Phenomenology of Dark Matter — Precision Calculation and Minimal Models

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Graduiertenkolleg "Strong and Weak Interactions — from Hadrons to Dark Matter" Graduate Days 2015 — 25 November 2015, Münster/Telgte

Curriculum Vitae

- 2000–2005 Studies of Physics Universität Karlsruhe and Université Joseph Fourier Grenoble Double diploma programme of Franco-German University
- 2005–2008 PhD thesis Université Joseph Fourier Grenoble, LPSC Grenoble Supervisor: Prof. Michael Klasen
- 2008–2010 Postdoctoral researcher Universität Würzburg, Institut für Theoretische Physik und Astrophysik Working group of Prof. Werner Porod
- 2010–2011 Young Investigator Group Leader Deutsches Elektronen Synchrotron (DESY) Hamburg Excellence Initiative "Connecting particles with the cosmos"
- since 2011 Maître de Conférences (tenured assistant professor) Université Savoie Mont Blanc, LAPTh Annecy-le-Vieux
- since 2015 Organizer of the "PhysTeV Les Houches" workshop series (with Fawzi Boudjema and Diego Guadagnoli)

Flavour violation in supersymmetric theories

— Constraints on the NMFV MSSM parameter space K. De Causmaecker, B. Fuks, B. Herrmann, N. Mahmoudi, B. O'Leary, W. Porod, S, Sekmen, N. Strobbe — arXiv:1509.05414

- Signatures of non-minimal squark flavour violation at the LHC
 A. Bartl, H. Eberl, B. Herrmann, K. Hidaka, W. Majerotto, W. Porod PLB 698 (2011) arXiv:1007.5483
 A. Bartl, H. Eberl, E. Ginina, B. Herrmann, K. Hidaka, W. Majerotto, W. Porod PRD 84 (2011) arXiv:1107.2775
 M. Bruhnke, B. Herrmann, W. Porod JHEP 1009 (2010) arXiv:1007.2100
- Reconstruction of the squark sector of the MSSM at the LHC
 B. Herrmann, W. Porod in progress

Footprints of SU(5)-like unification in observables at the LHC
 S. Fichet, B. Herrmann, Y. Stoll — PLB 742 (2015) — arXiv:1403.3397
 S. Fichet, B. Herrmann, Y. Stoll — JHEP 1505 (2015) — arXiv:1501.05307
 Y. Stoll, B. Herrmann, S. Fichet — to be published

Two Higgs Doublet Models (e.g. Inert Doublet Model)

— Constraints from dark matter, Higgs boson, vacuum stability A. Goudelis, B. Herrmann, O. Stal — JHEP 1309 (2013) — arXiv:1303.3010

— Constraints from LHC searches
 G. Bélanger, B. Dumont, A. Goudelis, B. Herrmann, S. Kraml, D. Sengupta — PRD 91(2015) — arXiv:1503.07367

Precision calculations for dark matter annihilation...

Minimal models for dark matter (and more)...

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Precision calculations for dark matter annihilation...

Minimal models for dark matter (and more)...

next 20-25 minutes...



Predicting the dark matter relic density more precisely

http://dmnlo.hepforge.org





The Minimal Supersymmetric Standard Model (MSSM)

SM Particles		Spin		Spin	Superpartners	
Quarks	$\begin{pmatrix} u_L & d_L \end{pmatrix}$	1/2	Q	0	$egin{pmatrix} ilde{u}_L & ilde{d}_L \end{pmatrix}$	Squarks
	u_R^\dagger	1/2	\bar{u}	0	$ ilde{u}_R^*$	
	d_R^\dagger	1/2	$ \overline{d}$	0	$ ilde{d}_R^*$	
Leptons	$\begin{pmatrix} \nu & e_L \end{pmatrix}$	1/2	L	0	$egin{pmatrix} ilde{ u} & ilde{e}_L \end{pmatrix}$	Sleptons
	e_R^\dagger	1/2	\bar{e}	0	$ ilde{e}_R^*$	
Higgs	$\begin{pmatrix} H_u^+ & H_u^0 \end{pmatrix}$	0	H_u	1/2	$\begin{pmatrix} \tilde{H}_u^+ & \tilde{H}_u^0 \end{pmatrix}$	Higgsinos
	$\begin{pmatrix} H_d^0 & H_d^- \end{pmatrix}$	0	H_d	1/2	$\begin{pmatrix} \tilde{H}_d^0 & \tilde{H}_d^- \end{pmatrix}$	
W bosons	W^0, W^{\pm}	1		1/2	$ ilde W^0, ilde W^\pm$	Winos
B boson	B^0	1		1/2	$ ilde{B}^0$	Bino
Gluon	g	1		1/2	\widetilde{g}	Gluino
Graviton	G	2		3/2	$ ilde{G}$	Gravitino

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SM Particles		Spin		Spin	Superpartners	
Quarks	$\begin{pmatrix} u_L & d_L \end{pmatrix}$	1/2	Q	0	$\left(\widetilde{u}_L \widetilde{d}_L \right)$	Squarks
	u_R^\dagger	1/2	$ar{u}$	0	$ ilde{u}_R^*$	
	d_R^\dagger	1/2	\overline{d}	0	$ ilde{d}_R^*$	
Leptons	$egin{pmatrix} u & e_L \end{pmatrix}$	1/2	L	0	$\begin{pmatrix} ilde{ u} & ilde{e}_L \end{pmatrix}$	Sleptons
	e_R^\dagger	1/2	\bar{e}	0	${ ilde e}_R^*$	
Higgs	$\begin{pmatrix} H_u^+ & H_u^0 \end{pmatrix}$	0	H_u			
	$\begin{pmatrix} H_d^0 & H_d^- \end{pmatrix}$	0	H_d	1/2	$ ilde{\chi}^0_{1,2,3,4}$	Neutralinos
W bosons	W^0, W^{\pm}	1		1/2	$\tilde{\chi}_{1.2}^{\pm}$	Charginos
B boson	B^0	1			,	
Gluon	g	1		1/2	\tilde{g}	Gluino
Graviton	C	?		3/9	\tilde{C}	Cravitino

Lightest neutralino is dark matter (WIMP) candidate "par excellence"

 $\tilde{\chi}_{1}^{0} = Z_{1\tilde{B}}\tilde{B} + Z_{1\tilde{W}}\tilde{W} + Z_{1\tilde{H}_{1}}\tilde{H}_{1} + Z_{1\tilde{H}_{2}}\tilde{H}_{2}$

Dark matter relic abundance — freeze-out picture



Time evolution of number density of the relic particle described by Boltzmann equation

$$\frac{\mathrm{d}n}{\mathrm{d}t} = -3Hn - \langle \sigma_{\mathrm{ann}}v \rangle \left(n^2 - n_{\mathrm{eq}}^2\right)$$

Prediction of dark matter relic density (if masses and interactions are known)

$$\Omega_{\chi} h^2 = \frac{m_{\chi} n_{\chi}}{\rho_{\rm crit}} \sim \frac{1}{\langle \sigma_{\rm ann} v \rangle}$$

Dark matter relic abundance — freeze-out picture



Dark matter relic abundance very precisely known Planck collaboration 2015

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$$\Omega_{\chi}h^{2} = \frac{m_{\chi}n_{\chi}}{\rho_{\rm crit}} \sim \frac{1}{\langle \sigma_{\rm ann}v \rangle}$$
(dis)favoured parameter regions...?
$$\Omega_{\rm CDM}h^{2} = 0.1199 \pm 0.0022$$

Dark matter relic abundance — freeze-out picture



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Prediction of dark matter relic density (if masses and interactions are known)



Computational tools allow an efficient calculation of the (neutralino) relic density: DarkSUSY Bergström, Edsjö, Gondolo et al. 2004-2015, micrOMEGAs Bélanger, Boudjema, Pukhov et al. 2003-2015, ...

A closer look on the (co)annihilation cross-section

Cross-section in Boltzmann equation includes the sum over all relevant processes

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Cross-section in Boltzmann equation includes the sum over all relevant processes

$$\frac{\mathrm{d}n}{\mathrm{d}t} = -3Hn - \langle \sigma_{\mathrm{ann}}v \rangle \left(n^2 - n_{\mathrm{eq}}^2\right)$$

$$\langle \sigma_{\mathrm{ann}}v \rangle = \sum_{i,j} \sigma_{ij} v_{ij} \frac{n_i^{\mathrm{eq}}}{n_\chi^{\mathrm{eq}}} \frac{n_j^{\mathrm{eq}}}{n_\chi^{\mathrm{eq}}}$$



A closer look on the (co)annihilation cross-section

Cross-section in Boltzmann equation includes the sum over all relevant processes



Only co-annihilations with almost mass-degenerate particles are numerical relevant Typical examples in MSSM: other neutralinos, charginos, stau, stop

Motivation for higher order corrections

All processes implemented in public codes — but only at the (effective) tree-level



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Higher-order loop corrections can give important contributions to cross-sections In particular, sizeable impact from QCD corrections due to strong coupling constant More precise theoretical predictions needed to keep up with experimental improvements



Motivation for higher order corrections

All processes implemented in public codes — but only at the (effective) tree-level



Higher-order loop corrections can give important contributions to cross-sections In particular, sizeable impact from QCD corrections due to strong coupling constant More precise theoretical predictions needed to keep up with experimental improvements

















Similar setup for use with DarkSUSY in development.

J. Edsjö, B. Herrmann, C. Niblaeus — in progress...





Loop corrections





Interlude — a few technical details

Loop diagrams include UV-divergent integrals → Renormalization!

Hybrid on-shell/ \overline{DR} renormalization scheme for the squark sector (3rd generation), which is applicable to all (co)annihilation processes



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Hybrid on-shell/ \overline{DR} renormalization scheme for the squark sector (3rd generation), which is applicable to all (co)annihilation processes



Loop diagrams contain **IR-divergencies** (soft and/or collinear), which vanish when taking into account the real emission of a gluon $(2 \rightarrow 3 \text{ processes})$

Dipole Subtraction Method and Phase Space Slicing Catani, Seymour (2001)

$$\sigma_{\rm NLO} = \int_{3} \left[\mathrm{d}\sigma^{\rm R} \Big|_{\epsilon=0} - \mathrm{d}\sigma^{\rm A} \Big|_{\epsilon=0} \right] + \int_{2} \left[\mathrm{d}\sigma^{\rm V} + \int_{1} \mathrm{d}\sigma^{\rm A} \right]_{\epsilon=0}$$



Neutralino pair annihilation into top quarks



Herrmann, Klasen, Kovařík — Phys. Rev. D 79: 061709 (2009) — arXiv:0901.0481 [hep-ph] Harz, Herrmann, Klasen, Kovařík, Steppeler — Phys. Rev. D 91: 034028 (2015) — arXiv:1409.2898 [hep-ph]

Neutralino pair annihilation into top quarks



Annihilation cross-section enhanced by up to 50% by radiative corrections Corrections can lead to **important shifts for preferred regions** (e.g. ~200 GeV for m_{stop})

Effective Yukawa couplings (as e.g. in micrOMEGAs) very good approximation around Higgsresonances, but other sub-channels can be dominant (here: Z/squark-exchange)

Generalization of previous results to annihilation of any gauginos into any quarks



Harz, Herrmann, Klasen, Kovařík, Steppeler — Phys. Rev. D 91: 034028 (2015) — arXiv:1409.2898 [hep-ph]

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Neutralino pair annihilation into gluons

Process occurring only at one-loop level (no renormalization needed)



Fedkevych, Klasen, Kovařík — in progress...







The observed Higgs mass favours large mass splitting of the stops, typically achieved through large trilinear coupling A_t

$$m_{h^0} = m_Z^2 \cos^2 2\beta + \frac{3g^2 m_t^4}{8\pi^2 m_W^2} \left[\log \frac{M^2}{m_t^2} + \frac{X_t^2}{M^2} \left(1 - \frac{X_t^2}{12M^2} \right) \right]$$
$$X_t = A_t - \mu \tan \beta \approx \sqrt{6}M = \sqrt{6}\sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}}$$

Harz, Herrmann, Klasen, Kovařík, Le Boulc'h — Phys. Rev. D 87: 054031 (2013) — arXiv:1212.5241 [hep-ph]



Neutralino-stop co-annihilation



Relative corrections of 40-50% observed for the coannihilation cross-section, leading to an **important shift** (up to almost 25% — more than Planck uncertainty!) for the predicted **neutralino relic density**

Coannihilation into SM-like Higgs and gluon most important (other final states generally subdominant)

Stop pair ar \tilde{h}_{i}^{V} $\tilde{h}_{$

Exchange of multiple gluons in the initial state (in addition to one-loop diagrams) — resummation to all orders using non-relativistic QCD



$$\sigma^{\text{Coul}} = \frac{4\pi}{\text{vm}_{\tilde{t}}^2} \Im\left\{ \mathsf{G}^{[1]} \left(\mathbf{r} = \mathbf{0}; \sqrt{\mathsf{s}} + \mathsf{i} \Gamma_{\tilde{t}} \right) \right\} \sigma^{\text{LO}}$$
$$\left[\mathsf{H}^{[1]} - \left(\sqrt{\mathsf{s}} + \mathsf{i} \Gamma_{\tilde{t}} \right) \right] \mathsf{G}^{[1]} = \delta^{(3)}(\mathbf{r})$$



Coulomb corrections dominant for small values of p_{cm} (Coulomb singularity), while fixed-order corrections dominant for high-momentum region

Stop pair annihilation into electroweak final states included — coloured final states to be implemented...

Harz, Herrmann, Klasen, Kovařík, Meinecke — Phys. Rev. D 91: 034012 (2015) — arXiv:1410.8063 [hep-ph] Herrmann, Klasen, Kovařík, Schmiemann — *in progress*...

Scale dependence of neutralino (co)annihilation

Loop calculation introduces a dependence on an unphysical parameter: renormalization scale — Evaluation of theory uncertainty by varying renormalization scale

scale-dependent parameters

Scale dependence of neutralino (co)annihilation

Loop calculation introduces a dependence on an unphysical parameter: renormalization scale — Evaluation of theory uncertainty by varying renormalization scale

 $\mu_{\rm R} = 500 \dots 2000 \,\,{
m GeV}$ $A_b, \theta_{\tilde{t}}, \theta_{\tilde{t}}, \alpha_s, m_b$ A_t , scale-dependent parameters 1.2 $\rightarrow th^0$, W α_S $\tilde{\boldsymbol{\chi}}_{1}^{0}\tilde{t}_{1}$ Scale uncertainty reduced at the 1.0 $\sigma^{ tree}$ ($\sigma^{ tree}/\sigma^{ tMO}$) one-loop level w.r.t. to tree-level GeV^{-2}) $(\sigma^{\text{NLO}}/\sigma^{\text{MO}})$ 0.8 result (as expected) $\sigma^{ t NLO}$ $(\sigma^{\text{NLO}}/\sigma^{\text{tree}})$ σν (10⁻⁸ (— main effect from mixing angle 0.6 and trilinear coupling 0.4 dependence of α_s subdominant 0.2 PRELIMINARY.6 1.0 100 300 0 200 400 500 p_{cm} (GeV)

Harz, Herrmann, Klasen, Kovařík, Steppeler — to be published...

Direct dark matter detection

Same topologies as neutralino pair annihilation into quarks



Direct dark matter detection

Same topologies as neutralino pair annihilation into quarks



Direct dark matter detection



Same topologies as neutralino pair annihilation into quarks

Calculation carried out at very low energy: $p_{cm} \sim 0$

- need to implement specific reduction procedure for threshold...

Finiteness (UV and IR) of loop calculation within effective theory verified



Recent experimental improvements (WMAP, Planck...) require more precise predictions of the dark matter relic density on the theory side

Package providing a calculation of neutralino (co)annihilation including QCD corrections

$$\begin{split} \tilde{\chi}\tilde{\chi}' &\to q\bar{q}' \\ \tilde{\chi}\tilde{q} &\to q'H/q'V \\ \tilde{q}\tilde{q}^* &\to HH/HV/VV \end{split}$$

numerically implemented results published

$$\begin{split} \tilde{\chi}\tilde{\chi}' &\to gg/\gamma\gamma \\ \tilde{q}\tilde{q}^* &\to q\bar{q}' \\ \tilde{q}\tilde{q} &\to qq \\ \tilde{\tau}\tilde{\tau}^* &\to qq' \end{split}$$

work in progress...

Impact of corrections on the relic density more important than current exp. uncertainty

Minimal models — Dark matter and Neutrino masses









Parameters of the model: (

Analysis of the relevant parameter space with respect to dark matter and flavour constraints — Markov Chain Monte Carlo technique for efficient scanning — study also dark matter annihilation channels, neutrino and collider aspects

Esch, Herrmann — in preparation...

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Loop corrections to DM annihilation...? Study of collider signatures...? — link to **DM@NL** – Large Hadron Collider / linear collider sonja Esch-visit Annecy march 2016 Detailed study of neutrino sector...? - mass patterns, astroparticle physics Graduiertenkolleg 2149 Graduiertenkolleg Group Research Training

