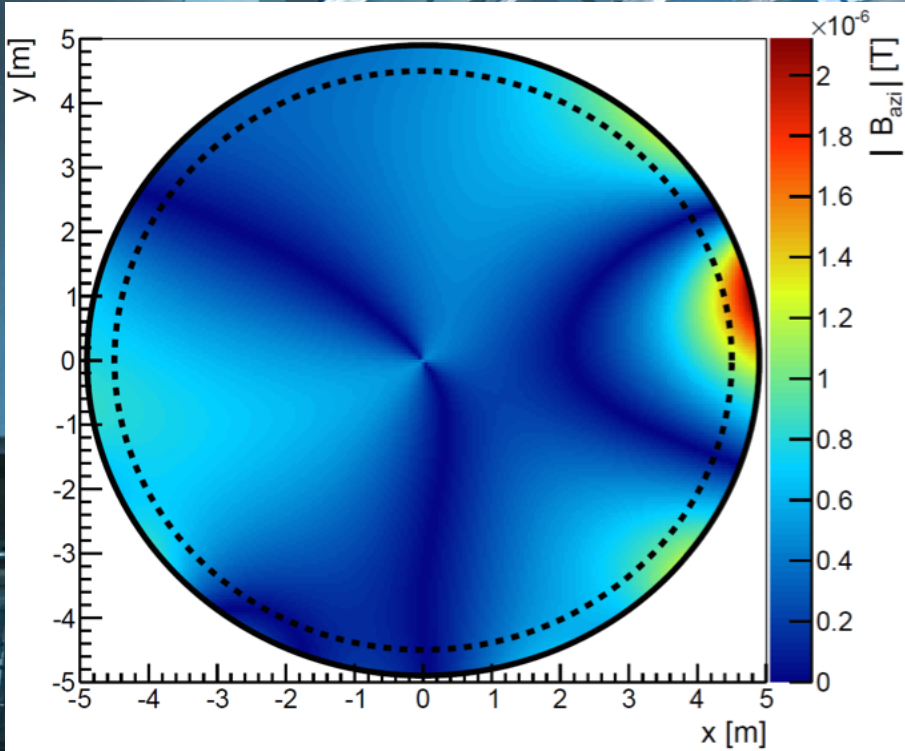
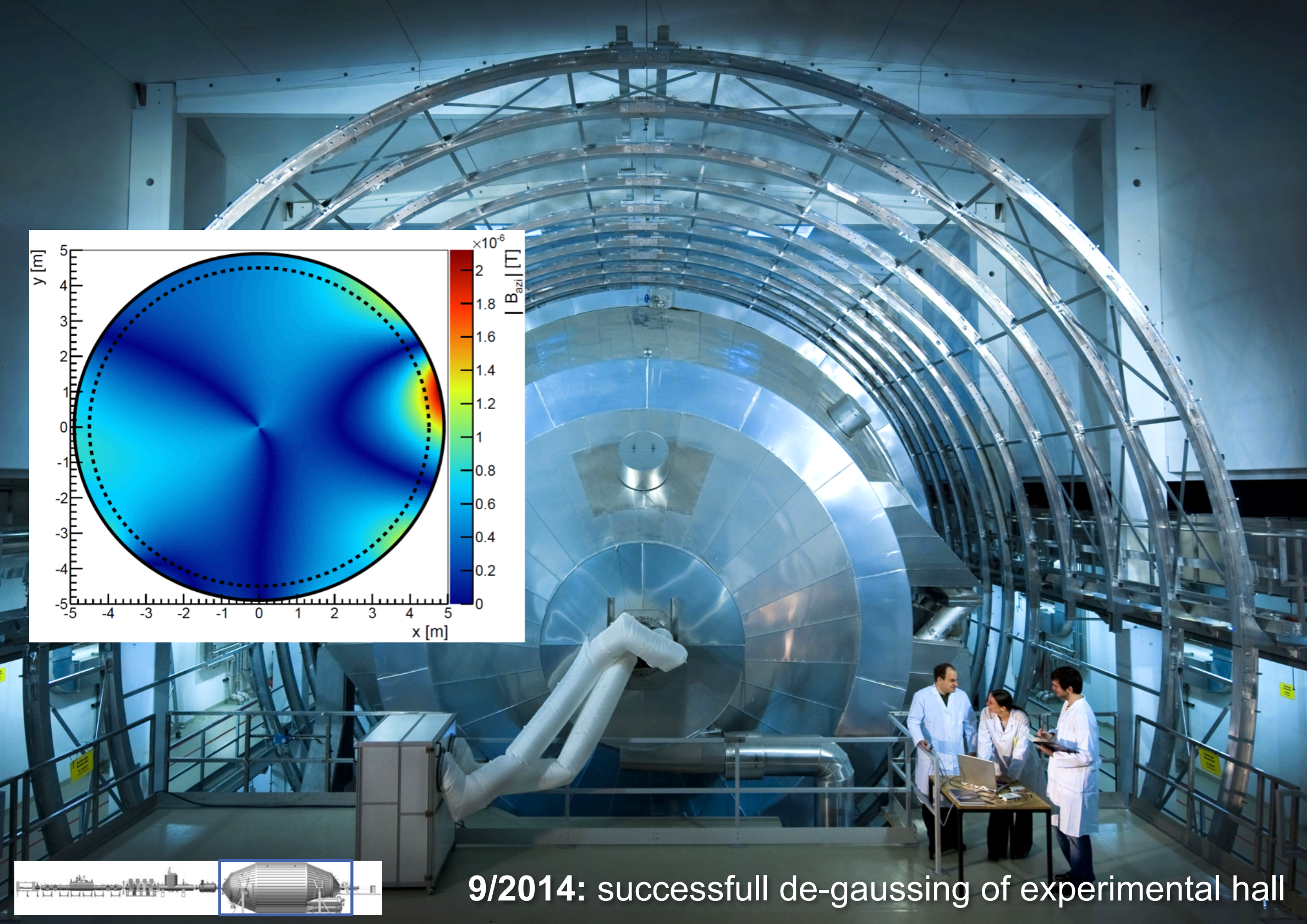
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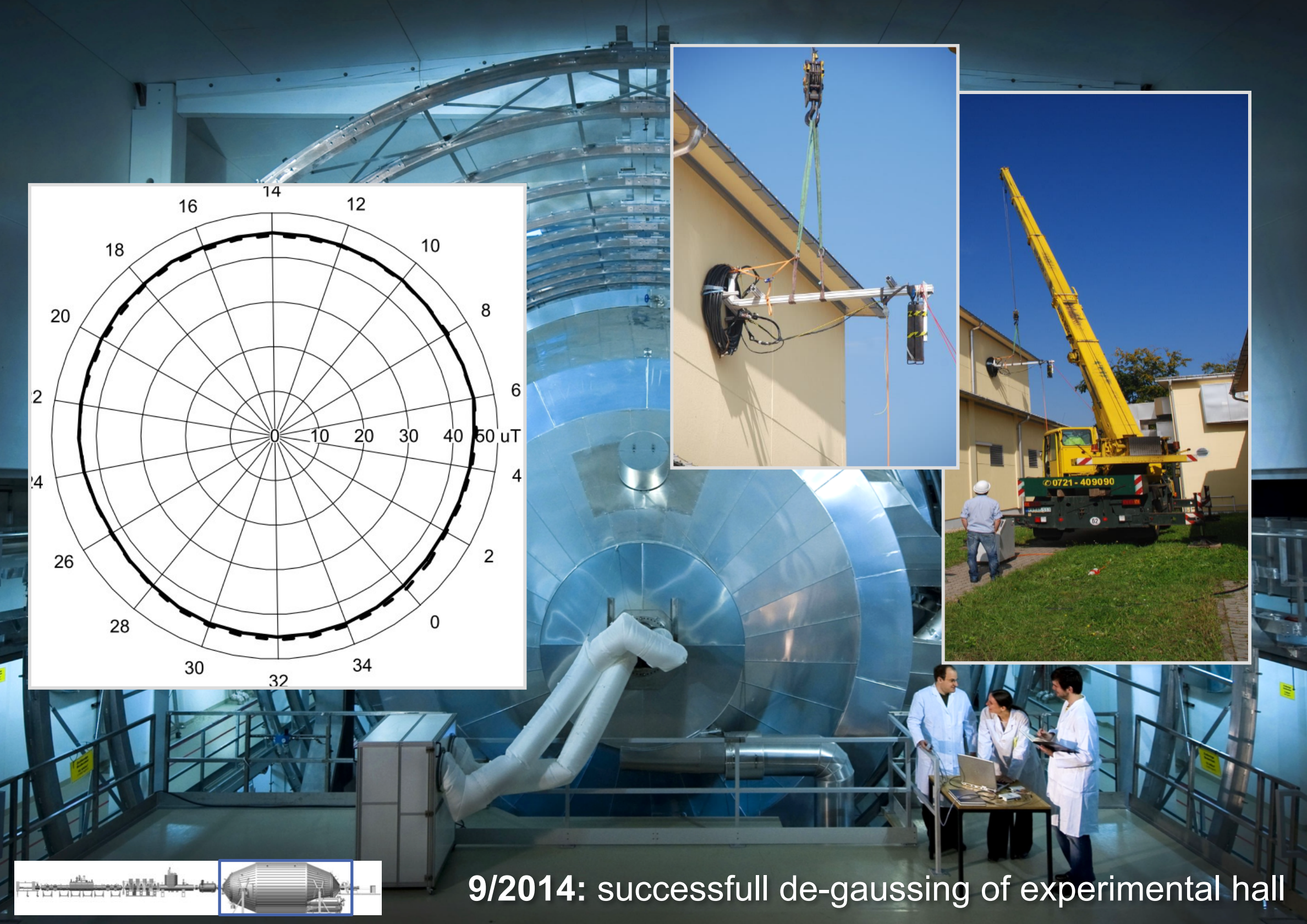
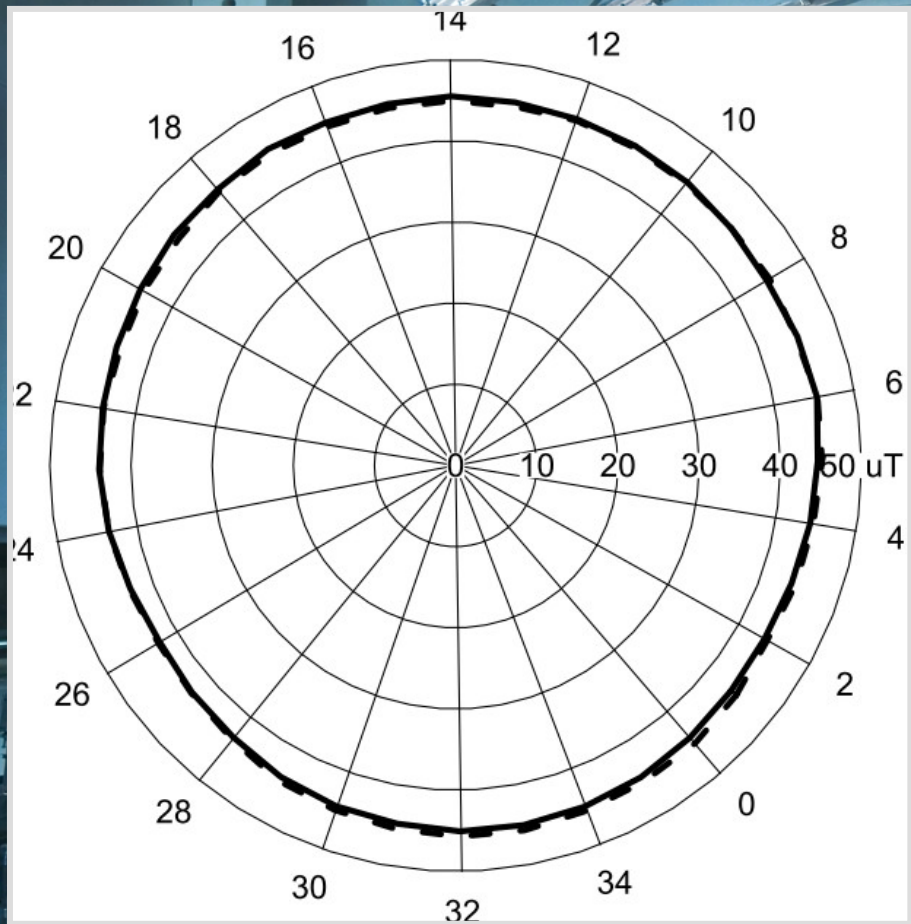
**Trivia question:**  
UHV recipient — LHC vs. KATRIN?







9/2014: successful de-gaussing of experimental hall



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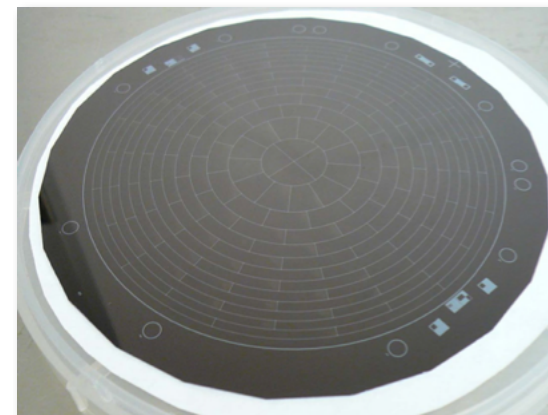
# Detector system

## Requirements:

- detection of  $\beta$ -electrons (mHz to kHz)
- high efficiency ( $> 90\%$ )
- low background ( $< 1$  mHz)
  - passive and active shielding
  - post-acceleration (10-30 kV)
- good energy resolution ( $\sim 1$  keV)

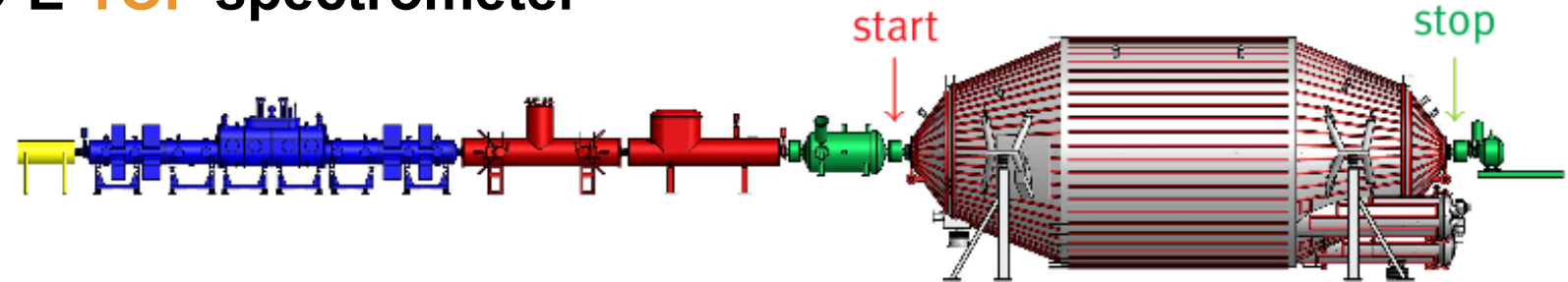
## Characteristics:

- 90 mm  $\varnothing$  Si PIN diode
- thin entry window (50 nm)
- segmented wafer (148 pixels)
  - compensate field inhomogeneities
  - radial-dependent background
  - investigate systematic effects
- detector magnet 3 - 6 T

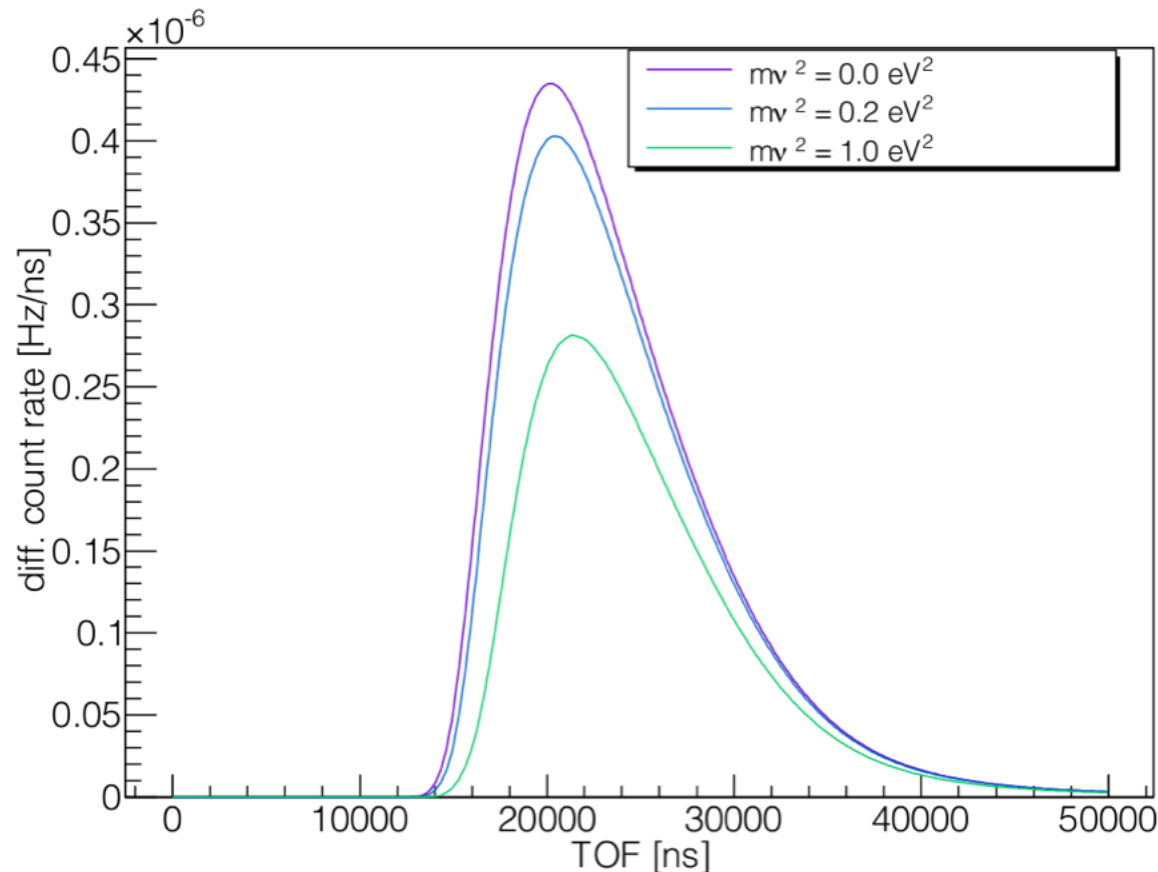


# 1<sup>st</sup> avenue: exploit differential $\beta$ spectrum

## Idea: Upgrade to MAC-E-TOF spectrometer



Comparison of TOF spectra for different neutrino masses for  $E_0 = 18574.0$  eV,  $U_{\text{ret}} = -18570.0$  eV



Spectrometer as 24 m long “delay line”  
→ very sensitive to small differences in surplus energy

TOF spectrum records full  $\beta$  spectrum  
→ save meas. time by using only few voltage settings of MAC-E filter

Coincidence requirement  
→ add. background suppression

### Technical realization?

- (a) pre-spectrometer as gated filter
- (b) radio frequency tagger

# 2<sup>nd</sup> avenue: alternative spectroscopic technique



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### Idea: Cyclotron Radiation Emission Spectroscopy (CRES)

[Formaggio & Monreal, PRD 80 (2009) 051301(R)]

$$\omega(\gamma) = \frac{\omega_c}{\gamma} = \frac{eB}{E_{\text{kin}} + m_e}$$

**Energy** measured via **cyclotron frequency** of single electrons in B field



# 2<sup>nd</sup> avenue: alternative spectroscopic technique



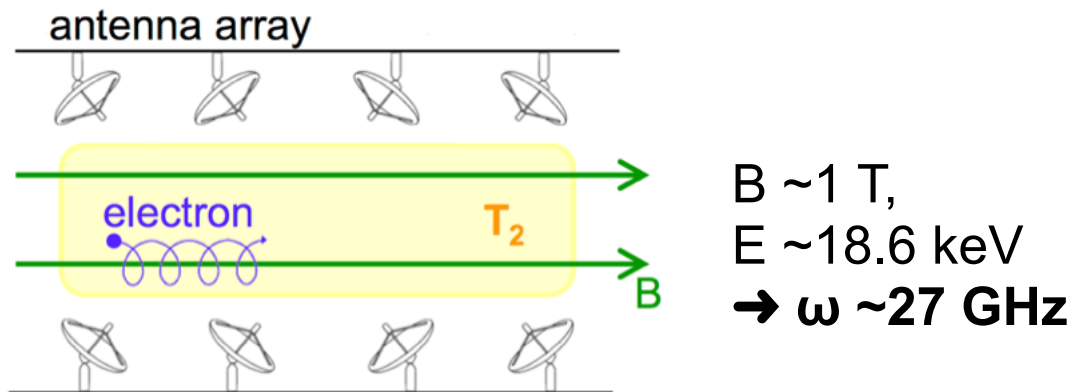
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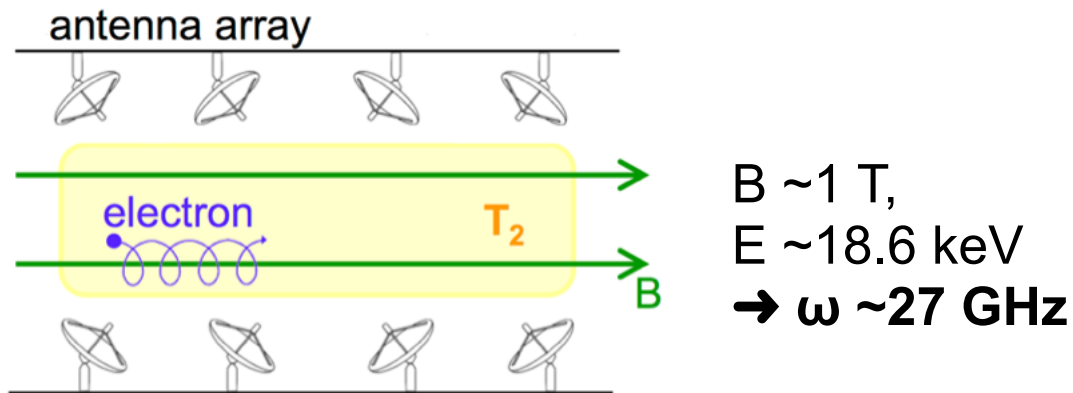
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“KATRIN”-like gaseous source:  
uniform B-field and low-pressure T<sub>2</sub> gas



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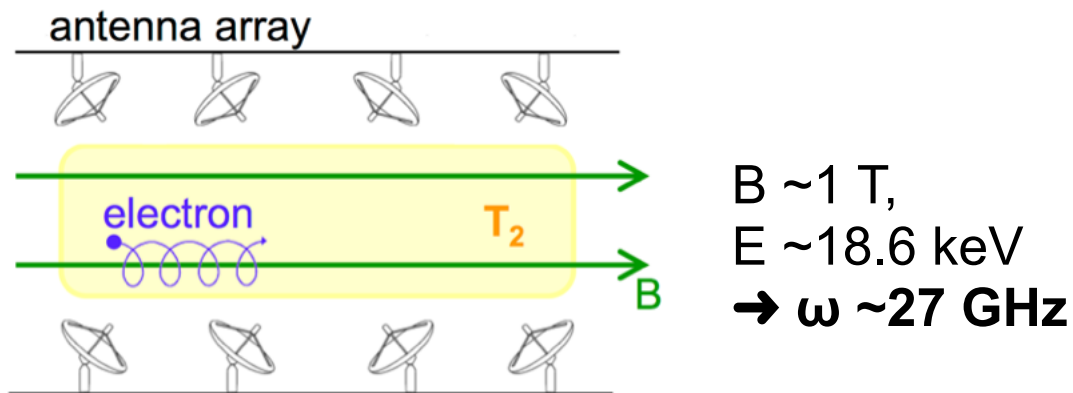
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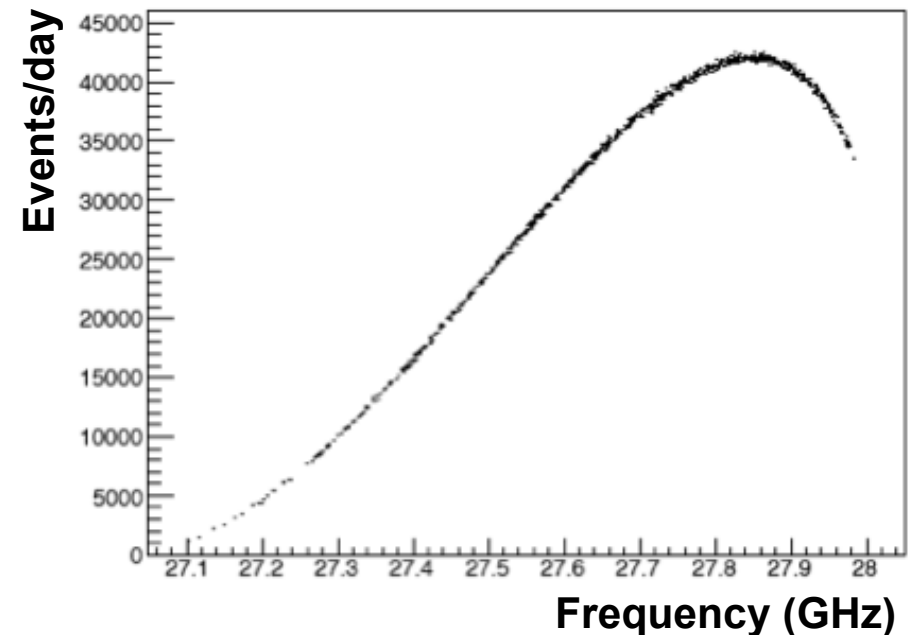
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# Why sterile neutrinos?

Hints of **eV-scale** sterile neutrinos?

Hints of **keV-scale** sterile neutrinos?

# Why sterile neutrinos?

## Hints of **eV-scale** sterile neutrinos?

May explain anomalous oscillation results from

- Short baseline accelerator experiments
- Gallium experiments
- Reactor experiments

## Hints of **keV-scale** sterile neutrinos?

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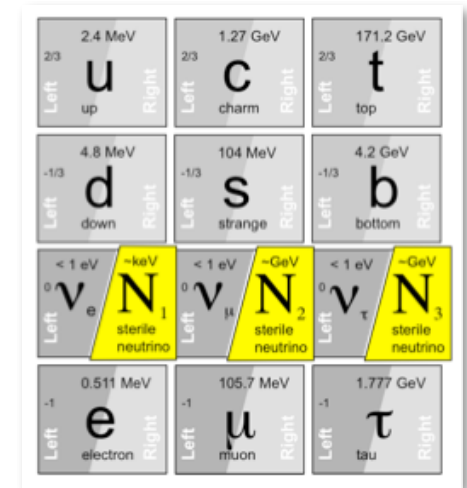
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Well motivated as natural extension of Standard Model ( $\nu$ MSM)



[e.g., Canetti, Drewes, Shaposhnikov (2013)]

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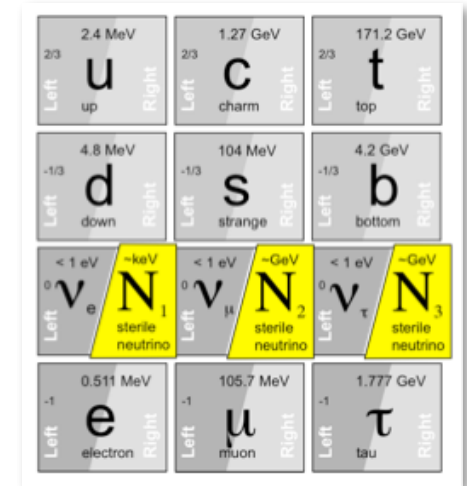
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In agreement with cosmological observations from small to large scales [e.g., Shi & Fuller (1999)]

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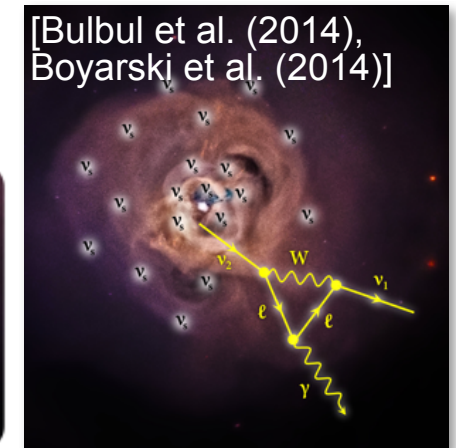
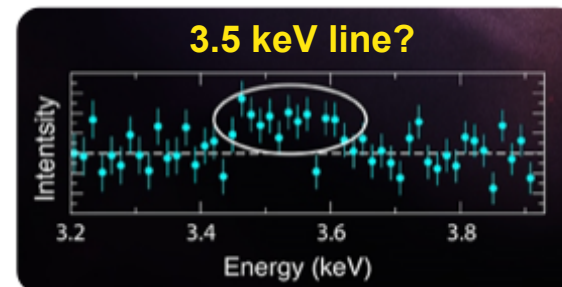
Well motivated as natural extension of Standard Model ( $\nu$ MSM)

2.4 MeV Left $u$ Right up	1.27 GeV Left $c$ Right charm	171.2 GeV Left $t$ Right top
4.8 MeV Left $d$ Right down	104 MeV Left $s$ Right strange	4.2 GeV Left $b$ Right bottom
< 1 eV Left $\nu_e$ Right sterile neutrino $N_1$	< 1 eV Left $\nu_\mu$ Right sterile neutrino $N_2$	< 1 eV Left $\nu_\tau$ Right sterile neutrino $N_3$
0.511 MeV Left $e$ Right electron	105.7 MeV Left $\mu$ Right muon	1.777 GeV Left $\tau$ Right tau

[e.g., Canetti, Drewes, Shaposhnikov (2013)]

In agreement with cosmological observations from small to large scales [e.g., Shi & Fuller (1999)]

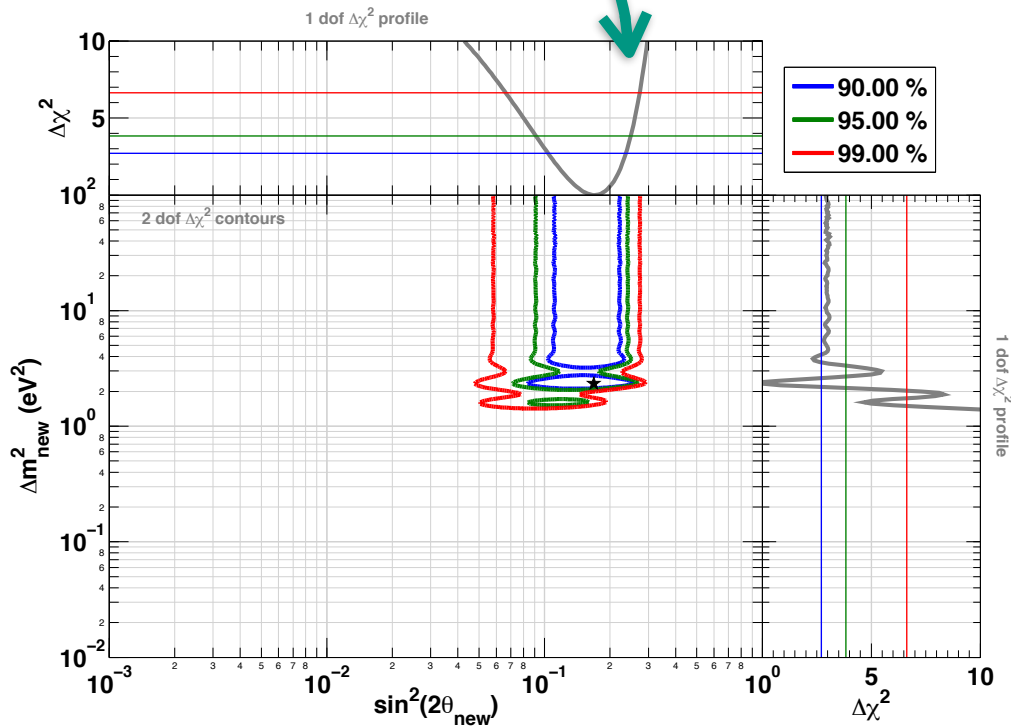
Recent indirect hints from X-ray astronomy?



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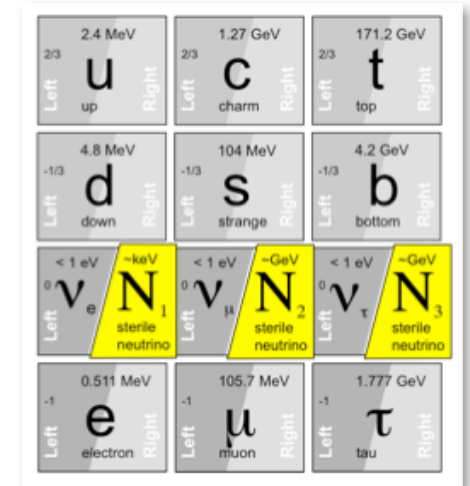
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[G. Mention et al. (2011), updated in White Paper (2014)]

## Hints of keV-scale sterile neutrinos?

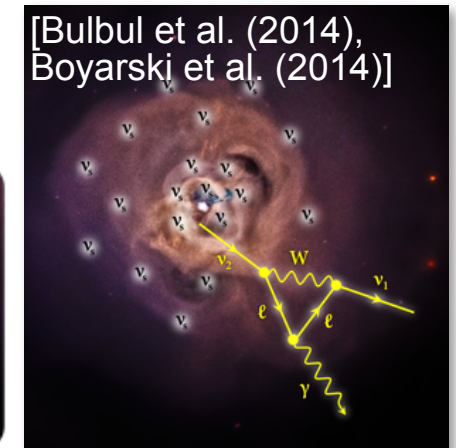
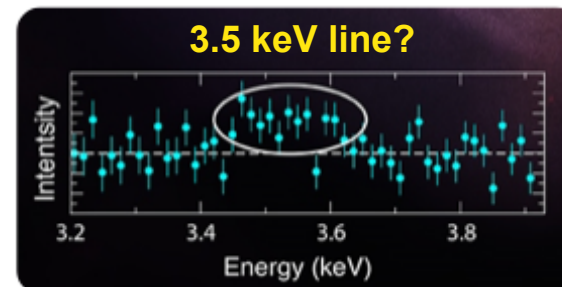
Well motivated as natural extension of Standard Model (νMSM)



[e.g., Canetti, Drewes, Shaposhnikov (2013)]

In agreement with cosmological observations from small to large scales [e.g., Shi & Fuller (1999)]

Recent indirect hints from X-ray astronomy?

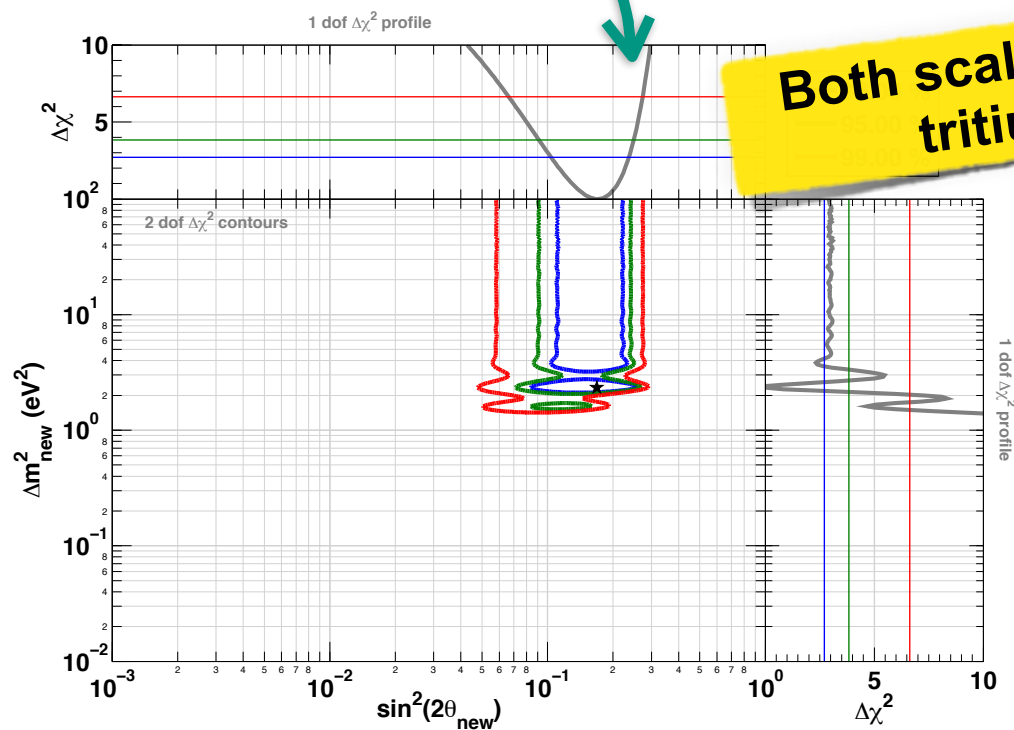


[Bulbul et al. (2014), Boyarski et al. (2014)]

# Why sterile neutrinos?

## Hints of eV-scale sterile neutrinos?

- May explain anomalous oscillation results from
- Short baseline accelerator experiments
- Gallium experiments
- Reactor experiments



[G. Mention et al. (2011), updated in White Paper (2014)]

## Hints of keV-scale sterile neutrinos?

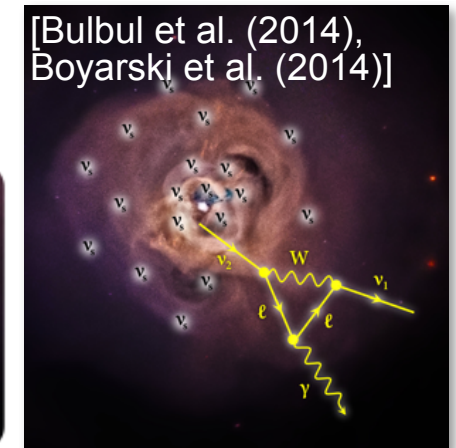
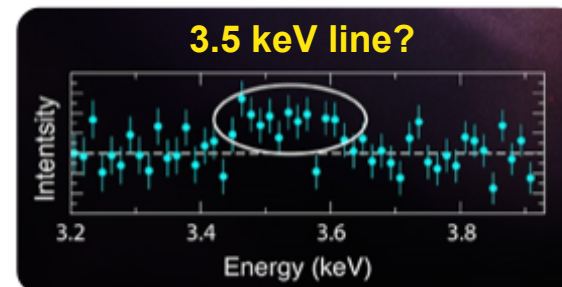
Well motivated as natural extension of Standard Model (νMSM)

2.4 MeV Left up Right	1.27 GeV Left charm Right	171.2 GeV Left top Right
4.8 MeV Left down Right	104 MeV Left strange Right	4.2 GeV Left bottom Right
< 1 eV Left e Right	~keV Left ν <sub>μ</sub> Right	< 1 eV Left ν <sub>τ</sub> Right
0.511 MeV Left electron Right	105.7 MeV Left muon Right	1.777 GeV Left tau Right

[e.g., Canetti, Drewes, Shaposhnikov (2013)]

In agreement with cosmological observations from small to large scales [e.g., Shi & Fuller (1999)]

Recent indirect hints from X-ray astronomy?



[Bulbul et al. (2014), Boyarski et al. (2014)]